Isolation by Distance in Populations with Long-Range Dispersal

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ABSTRACT Limited dispersal results in isolation by distance in spatially structured populations, in which individuals found further apart tend to be less related to each other. Models of populations undergoing short-range dispersal predict a close relation between the distance individuals disperse and the length scale over which two sampled individuals are likely to be closely related. In this work, we study the effect of long jumps on patterns of isolation by distance by replacing the typical short-range dispersal kernel with a long-range, power-law kernel. We find that incorporating long jumps leads to a slower decay of relatedness with distance, and that the quantitative form of this slow decay contains visible signatures of the underlying dispersal process.

KEYWORDS Isolation By Distance; Identity By Descent; Dispersal Rate; Long-Range Dispersal

Introduction

Direct measurement of dispersal in natural populations is often difficult or impossible due to practical difficulties in tracking large numbers of individuals over long periods of time. It is often more feasible to instead infer dispersal from spatial patterns of genetic diversity (Cayuela *et al.* 2018; Bradburd and Ralph 2019; Battey *et al.* 2020). Populations with limited dispersal should exhibit "isolation by distance": the more distant individuals are from each other in space, the less related they tend to be (Wright 1946; Rohlf and Schnell 1971; Slatkin 1991). While in general the spatial pattern of genetic diversity depends on selection (Barton *et al.* 2013; Allman and Weissman 2018), for populations evolving neutrally the strength of isolation by distance is simply determined by the balance between dispersal and mutation, and thus if the mutation rate is known, the dispersal rate can be inferred directly (Malécot 1975; Slatkin and Arter 1991; Slatkin 1993).

For populations spread over a fairly continuous range, rather than being clumped into a small number of discrete subpopulations, dispersal is often assumed to be short-range, with displacement approximately following a normal distribution (Barton *et al.* 2002; Ringbauer *et al.* 2017). If dispersal is unbiased and homogeneous, it is then characterized by a single parameter, the dispersal rate D (the diffusion constant). Pairwise genetic similarity is predicted to decay exponentially with distance, with a decay rate of $\sqrt{\mu/D}$, where μ is the mutation rate (Kimura and Weiss 1964; Slatkin and Barton 1989; Rousset 1997, 2000; Barton *et al.* 2002; Ralph and Coop 2013).

However, many populations exhibit at least occasional long-range dispersal (Clark 1998; Willson 1993; Atkinson *et al.* 2002; Baguette 2003; Brockmann *et al.* 2006; Dai *et al.* 2007; Fric and Konvicka 2007; Devaux *et al.* 2007; Aguillon *et al.* 2017; Vallaeys *et al.* 2017). For many other populations, particularly non-animal ones, very little is known about dispersal, and we would like to know how to use genetic data to determine if it is short- or long-range (Nathan *et al.* 2003; Cayuela *et al.* 2018). Even for populations where dispersal is primarily short-range, even very rare long jumps that are unlikely to be observed directly (Koenig *et al.* 1996) can have large effects on the population's evolution (Ibrahim *et al.* 1996; Mancinelli *et al.* 2003; Bialozyt *et al.* 2006; Brockmann and Hufnagel 2007; Wingen *et al.* 2007; Fayard *et al.* 2009; Hallatschek and Fisher 2014; Paulose *et al.* 2019; Paulose and Hallatschek 2020), so we would also like to be able to infer their pattern and frequency.

In this work, we further explore the effects of long-range (power-law) dispersal in neutrally evolving, demographically stable populations with constant density. This problem was previously studied by Nagylaki (1976) in one dimension for moderately long-range dispersal with finite mean distance and by Chave and Leigh Jr (2002) in two dimensions for Cauchy-distributed dispersal. Recently, Janakiraman (2017) studied an analogous problem in chemical physics in the same regime as Nagylaki (1976) and found complementary results. We unify and extend this work to cover arbitrary power law dispersal tails in both one and two spatial dimensions, and find expressions for how the pattern of isolation by distance reflects the underlying dispersal process at both large and small distances. We also find

Figure 1 **α controls the size and number of long-range jumps.** Two-dimensional Lévy flights with α values 0.95, 1.5, and 2. When α = 2, Lévy flight motion reduces to normal diffusion without any long-range jumps. Lévy flights with α < 2 have divergent mean squared displacement and a power law dispersal kernel proportional to $x^{-\alpha-1}$. For α < 1, the mean displacement also diverges, and large jumps become noticeably more prevalent than for 1 < α < 2. Circles mark the beginnings of the trajectories, while squares mark the ends. The generalized dispersal constants D_α are chosen such that the flights all have the same characteristic displacement at the time step marked by triangles. On shorter time scales, the diffusive trajectory tends to have the largest displacement, while on longer time scales the trajectory with α < 1 tends to have the largest displacement.

how the distribution of time to the most recent common ancestor
of a pair of individuals depends on the distance between them.
Forien (2019) recently proposed a similar model for populations
evolving according to a spatial Λ-Fleming-Viot process with
long-range dispersal; our analysis may also apply to this model.

Model

We consider two lineages sampled in the present a distance x
apart, and trace their lineages backwards through time until
they coalesce. We assume that individuals move through an
infinite range according to a distribution that falls off as a power
law at long dispersal distances. For concreteness, we will mostly
assume that the range is one-dimensional and that the lineages
follow Lévy flights, a flexible, mathematically tractable way to
model dispersal with power-law tails (Jespersen *et al.* 1999; Metzler and Klafter 2000; Metzler *et al.* 2009). Lévy flights occur
naturally as the limit of any trajectory composed of many independent, identically distributed jumps, including the special
case of diffusive motion in which the jumps have a finite variance. In this finite-variance case, we also consider power-law
dispersal kernels that cannot be described by a Lévy flight.
For a lineage following a Lévy flight, its displacement y from
its present position t generations in the past follows a Lévy

Symbol	Definition
x	Distance between samples
ρ	Population density
α	Stability parameter of dispersal tail
D_α	Generalized dispersal constant
μ	Mutation rate
ψ	Probability of identity
$\bar{x} = (D_\alpha/\mu)^{1/\alpha}$	Characteristic length scale of identity

alpha-stable distribution:

$$K_1(y|t) = \frac{1}{2\pi} \int_{-\infty}^{\infty} dk \exp\left(-iky - D_\alpha t|k|^\alpha\right). \quad (1)$$

The parameter D_α is a generalized diffusion constant with units of length$^\alpha$/time. It sets the scale of dispersal: at time t, a typical lineage will have a displacement $\bar{y}(t) \sim (D_\alpha t)^{1/\alpha}$. As shown in Fig. 1, the "stability parameter" α, $0 < \alpha \leq 2$, controls the tail of dispersal: for $\alpha < 2$, the probability that at time t a lineage has moved an abnormally long distance $y \gg \bar{y}(t)$ is $K_1(y|t) \propto y^{-1-\alpha}$, a power law tail. In the limiting case $\alpha = 2$, dispersal reduces to ordinary short-range diffusion and K_1 is just a normal distribution with standard deviation $\sqrt{Dt}$. In two dimensions, the definition (1) must be adjusted (see (43) in the Methods), but the only change to the approximate description above is that the tail is $\propto y^{-2-\alpha}$.

When the two lineages encounter each other, they coalesce at rate $1/\rho$, where ρ is the density of the population. Technically, for $\alpha \leq 1$ in one dimension the two lineages will never be at exactly the same position. In two dimensions, all values of α have this problem, including ordinary diffusion (Mörters and Peres 2010). So really there must be some small distance δ within which lineages coalesce at rate $\sim 1/(\delta^d\rho)$, where d is the dimensionality of the range. At these small scales, even the model of independent diffusion of lineages breaks down (Barton *et al.* 2010). But we will see below that this coalescence length scale does not affect isolation by distance on larger scales $x \gg \delta$.

We are interested in the probability ψ of identity by descent of our sample pair as a function of the distance between them, x, which we will also refer to as the homozygosity or relatedness. If the time to their most recent common ancestor is T and the mutation rate is μ, then ψ is given by:

$$\psi(x) = E\left[e^{-2\mu T}|x\right]. \quad (2)$$

Although usually it is ψ rather than the coalescence time T itself that is directly observable, T is important for, e.g., determining whether it is reasonable to assume stable demography, so we will also find expressions for its distribution $p(t|x)$.

Results

In this section, we will describe our main results and provide brief sketches of the logic behind key features. Roughly speaking, the basic intuition is that the sampled pair will be identical if their lineages coalesce within the past $\sim 1/\mu$ generations. In this time, they will disperse a typical distance $\sim \bar{x} \equiv (D_\alpha/\mu)^{1/\alpha}$, so this is the key length scale over which identity decays: pairs separated by $x \ll \bar{x}$ should be relatively closely related, while identity between pairs separated by $x \gg \bar{x}$ should be rare.

Figure 2 The form of isolation by distance in one dimension is universal at long distances and varies at short distance. Approximate form for the probability of identity as a function of distance, $\psi(x)$, for different dispersal kernels α. Different regimes of the parameter space are separated by solid lines, and labelled by their qualitative dynamics. Coalescence for distant pairs, $x \gg \overline{x}$, where $\overline{x} = (D_\alpha/\mu)^{1/\alpha}$ is the characteristic length scale of identity, occurs via one long jump for all α, leading to the power law scaling at large distances predicted by (4). Coalescence for nearby pairs, $x \ll \overline{x}$, depends on the value of α considered. For $\alpha > 2$, the motion of lineages across short distances is diffusive and ψ scales exponentially, as shown in (3). For $1 < \alpha < 2$, short distances are covered via many small jumps, but lineages spread faster than they would under diffusion, leading to the broader scaling found in (5). For $\alpha < 1$, even short distances are covered by one quick jump, leading to the power law shown in (6). Lineages that do not coalesce quickly (at $t \ll 1/\mu$) will likely never coalesce, and probability of identity is limited by δ, rather than μ, as shown in (7). We use "$\sim$" to denote proportionality in the limit of large population density where $\psi(0) \ll 1$.

For the classic case of diffusive motion ($\alpha = 2$), this length scale is $\overline{x} = \sqrt{D/\mu}$, and the probability of identity falls off exponentially (Barton *et al.* 2002):

$$\psi(x) = \frac{e^{-x/\overline{x}}}{4\rho\overline{x}\mu + 1}. \tag{3}$$

Here we generalize (3) to $\alpha \neq 2$, and find simple approximate expressions for ψ in different parameter regimes, illustrated in Fig. 2, including a universal form for all dispersal kernels at long distances. Intuitively, long-range dispersal broadens the distribution of coalescence times for pairs at a given separation x, creating more overlap in the distributions for different x values (Fig. 3).

Distant pairs

For distant samples, $x \gg \overline{x}$, we expect substantial isolation by distance. For the pair to coalesce, their lineages must approach within δ of each other. The most likely way for this to happen within time $\lesssim 1/\mu$ if for one lineage to cover the distance in a long jump. Since such jumps occur at rate $\sim D_\alpha x^{-\alpha-1}\delta$, this occurs with probability $\sim D_\alpha x^{-\alpha-1}\delta/\mu$. The lineages must then coalesce within their neighborhood of $\sim \delta\rho$ individuals before they mutate, which occurs with probability $\sim 1/(\mu\rho\delta)$. We therefore expect that the probability of identity is $\psi \sim D_\alpha x^{-\alpha-1}\delta/\mu/(\mu\rho\delta) = D_\alpha x^{-\alpha-1}/(\mu^2\rho)$, i.e., that there is a power-law dependence of identity on distance, with the same exponent as that of dispersal. A more careful calculation (see Methods) recovers the result Eq. (37) of Nagylaki (1976):

$$\psi(x \gg \overline{x}) \approx \frac{\Gamma(1+\alpha)\sin(\pi\alpha/2)}{2\pi}\frac{D_\alpha}{\mu^2\rho x^{1+\alpha}}, \tag{4}$$

up to an additional prefactor that enters when $\psi(0)$ is close to one. We confirm (4) with simulations and numerical analysis (Fig. 4).

Nearby pairs: moderately long-range dispersal

For moderately long-rage dispersal, $1 < \alpha \leq 2$, nearby pairs should be nearly as related as pairs from the same deme: $\psi(x \ll \overline{x}) \approx \psi(0)$. In the time $\sim 1/\mu$ before they mutate, the lineages will disperse over a distance $\sim \overline{x}$. There are $\sim \rho\overline{x}$ individuals in this neighborhood, so the overall average rate of coalescence is $\sim 1/(\rho\overline{x})$. If this is large compared to the mutation rate, then homozygosity will be high, $\psi(0) \approx 1$, while if coalescence is slow compared to mutation, homozygosity will be low, $\psi(0) \sim 1/(\mu\rho\overline{x})$. A more detailed analysis (see Methods) recovers the result of Nagylaki (1976) (Eq. (39)) for $\psi(0)$, and the leading distance dependence was found by Janakiraman (2017) (Eq. (C1)):

$$\psi(x \ll \overline{x}) \approx \frac{1 - \frac{\alpha\sin(\pi/\alpha)}{\Gamma(\alpha)\cos(\pi(1-\alpha/2))}\left(\frac{x}{\overline{x}}\right)^{\alpha-1}}{2\alpha\sin(\pi/\alpha)\rho\overline{x}\mu + 1}. \tag{5}$$

We confirm (5) with simulations and numerical analysis (Fig. 4). The scaling in (5) reflects the superdiffusive dispersal of the lineages.

Nearby pairs: very long-range dispersal

For very long-range dispersal in which the mean jump size diverges, $\alpha < 1$, nearby lineages are likely to either coalesce very quickly or to disperse across the whole range before coalescing (Palyulin *et al.* 2014). This "now-or-never" dynamic has the interesting effect of making the local probability of identity by descent independent of the mutation rate (Fig. 5), since the main competition is between coalescence and long-range dispersal rather than between coalescence and mutation. Intuitively, the pair of lineages take time $t_0 \sim x^\alpha/D_\alpha$ to disperse across the distance between them. From that time on, they are roughly evenly spread over a range $\sim (D_\alpha t)^{1/\alpha}$, and so coalesce at rate $\sim (D_\alpha t)^{-1/\alpha}/\rho$. Integrating this rate over time starting from t_0 out to $\sim 1/\mu$, we find that $\psi \sim 1/(\rho D_\alpha x^{1-\alpha})$, with the upper limit of integration only negligibly decreasing ψ. So ψ again follows a power law, although a different one from the long-distance $1/x^{1+\alpha}$. We calculate ψ more carefully in the Methods to find:

$$\psi(\delta \ll x \ll \overline{x}) \approx \frac{\Gamma(1-\alpha)\sin(\pi\alpha/2)}{2\pi}\frac{1}{\rho D_\alpha x^{1-\alpha}}, \tag{6}$$

again up to an additional prefactor that enters when $\psi(0)$ is close to one. We confirm (6) with simulations and numerical analysis (Fig. 4 and Fig. 5).

The power law in (6) makes it diverge at very short distances, where it breaks down. Instead, for individuals within the same deme, $x < \delta$, ψ flattens out. Roughly speaking, individuals

Figure 3 Long-range jumps affect when and where lineages coalesce. Qualitative illustrations of lineage dynamics for each of the three α regimes in 1 dimension. Typical histories are shown for nearby samples ($x \ll \overline{x}$, blue) and distant samples ($x \gg \overline{x}$, red). **Left**: For diffusive motion, $\alpha = 2$, initial separation x is a relatively good predictor of coalescence time. **Center**: For moderately long-range dispersal, $1 < \alpha < 2$, large jumps broaden the spatial and temporal ranges over which lineages coalesce. Lineages at large separations $x \gg \overline{x}$ are occasionally able to coalesce at times comparable to $1/\mu$. **Right:** For very long-range dispersal, $\alpha < 1$, large jumps are common. This allows for rapid coalescence of lineages at both small and large distances, but also lets lineages jump very far away from each other and avoid coalescing.

1 coalescence at rate $1/(\rho\delta)$ and disperse outside of coalescence
2 range at rate $\sim D_\alpha \delta^{-\alpha}$. When coalescence is faster, probability
3 of identity is high, $\psi(0) \approx 1$, while when dispersal is faster it
4 is low, $\psi \sim 1/(\rho\delta)/(D_\alpha\delta^{-\alpha}) = 1/(\rho D_\alpha \delta^{1-\alpha})$. A more careful
5 calculation gives (see Methods):

$$6 \quad \psi(x \ll \delta) \approx 1 \left/ \left[1 + \frac{2^{(\alpha+3)/2}\pi}{\Gamma(1/2 - \alpha/2)} \rho D_\alpha \delta^{1-\alpha} \right], \right. \quad (7)$$

although these numerical factors depend on the details of the
model and are only valid when $\psi(0)$ is small. We confirm (7)
with simulations (Fig. 5).

Finite Variance Dispersal

In addition to considering Lévy flight dispersal kernels with
$0 < \alpha \leq 2$, we consider F-distribution kernels (see Methods)
with steeper power law tails $\alpha > 2$. These have finite variance
and approach a diffusion after infinitely many steps, but at any
finite time will be different, particularly in the tail. As shown in
Fig. 6, for nearby pairs, $x \ll \overline{x}$, relatedness decays exponentially,
as in the case of purely diffusive motion. However, for distant
pairs, $x \gg \overline{x}$, coalescence tends to occur via one long jump,
and the probability of identity follows the power law jump tail
$\propto x^{-\alpha-1}$ (Fig. 6), i.e., the same behavior as the infinite-variance
Lévy flights described above.

Two Dimensions

While we have focused above on the case of one spatial dimen-
sion, the same intuition applies in two dimensions. In some
ways the results are even simpler—the same expressions apply
for both $\alpha < 1$ and $1 < \alpha < 2$ (Fig. 7). For distant individu-
als, $x \gg \overline{x}$, the dynamics of coalescence are again dominated by
large jumps in a single generation, and the probability of identity

by descent again has the same power law tail as the dispersal
kernel, which is now $\alpha + 2$:

$$31 \quad \psi(x \gg \overline{x}) \approx \sin\left(\frac{\pi\alpha}{2}\right) \frac{2^{\alpha-1}\Gamma(1 + \alpha/2)^2}{\pi^2} \frac{D_\alpha}{\rho\mu^2 x^{2+\alpha}}. \quad (8)$$

For nearby individuals, $\delta \ll x \ll \overline{x}$, all values of $\alpha < 2$ dis-
play "now-or-never" behavior similar to that of $\alpha < 1$ in one
dimension. Coalescence tends to occur via one "quick jump",
rather than many small jumps, and the competition is with jump-
ing very far away rather than with mutating. The same intuitive
argument used above applies, now with the lineages spread
over an area of $\sim (D_\alpha t)^{2/\alpha}$ and therefore coalescing at rate
$\sim (D_\alpha t)^{-2/\alpha}/\rho$, for a probability of identity $\psi \sim 1/(\rho D_\alpha x^{2-\alpha})$,
independent of the mutation rate. We calculate this more care-
fully in the Methods, finding:

$$42 \quad \psi(\delta \ll x \ll \overline{x}) \approx \frac{\Gamma(1 - \alpha/2)}{2^\alpha \pi \Gamma(\alpha/2)} \frac{1}{\rho D_\alpha x^{2-\alpha}}. \quad (9)$$

43 Note that (9) reduces to Eq. (A6) in Chave and Leigh Jr (2002)
44 when $\alpha = 1$.

45 Again, (9) must be cut off at very small distances, $x \lesssim \delta$, at
46 which point homozygosity flattens out to its $x = 0$ value:

$$47 \quad \psi(x \ll \delta) \approx 1 \left/ \left[1 + \frac{2^{2+\alpha/2}\pi}{\Gamma(1 - \alpha/2)} \rho D_\alpha \delta^{2-\alpha} \right]. \right. \quad (10)$$

48 Just as in one dimension, the numerical factors in (10) depend on
49 the details of the model and are only valid when $\psi(0) \ll 1$, but
50 the basic scaling is general and arises from the race between coa-
51 lescence at rate $\sim 1/(\rho\delta^2)$ and dispersing outside coalescence
52 range at rate $\sim D_\alpha \delta^{-\alpha}$.

Figure 4 Isolation by distance follows the same power law as dispersal. Each panel shows the scaled probability of identity between a sampled pair of individuals, $\psi\rho\bar{x}\mu$, as a function of the scaled distance $x/\bar{x}$ between them. Points show simulation results, black curves show numerical solutions of $\psi(x)$ calculated from (27) with $\delta = 0$, and magenta lines show the power law that emerges at large distances (4). Red curves show the asymptotic behavior predicted at short distances by (6) ($\alpha < 1$), (5) ($1 < \alpha < 2$), and (40) ($\alpha = 1$). For all plots, error bars show 68% percentile bootstrap confidence intervals (see Methods). $\rho = 100$ in all plots, and data with $\rho = 10$ and $\rho = 1$ (not shown) yield indistinguishable plots.

Figure 5 For very long-range dispersal, $\alpha < 1$, relatedness at short distances is independent of mutation rate. Nearby lineages at $x \ll \bar{x}$ either coalesce quickly and are identical, or jump very far away from each other and never coalesce. Points show simulation results, and red and magenta lines show the asymptotic predictions of (6) and (7), respectively. The black curve shows a numerical solution of $\psi(x)$ calculated from (27) with $\mu = 10^{-4}$. $\rho = 100$ in all plots, and data with $\rho = 10$ and $\rho = 1$ (not shown) yield indistinguishable plots.

Figure 7 The form of isolation by distance in two dimensions is universal at long distances. Approximate form for the probability of identity as a function of distance, $\psi(x)$, for different dispersal kernels α. Different regimes of the parameter space are separated by solid lines, and labelled by their qualitative dynamics. Coalescence for distant pairs, $x \gg \bar{x}$, typically occurs via one long jump, which leads to the power law scaling at large distances predicted by (8). Nearby pairs, $x \ll \bar{x}$, typically either coalesce very quickly or disperse far away from each other, so the probability of identity is nearly independent of the mutation rate, as shown in (9). This quick coalescence is effectively diffusive for $\alpha > 2$, while for $\alpha < 2$, it is typically driven by a single jump. We use "$\sim$" to denote proportionality in the limit of large population density where $\psi(0) \ll 1$.

Figure 6 Even for $\alpha > 2$, relatedness still follows the same power law as dispersal, rather than the diffusive prediction. Points show simulation results with $\rho = 100$; $\rho = 10$ and $\rho = 1$ yield indistinguishable plots. Since the dispersal kernel has finite variance, it approaches a diffusion, and at short distances $x \ll \bar{x}$ the probability of identity is well-approximated by the diffusive prediction (3) (red curve). But at long distances $x \gg \bar{x}$, relatedness is driven by rare long-range jumps, and therefore has the same power-law tail as dispersal, (31) (magenta line).

Discussion

Limited dispersal produces a correlation between spatial and genetic distance (Wright 1946; Malécot 1975; Slatkin 1991; Slatkin and Arter 1991; Slatkin 1993). While most previous models have only considered diffusive dispersal, dispersal can be long-range in many natural populations. Nagylaki (1976) was the first to generalize classic diffusive models of isolation by distance by allowing dispersal distance to have a power law tail (with $1 < \alpha < 2$). This groundbreaking work has largely been neglected; it was last cited by Chave and Leigh Jr (2002), who extended the results to two dimensions for the special case of Cauchy flights ($\alpha = 1$) in a paper modeling ecological diversity. We hope that this paper will reintroduce these results to population genetics now that the field may have sufficient data to apply them. We also extend this previous work by considering Lévy flights for all $\alpha \leq 2$ in both one and two dimensions, as well as steeper power law kernels ($\alpha > 2$) with finite variance. We find that, for all α, long-range dispersal leads to much more long-range relatedness than diffusive dispersal, with relatedness having the same power law tail in distance as the dispersal kernel. This is true even for steep kernels with finite variance. In this case, even though a diffusive approximation can fit the pattern of isolation by distance between nearby individuals, it will greatly underestimate the degree of relatedness between distant individuals.

Standard methods for inferring dispersal from pairwise measures of relatedness or autocorrelations in allele frequency typically assume either short-range, diffusive motion (Rousset 1997, 2000; Robledo-Arnuncio and Rousset 2010; Ringbauer *et al.* 2017; Bradburd *et al.* 2018) (perhaps with recent long-range admixture (Bradburd *et al.* 2016)) or a small number of discrete demes (Slatkin 1991; Whitlock and McCauley 1999; Rousset and Leblois 2011; Petkova *et al.* 2016; Al-Asadi *et al.* 2019; Lundgren and Ralph 2019). Methods using cline theory to infer dispersal from the width of hybrid zones make similar assumptions about the motion of lineages being short-range and diffusive (Barton 1983; Barton and Hewitt 1985; Sotka and Palumbi 2006; Rieux *et al.* 2013; Gagnaire *et al.* 2015; Cayuela *et al.* 2018), and methods for non-stable demographies based on historical biogeography or coalescent theory tend to also assume a small number of discrete demes (Sanmartín *et al.* 2008; Ree and Smith 2008; Hey 2010). Other genetic methods such as parentage analysis are better equipped to infer long-range dispersal on continuous ranges, but these techniques require exhaustive sampling of the population to ensure that the parents of each individual can be located (Adams *et al.* 1992; Jones and Ardren 2003; Bacles and Ennos 2008; Wang and Santure 2009). More recent methods for pollen dispersal have been developed that allow for the inference of long-range dispersal without the need for exhaustive sampling, but knowledge of maternal genotypes for all sampled individuals is still required (Austerlitz *et al.* 2004; Robledo-Arnuncio *et al.* 2006). For plant species where this data is available, our results could serve as the basis for complementary inference methods. While the pollen dispersal methods are focused on inferring the dispersal kernel over a single generation, isolation by distance reflects the history of dispersal over many generations, so a comparison of the results could reveal changes in dispersal over time.

One key open question is to what extent it is possible to detect the genetic traces of rare long-range dispersal in natural populations, and if so how well the form of long-range dispersal (e.g., the tail exponent α) can be determined. Austerlitz *et al.* (2004) were able to detect long-range pollen dispersal in tree populations with the seed-specific TwoGener method, which suggests that similar patterns should be detectable from the genomes of the trees themselves. Aguillon *et al.* (2017) found clear patterns of isolation by distance across scales between 500 meters and 10 kilometers in Florida Scrub-Jays, and directly measured long-range dispersal; a good first test of an inference method would be to apply it to such a dataset to see if it can recover the known dispersal pattern.

Along with predicting characteristic scaling of identity by descent with distance, our results predict characteristic scalings with the mutation rate μ, and also a scaling of the typical length scale of identity $\bar{x}$ with μ. While mutation rate cannot be scanned directly as distance can, μ here should be understood as referring to the mutation rate in a block of non-recombining genome, and so a wide range of μ values can be scanned by considering identity by descent in blocks of varying size (Weissman and Hallatschek 2017). This will be valid as long as recombination is rare relative to mutation, or even if recombination is frequent as long as "μ" is understood to mean the sum of the block mutation and recombination rates. This suggests that it should be possible to measure identity by descent statistics corresponding to "μ" values ranging over five orders of magnitude in a single sample (Harris and Nielsen 2013).

Our analysis has focused on the lineages of a particular pair of individuals sampled a fixed distance apart in an infinitely large habitat. We can also consider a pair of individuals sampled from random locations from a habitat with finite length L; the mean coalescence time between them would then be the "effective population size". The pair will typically be sampled a distance $\sim L$ from each other, and so it will typically take a time $\sim L^\alpha / D_\alpha$ for their lineages to overlap in space. At this point the ancestry is effectively well-mixed, and it takes time proportional to the total population size $\sim L^d \rho$, where $d = 1$ or 2 is the dimension of the habitat. For $L^\alpha / D_\alpha \ll L^d \rho$, the mixing time has little effect, while for $L^\alpha / D_\alpha \gtrsim L^d \rho$ spatial structure substantially increases genetic diversity. For short-range dispersal, $\alpha = 2$, structure will be strong in a one dimensional habitat of length $L \gtrsim D \rho$ (Maruyama 1971), while in two dimensions its strength depends only on the local neighborhood size $D \rho$ (Maruyama 1972). The amount of population structure thus either increases with the spatial extent of the population (at fixed density) or is insensitive to it. With long-range dispersal, however, we see a new qualitative pattern. For $\alpha < d$, i.e., for any amount of long-range dispersal in two dimensions, the effect of structure on overall diversity counterintuitively becomes *weaker* as the range size L grows, because L^α / D_α grows more slowly than $L^d \rho$. It seems plausible that in many natural populations, spatial structure may be strong over short time scales (as measured by $\psi(x)$ at large μ) while having little effect on the average coalescence time.

Our use of stable distributions for the dispersal kernel has been partly motivated by the fact that any isotropic single-generation dispersal kernel will converge to a stable one if it is repeated over many independent generations. But as we have noted, this is only true asymptotically, and in any real population there will be correlations across generations, spatial inhomogeneities, shifts in dispersal over time, limits due to finite range size, and many other effects that cannot be captured by a stable distribution. It is therefore better to see it as a simple null model, one step closer to reality than the purely diffusive one, that can serve as a background against which to measure

all these other effects.

What other processes could produce similar patterns to long-range dispersal? One obvious one is if individuals are performing something more like a "Lévy walk" than a Lèvy flight, in which dispersal in any one generation is short but can be correlated across many generations (Zaburdaev *et al.* 2015). Such an effect can be produced at the level of alleles by hitchhiking on beneficial substitutions (Allman and Weissman 2018). But this should be readily distinguishable from neutral long-range dispersal by considering the distribution of relatedness across multiple individuals and loci – hitchhiking will produce long-range relatedness at the same few loci across all individuals, whereas neutral effects will be more evenly distributed. It is an open question whether other neutral processes, in particular demographic fluctuations, might produce similar patterns.

Methods

Simulation Methods

All simulation code and displayed data are available at https://github.com/weissmanlab/Long_Range_Dispersal. We simulate our model in two stages. First, for each value of present-day separation x, dispersal constant D_α, and tail parameter α, we simulate dispersal of the lineages, ignoring coalescence and mutation. Then, for each value of ρ and μ, we calculate the expected homozygosity and coalescence time distribution for each simulated trajectory. We then average over many independent trajectories. A major advantage of this two part method is that the same dispersal simulations can be used for calculating the homozygosity and coalescence time distribution for multiple choices of ρ and μ. Another advantage is that the second part of the method, in which conditional expectations are calculated for previously generated paths, is entirely deterministic. This reduces computational costs and noise in the estimations.

We simulate lineage motion using a discrete time random walk,

$$X_{t+1} = X_t + \Delta X_t, \tag{11}$$

where X_t represents the signed distance between two lineages at a given time (ignoring coalescence., i.e., assuming $\rho \to \infty$), and the step size, ΔX_t, is a real valued random variable drawn from the dispersal distribution at each integer time t. For dispersal, we primarily use Lévy alpha-stable distributions, so ΔX_t has distribution:

$$K(y) = \frac{1}{2\pi} \int_{-\infty}^{\infty} dk \exp\left(-iky - 2D_\alpha |k|^\alpha\right). \tag{12}$$

Note that this differs from (1) because $\Delta t = 1$, and there is an extra factor of two because ΔX_t is the sum of the two lineages' independent jumps, i.e., K is the convolution of K_1 with itself. To simulate steeper tails with $\alpha > 2$, we use an F-distribution, defined below in (15). We use the GNU Scientific Library's efficient pseudorandom generators for both stable distributions and the F-distribution (Galassi *et al.* 2009); because these are available only for the one-dimensional distributions, all our simulations are in one dimension.

For each simulated trajectory $\{x_t\}$, we then compute the conditional distribution of coalescence times $p(t|\{x_{t'\leq t}\})$ and

conditional mean homozygosity $\psi(x|\{x_{t'\leq t}\})$:

$$p(t \geq 1|\{x_{t'\leq t}\}) = \left(1 - e^{-\frac{1}{\rho}R(x_t)}\right) \exp\left[-\frac{1}{\rho}\sum_{t'=1}^{t-1} R(x_{t'})\right], \tag{13}$$

$$\psi(x|\{x_{t'\leq t}\}) = \sum_{t=1}^{\infty} p(t|\{x_{t'\leq t}\})e^{-2\mu t}. \tag{14}$$

We start (13) and (14) at $t = 1$ because we assume that the individuals are sampled immediately after dispersal, so no coalescence takes place at $t = 0$. $R(x)$ in (13) is a rectangular function representing a uniform rate of coalescence of all lineages within a distance δ:

$$R(x) \equiv \begin{cases} \frac{1}{2\delta} & \text{if } |x| < \delta \\ 0 & \text{otherwise.} \end{cases}$$

For every time-step the lineages spend in this region, there is a probability of coalescence $1 - e^{-\frac{1}{2\rho\delta}}$. We discuss issues with the microscopic interpretation of this model after we introduce our analytical model below.

To obtain the unconditioned values $p(t|x)$ and $\psi(x)$, we then average (13) and (14) across all simulated trajectories. All error bars in plots show 68% confidence intervals, as determined by the percentile bootstrap with 10,000 bootstrap samples (Davison and Hinkley 1997). At large distances, the distribution of the probability of identity across sample trajectories is highly skewed, with most trajectories having very low probabilities of identity, but a few having the lineages rapidly jump close to each other and having a high probability of identity. This means that we cannot quantify the uncertainty in our estimates using, for example, the standard error of the mean, but it also means that we must simulate many independent trajectories to get good enough coverage for the bootstrap to be accurate (Chernick 2011).

We set $\delta = 0.5$ for all simulations. For the simulations of mean homozygosity ψ shown in Fig. 4 and Fig. 6, we simulate 250,000 independent runs of 1000 generations each for each combination of initial separation x and tail parameter α. We set the dispersal constant D_α indirectly by setting the characteristic spread c of two lineages after one generation ($t = 1$), $c = (2D_\alpha)^{1/\alpha}$, to be fixed at $c = 250$ for $\alpha < 2$, and $c = 179.68$ for $\alpha = 2.05$ (see below for the definition of c for $\alpha > 2$). For the largest initial separations, $x = e^{10}$ and e^{11}, coalescence within 1000 generations is very rare, so we increase the number of runs to 1.25×10^6. For Fig. 5 and Fig. 9, we choose D_α such that $c = 0.2$, and simulate 10,000 independent runs of length 1000 generations each.

For the simulations of the cumulative distribution of coalescence times $P(t)$ shown in Fig. 8, we set initial separation $x = 0$ and generate 10,000 independent trajectories of 1.5 million generations each for each combination of c and tail parameter α. We set $c = 3.59$ for $\alpha > 2$, $c = 5$ for $1 < \alpha < 2$, and $c = 1$ for $\alpha \leq 1$.

Dispersal kernel for $\alpha > 2$

To simulate dispersal kernels with tail exponents $\alpha > 2$, we draw ΔX_t from a Fisher F-distribution:

$$K(y) = \frac{\Gamma(2\alpha)}{\Gamma^2(\alpha)\omega}(y/\omega)^{\alpha-1}\left(1 + (y/\omega)\right)^{-2\alpha}. \tag{15}$$

At long times, the displacement distribution approaches that of a diffusive kernel, with dispersal constant equal to half the mean squared single-generation displacement of one lineage:

$$D = \frac{c^2}{2} = \left(\frac{\alpha(2\alpha-1)}{4(\alpha-2)(\alpha-1)^2} + \frac{\alpha^2}{4(\alpha-1)^2}\right)\omega^2.$$

Analytical model in one dimension

Generic dispersal

We want to find a tractable analytical approximation to the model described above. For recurrent motion, the lineages will sometimes be in exactly the same place, and we can model coalescence with a δ distribution, i.e., as taking place at rate $\frac{1}{\rho}\delta(X_t)$. For transient motion, however, they will never coincide (Palyulin *et al.* 2014), and we must allow coalescence to take place at a finite distance. Let the coalescence kernel be some normalized distribution $\mathcal{N}(x)$ symmetric about $x = 0$ and with width $\sim \delta$, with coalescence taking place at rate $\frac{1}{\rho}\mathcal{N}(X_t)$. The δ-distribution is just the limit of $\mathcal{N}$ as δ goes to 0, so we can treat the two cases together. Forien (2019) avoids this issue by using a spatial Λ-Fleming-Viot model, in which dispersal and coalescence are produced by the same long-range events, but we wish to keep long-range dispersal as a distinct process from short-range coalescence. As mentioned in the Simulation Model section, this creates issues with the microscopic interpretation of the model, which we discuss below in "Breakdown of models at small scales".

The coalescence time distribution is then:

$$p(t|x) = E\left[\frac{1}{\rho}\mathcal{N}(X_t)\exp\left(-\frac{1}{\rho}\int_0^t d\tau\,\mathcal{N}(X_\tau)\right)\bigg| X_0 = x\right], \quad (16)$$

and the probability of identity is its Laplace transform:

$$\psi(x) = \int_0^\infty dt\, p(t|x)e^{-2\mu t}. \quad (17)$$

There are several different ways to derive an explicit expression for ψ from (17), including balancing mutation, coalescence, and dispersal over an infinitesimal time step (Malécot 1975; Barton *et al.* 2002) or, for Lévy flights, using a fractional diffusion equation (Janakiraman 2017) (see Appendix). Here we start with a generalization of Barton and Wilson (1995)'s expression for $p(t|x)$ that is valid for any two-lineage dispersal kernel K, not necessarily a Lévy flight, assuming that $\mathcal{N}(x) = \delta(x)$:

$$p(t|x) = \frac{1}{\rho}K(x|t) - \int_0^t dt'\, p(t - t'|0)\frac{1}{\rho}K(x|t'). \quad (18)$$

To interpret (18), notice that the first term is the probably of coalescing at time t neglecting the possibility that the lineages have coalesced more recently. The second term corrects for these more recent coalescences: for every trajectory where the lineages coincide at $t' < t$, we subtract off the probability that the lineages would coalesce at t' and then again exactly at t. Notice that we do not need to correct again for lineages that coincide three times, at $t'' < t' < t$: the factor of p guarantees that each trajectory is weighted appropriately.

We can immediately find a simple expression for ψ for recurrent dispersal by taking the Laplace transform of (18):

$$\psi(x) = \frac{1 - \psi(0)}{\rho}\widetilde{K}(x, 2\mu), \quad (19)$$

where tilde denotes the Laplace transform. Plugging in $x = 0$, we can solve (19) for $\psi(0)$ and express $\psi(x)$ purely in terms of the dispersal kernel:

$$\psi(x) = \frac{\widetilde{K}(x, 2\mu)}{\rho + \widetilde{K}(0, 2\mu)}. \quad (20)$$

Note that, for diffusive dispersal, (19) reduces to the classical Wright-Malécot formula for isolation by distance (Barton *et al.* 2002).

For transient dispersal, we must consider a coalescence kernel of finite width, and (18) generalizes to:

$$\begin{aligned} p(t|x) = {} & \int dy \frac{1}{\rho}\mathcal{N}(y)K(x - y|t) \\ & - \int_0^t dt'\int dy\, p(t - t'|y)\frac{1}{\rho}\mathcal{N}(y)K(x - y|t'). \end{aligned} \quad (21)$$

(21) is exactly the same as (18) except that now we must integrate over possible locations y of coalescence at both t and t'. Taking the Laplace transform of (21) now gives:

$$\psi(x) = \frac{1}{\rho}\int dy(1 - \psi(y))\mathcal{N}(y)\widetilde{K}(x - y, 2\mu). \quad (22)$$

To simplify (22), we can make the ansatz that $1 - \psi(y)$ is nearly constant over all separations $|y| \lesssim \delta$ where $\mathcal{N}(y)$ is non-negligible, allowing us to pull it out of the integral:

$$\psi(x) \approx \frac{1 - \psi(0)}{\rho}\int dy\mathcal{N}(y)\widetilde{K}(x - y, 2\mu). \quad (23)$$

This ansatz will necessarily be accurate when identity is low, $\psi(0) \ll 1$ because $1 - \psi$ will be close to 1 for all y. However, for $1 - \psi(0) \ll 1$, the approximation can become inaccurate; we discuss this below. At long distances $x \gg \delta$, $\widetilde{K}$ will also be roughly constant in the integral, and we simply recover (19), although now only as an approximation:

$$\psi(x \gg \delta) \approx \frac{1 - \psi(0)}{\rho}\widetilde{K}(x, 2\mu). \quad (24)$$

We see that the details of the short-range behavior only affect the long-range probability of identity by descent through the overall factor $1 - \psi(0)$ (Barton *et al.* 2002). Mathematically, the main challenge is to find simple expressions for $\psi(0)$ and especially $\widetilde{K}$.

Because (24) is invalid for $x = 0$, we cannot solve it directly for $\psi(0)$ as we could with (19), and so we must also work with (23). We can simplify the convolution in (23) by taking the spatial Fourier transform $\mathcal{F}\{\cdot\}$:

$$\widehat{\psi}(k) \approx \frac{1 - \psi(0)}{\rho}\widehat{\mathcal{N}}(k)\widehat{\widetilde{K}}(k, 2\mu). \quad (25)$$

where $\widehat{\psi}$ and $\widehat{\widetilde{K}}$ are the Fourier transforms of ψ and $\widetilde{K}$.

Lévy flight dispersal

For Lévy flights, the characteristic function is $\widehat{K}(k|t) = \exp(-2D_\alpha t|k|^\alpha)$ and the Fourier-Laplace transform is $\widehat{\widetilde{K}}(k, 2\mu) = 1/(2\mu + 2D_\alpha|k|^\alpha)$. (25) for $\widehat{\psi}$ is correspondingly simple:

$$\frac{\widehat{\psi}(k)}{1 - \psi(0)} \approx \frac{\widehat{\mathcal{N}}(k)}{2\rho(\mu + D_\alpha|k|^\alpha)}. \quad (26)$$

To get an explicit expression for ψ, we need to specify a form for the coalescence kernel $\mathcal{N}$. We will use a normal distribution with standard deviation δ, which has the simple Fourier transform $\widehat{\mathcal{N}}(k) = \exp(-\delta^2 k^2/2)$. Then we can invert the Fourier transform in (26):

$$\frac{\psi(x)}{1 - \psi(0)} \approx \frac{1}{2\pi\rho}\int_0^\infty dk\frac{\cos(kx)e^{-\delta^2 k^2/2}}{\mu + D_\alpha k^\alpha}, \quad (27)$$

which can be re-expressed in dimensionless units as

$$\frac{\psi(x)}{1-\psi(0)} \approx \frac{1}{2\pi\rho\mu\bar{x}} \int_0^\infty d\kappa \frac{\cos(\kappa x/\bar{x})e^{-(\delta/\bar{x})^2\kappa^2/2}}{1+\kappa^\alpha}. \quad (28)$$

Examining (27), we see that the power law tail in the integrand can be cut off either when oscillations in the cosine factor become rapid at $k \sim 1/x$ or by the normal factor at $k \sim 1/\delta$. As long as we are sampling pairs that are outside the immediate range of coalescence, $x \gg \delta$, the former cutoff will happen at lower k, and therefore the normal factor can be neglected, leaving (in dimensionless form):

$$\frac{\psi(x \gg \delta)}{1-\psi(0)} \approx \frac{1}{2\pi\rho\mu\bar{x}} \int_0^\infty d\kappa \frac{\cos(\kappa x/\bar{x})}{1+\kappa^\alpha}. \quad (29)$$

(29) can equivalently be derived directly from (24) by substituting in the Lévy flight dispersal kernel and writing $\widetilde{K}$ as the inverse Fourier transform of $\widehat{\widetilde{K}}$.

We can solve (28) for $\psi(x)$ by first evaluating it at $x = 0$ to find $\psi(0)$; we do this below. But it is interesting that the ratio $\Psi(x) \equiv \psi(x)/(1-\psi(0))$ has the simplest relationship to the underlying parameters, as shown by Rousset (1997) for short-range dispersal. Ψ is closely related to Rousset (2000)'s statistic a_r: $a_r = \Psi(0) - \Psi(r)$. It is also related to the expected pairwise F_{ST} between demes separated by x:

$$E\left[F_{\text{ST}}(x)\right] = \frac{\Psi(0) - \Psi(x)}{2 + \Psi(0) - \Psi(x)}. \quad (21)$$

Probability of identity for distant pairs $x \gg \bar{x}$, $\alpha < 2$
For large $x \gg (D_\alpha t)^{1/\alpha}$, the dispersal kernel has a simple asymptotic form for $\alpha < 2$ (Nolan (2018), Theorem 1.12):

$$K(x \gg (D_\alpha t)^{1/\alpha}|t) \approx \frac{2\Gamma(\alpha+1)}{\pi} \sin\left(\frac{\pi\alpha}{2}\right) \frac{D_\alpha t}{x^{\alpha+1}}.$$

Plugging this into (24) and evaluating the Laplace transform gives (4), which was originally found by Nagylaki (1976):

$$\frac{\psi(x \gg \bar{x})}{1-\psi(0)} \approx \frac{\Gamma(\alpha+1)}{2\pi} \sin\left(\frac{\pi\alpha}{2}\right) \frac{D_\alpha}{\rho\mu^2 x^{\alpha+1}}. \quad (30)$$

Probability of identity for distant pairs $x \gg \bar{x}$, $\alpha > 2$
There is no stable distribution with $\alpha > 2$, but in discrete-time models such as the one we use in our simulations, we can consider single-generation jump kernels $K(y|1)$ with power-law tails with $\alpha > 2$. These will approach a diffusion with diffusion constant $D = \text{Var}(K)/2$. At long distances $y \gg \sqrt{Dt}$, however, the tail will still be dominated by the probability of taking a single large jump (Vezzani et al. 2019), so for $x \gg \bar{x}$, we will have $K(x|t \lesssim 1/\mu) \approx K(x|1)t$. Plugging this into (24) and evaluating the Laplace transform gives:

$$\frac{\psi(x \gg \bar{x})}{1-\psi(0)} \approx \frac{K(x|1)}{4\rho\mu^2}. \quad (39)$$

For the F-distribution kernel (15) used in the simulations, this is

$$\frac{\psi(x \gg \bar{x})}{1-\psi(0)} \approx \frac{\Gamma(2\alpha)}{4\Gamma^2(\alpha)} \frac{D_\alpha}{\rho\mu^2 x^{\alpha+1}}, \quad (31)$$

where we have defined $D_\alpha \equiv \omega^\alpha/\text{generation}$. (31) is confirmed by simulations (Fig. 6). We can then use the classic diffusive expression for $\psi(0)$ to get an explicit expression for probability of identity at large distances:

$$\psi(x \gg \bar{x}) \approx \left(1 + \frac{1}{4\rho\bar{x}\mu}\right)^{-1} \frac{\Gamma(2\alpha)}{4\Gamma^2(\alpha)} \frac{D_\alpha}{\rho\mu^2 x^{\alpha+1}}. \quad (32)$$

Moderately long-range dispersal, $1 < \alpha < 2$
For $\alpha > 1$, (19) and (29) are exact for all x. Evaluating (29) for $x = 0$ gives $\psi(0)$:

$$\psi(0) = \frac{1}{2\alpha \sin(\pi/\alpha)\rho\bar{x}\mu + 1}. \quad (33)$$

Plugging (33) into (30) lets us solve for $\psi(x)$ at large distances $x \gg \bar{x}$:

$$\psi(x \gg \bar{x}) \approx \frac{\sin(\pi\alpha/2)\Gamma(\alpha+1)/(2\pi)}{1 + 1/(2\alpha \sin(\pi/\alpha)\rho\bar{x}\mu)} \frac{D_\alpha}{\rho\mu^2 x^{\alpha+1}}. \quad (34)$$

For $0 < x \ll \bar{x}$, Janakiraman (2017) (Eq. (C1)) found that to leading order ψ falls off as:

$$\psi(x \ll \bar{x}) \approx \psi(0) \left[1 - \frac{\alpha \sin(\pi/\alpha)}{\Gamma(\alpha)\cos(\pi(1-\alpha/2))} \left(\frac{x}{\bar{x}}\right)^{\alpha-1}\right]. \quad (35)$$

When $\alpha = 2$, the above expression is equivalent to the classic diffusive result (3) for $x \ll \bar{x}$, which can be found by integrating (22) with $\delta = 0$.

Very long-range dispersal, $\alpha < 1$
For $\alpha < 1$, the finite width δ of the coalescence kernel is important for determining $\psi(0)$. Dropping the cosine factor in (28) gives:

$$\frac{\psi(0)}{1-\psi(0)} \approx \frac{1}{2\pi\rho\mu\bar{x}} \int_0^\infty d\kappa \frac{e^{-(\delta/\bar{x})^2\kappa^2/2}}{1+\kappa^\alpha}$$

$$\approx \frac{\Gamma(1/2 - \alpha/2)}{2^{(\alpha+3)/2}\rho D_\alpha \delta^{1-\alpha}},$$

where in evaluating the integral we have assumed that $\delta \ll \bar{x}$, i.e., that the mutation rate is not extremely large. We see that on small scales, the probability of identity by descent is independent of the mutation rate, i.e., there is a large probability that individuals from the same deme are differentiated even for infinitesimal mutation rates:

$$\psi(0) \approx 1 \left/ \left[1 + \frac{2^{(\alpha+3)/2}\pi}{\Gamma(1/2-\alpha/2)}\rho D_\alpha \delta^{1-\alpha}\right]\right. \quad (36)$$

Very long-range dispersal of nearby lineages causes them to quickly wander away from each other, and for infinite range size many pairs will never coalesce. While (36) is only accurate for $\psi(0) \ll 1$, the independence from mutation rate should persist even for large $\psi(0)$.

Plugging (36) for $\psi(0)$ into (30) gives an explicit expression for the probability of identity of distant pairs with $x \gg \bar{x}$:

$$\psi(x \gg \bar{x}) \approx \frac{\Gamma(\alpha+1)\sin(\pi\alpha/2)/(2\pi)}{1 + \Gamma(1/2-\alpha/2)/(2^{(\alpha+3)/2}\pi\rho D_\alpha \delta^{1-\alpha})} \frac{D_\alpha}{\rho\mu^2 x^{\alpha+1}}. \quad (80)$$

For pairs that are nearby but still well outside of coalescence range, $\delta \ll x \ll \bar{x}$, the integral in (29) is dominated by $\kappa \gg 1$ and is approximately:

$$\frac{\psi(\delta \ll x \ll \bar{x})}{1-\psi(0)} \approx \frac{\Gamma(1-\alpha)\sin(\pi\alpha/2)}{2\pi} \frac{x^{\alpha-1}}{\rho D_\alpha}. \quad (84)$$

Again, the probability of identity is independent of the mutation rate to lowest order. Substituting in (36) gives an explicit expression for ψ:

$$\psi(\delta \ll x \ll \bar{x}) \approx \frac{\Gamma(1-\alpha)\sin(\pi\alpha/2)/(2\pi)}{1 + \Gamma(1/2-\alpha/2)/(2^{(\alpha+3)/2}\pi\rho D_\alpha \delta^{1-\alpha})} \frac{x^{\alpha-1}}{\rho D_\alpha}. \quad (37)$$

Marginal case $\alpha = 1$

The analysis of the marginal case $\alpha = 1$ is essentially the same as for $\alpha < 1$ above, but we have separated it out because the form of the final expressions is very different. As with $\alpha < 1$, the finite coalescence width δ is important for $x = 0$:

$$\frac{\psi(0)}{1 - \psi(0)} \approx \frac{1}{2\pi\rho D_1} \int_0^\infty dk \frac{e^{-(\delta/\overline{x})^2 \kappa^2/2}}{1 + \kappa}$$
$$= \frac{2\ln(\overline{x}/\delta) + \ln 2 - \gamma}{4\pi D_1},$$

where $\gamma \approx 0.58$ is Euler's constant. Again assuming $\delta \ll \overline{x} = D_1/\mu$, the constant terms in the numerator can be neglected and ψ is approximately:

$$\psi(0) \approx \left(1 + \frac{2\pi\rho D_1}{\ln(\overline{x}/\delta)}\right)^{-1}. \tag{38}$$

Recall that the ansatz we used to derive (38) is only justified when $\psi(0) \ll 1$.

For pairs that are nearby but still well outside of coalescence range, $\delta \ll x \ll \overline{x}$, (29) gives:

$$\frac{\psi(\delta \ll x \ll \overline{x})}{1 - \psi(0)} \approx \frac{\ln(\overline{x}/x) - \gamma}{2\pi\rho D_1}. \tag{39}$$

Plugging the expression (38) for $\psi(0)$ into (39) and (30) gives explicit expressions for $\psi(x)$ at both short and long distances:

$$\psi(\delta \ll x \ll \overline{x}) \approx \frac{\ln(\overline{x}/x)}{2\pi\rho D_1 + \ln(\overline{x}/\delta)} \tag{40}$$

$$\psi(x \gg \overline{x}) \approx \frac{1}{2\pi\rho D_1 + \ln(\overline{x}/\delta)} \left(\frac{\overline{x}}{x}\right)^2. \tag{41}$$

Analytical model in two dimensions

Generic dispersal

For generic dispersal, the solution for ψ in two dimensions can again be found from (22), now with the integral over two spatial dimensions. The Fourier transform $\widehat{\psi}$ has the same form as the one-dimensional equation (25):

$$\widehat{\psi}(k) \approx \frac{1 - \psi(0)}{\rho} \widehat{\mathcal{N}}(k) \widehat{\widetilde{K}}(k, 2\mu), \tag{42}$$

where again we make the ansatz that $1 - \psi(x)$ is approximately constant over the x values where $\mathcal{N}(x)$ is non-negligible. This is again accurate for $\psi(0) \ll 1$, but may need to be adjusted for $1 - \psi(0) \ll 1$. While (42) looks exactly like the one dimensional expression (25), its interpretation is different: k is now the radial coordinate in two-dimensional k-space, and if we want to transform back to real space, we must use the two-dimensional inverse Fourier transform. For pairs that are far outside coalescence range, $x \gg \delta$, the simple relation (24) between $\psi(x)$ and $\widetilde{K}(x, 2\mu)$ still holds.

Lévy flight dispersal

For a two-dimensional Lévy flight, the dispersal kernel takes the form of an isotropic stable distribution (Zolotarev 1981):

$$K(y|t) = \frac{1}{2\pi} \int_0^\infty dk\, k J_0(ky) \exp\left(-2D_\alpha t k^\alpha\right), \tag{43}$$

where $K(y|t)$ is the probability density of being at a particular point a distance y away from the position at time 0, and

J_0 is the zeroth Bessel function of the first kind. (43) is the two-dimensional inverse Fourier transform (equivalently, the inverse zeroth-order Hankel transform) of the characteristic function $\widehat{K}(k|t) = \exp\left(-2D_\alpha t k^\alpha\right)$. The Fourier-Laplace transform is again $\widehat{\widetilde{K}}(x, 2\mu) = 1/(2\mu + 2D_\alpha|k|^\alpha)$. At large distances, $y \gg (D_\alpha t)^{1/\alpha}$, K has a power-law tail (Nolan 2013):

$$K(y \gg (D_\alpha t)^{1/\alpha}|t) \approx \frac{\alpha^2 \Gamma(\alpha/2)^2}{2^{1-\alpha}\pi^2} \sin\left(\frac{\pi\alpha}{2}\right) \frac{D_\alpha t}{x^{\alpha+2}}. \tag{44}$$

In two dimensions, we must allow coalescence to take place at a finite distance for all α (Mörters and Peres 2010). For the coalescence kernel, we use an isotropic normal distribution $\mathcal{N}(x)$ with mean zero and standard deviation δ, with coalescence taking place at rate $\frac{1}{\rho}\mathcal{N}(X_t)$. Inverting the Fourier transform in (42) then gives:

$$\frac{\psi(x)}{1 - \psi(0)} \approx \frac{1}{4\pi\rho} \int_0^\infty dk \frac{k J_0(kx) e^{-\delta^2 k^2/2}}{\mu + D_\alpha k^\alpha} \tag{45}$$

$$= \frac{1}{4\pi\rho\mu\overline{x}^2} \int_0^\infty d\kappa \frac{\kappa J_0(\kappa x/\overline{x}) e^{-(\delta/\overline{x})^2 \kappa^2/2}}{1 + \kappa^\alpha}, \tag{46}$$

The analysis of (46) parallels that of the one-dimensional case, but all $\alpha < 2$ can be treated together for all distances x, not just $x \gg \overline{x}$, and so we can conduct one unified analysis moving from short distances to long ones.

Probability of identity for co-located pairs, $x = 0$

For pairs sampled from the same location, $x = 0$, the Bessel function in (46) is simply equal to one and can be dropped:

$$\frac{\psi(0)}{1 - \psi(0)} \approx \frac{1}{4\pi\rho\mu\overline{x}^2} \int_0^\infty d\kappa \frac{\kappa e^{-(\delta/\overline{x})^2 \kappa^2/2}}{1 + \kappa^\alpha}$$
$$\approx \frac{\Gamma(1 - \alpha/2)}{2^{2+\alpha/2}\pi\rho D_\alpha \delta^{2-\alpha}}, \tag{47}$$

where in the last line we have assumed that $\delta \ll \overline{x}$. Intuitively, (47) can be understood as roughly the ratio between the time to coalesce, i.e., the neighborhood size $\sim \rho\delta^2$ and the time $\sim \delta^\alpha/D_\alpha$ that the lineages will spend in the same neighborhood before jumping apart. Note that mutation does not enter: in two dimensions, all $\alpha < 2$ act like $\alpha < 1$ does in one dimension, where locally mutation is irrelevant. Again, (47) is only accurate for $\psi(0) \ll 1$.

Solving (47) for ψ gives:

$$\psi(0) \approx \left(1 + \frac{2^{2+\alpha/2}\pi}{\Gamma(1 - \alpha/2)} \rho D_\alpha \delta^{2-\alpha}\right)^{-1}. \tag{48}$$

For $\alpha = 2$, integrating (46) with $x = 0$ recovers the classic diffusive result in two dimensions, which we expect to hold for pairs in contact when $\alpha \geq 2$ (Barton et al. 2002):

$$\psi(0) \approx \frac{\ln(\overline{x}/\delta)}{\ln(\overline{x}/\delta) + 4\pi\rho D_2}. \tag{49}$$

Probability of identity for separated but nearby pairs, $\delta \ll x \ll \overline{x}$

For pairs that are outside coalescence range, $x \gg \delta$, we can find ψ from (24):

$$\frac{\psi(\delta \ll x \ll \overline{x})}{1 - \psi(0)} \approx \frac{1}{4\pi\rho\mu\overline{x}^2} \int_0^\infty d\kappa \frac{\kappa J_0(\kappa x/\overline{x})}{1 + \kappa^\alpha}. \tag{50}$$

This looks different from the one-dimensional equation (27) because now we had to apply the two-dimensional inverse Fourier transform to $\widetilde{\widehat{K}}$ to obtain $\widetilde{K}$. For nearby pairs $x \ll \overline{x}$, the integral in (50) is dominated by $\kappa \gg 1$ and for $\alpha < 2$ we can approximate the denominator in the integrand as $1 + \kappa^\alpha \approx \kappa^\alpha$, giving:

$$\frac{\psi(\delta \ll x \ll \overline{x})}{1 - \psi(0)} \approx \frac{\Gamma(1 - \alpha/2)}{\Gamma(\alpha/2)2^\alpha \pi \rho D_\alpha} x^{\alpha - 2}. \quad (51)$$

The convergence of (50) to (51) is however quite slow in $x/\overline{x}$ when α is close to 0 or 2. For example, for $x/\overline{x} = 0.01$, the two expressions differ by $\approx 30 - 40\%$ for $\alpha = 0.25$ and $\alpha = 1.75$, and only approach to within 10% of each other at extreme values of $x/\overline{x}$ ($\approx 10^{-5}$ and $\approx 10^{-4}$ for $\alpha = 0.25$ and $\alpha = 1.75$, respectively).

Plugging (48) for $\psi(0)$ into (51) lets us solve for ψ:

$$\psi(\delta \ll x \ll \overline{x}) \approx \left(1 + \frac{\Gamma(1 - \alpha/2)}{2^{2+\alpha/2}\pi \rho D_\alpha \delta^{2-\alpha}}\right)^{-1} \frac{\Gamma(1 - \alpha/2)}{\Gamma(\alpha/2)2^\alpha \pi} \frac{x^{\alpha-2}}{\rho D_\alpha}. \quad (52)$$

We see that in two dimensions, relatedness at short distances is independent of μ to leading order for all $\alpha < 2$. However, the slow convergence mentioned above means that for most biologically reasonable parameter values, this should be interpreted as meaning that the dependence on mutation rate is weak rather than negligible.

For $\alpha = 1$ and $\delta \ll x \ll \overline{x}$ we recover Eq. (A6) of Chave and Leigh Jr (2002) for Cauchy dispersal. Note that they consider distances large compared to the typical single-generation dispersal distance, $c \equiv (2D_\alpha)^{1/\alpha}$, but small compared to $\overline{x}$, and thus our result for $\delta \ll x \ll \overline{x}$ is consistent with their findings.

For $\alpha = 2$ and $\delta \ll x \ll \overline{x}$ we can recover the known result for diffusive motion by approximating (46) as

$$\frac{\psi(\delta \ll x \ll \overline{x})}{1 - \psi(0)} \approx \frac{1}{4\pi \rho \mu \overline{x}^2} \int_0^\infty d\kappa \frac{J_0(\kappa x/\overline{x})\kappa}{1 + \kappa^2}. \quad (53)$$

Integrating (53) confirms that we find logarithmic scaling of $\psi(x)$ at short distances (Barton *et al.* 2002):

$$\frac{\psi(x)}{1 - \psi(0)} \approx \frac{1}{4\pi \rho D_2} \ln(\overline{x}/x), \quad (54)$$

which we expect to hold at $\delta \ll x \ll \overline{x}$ for all $\alpha \geq 2$.

Probability of identity by descent for distant pairs, $x \gg \overline{x}$

The probability of identity by descent for distant pairs $x \gg \overline{x} \gg \delta$ can be immediately be read off from (24) by substituting in the tail of the two-dimensional dispersal kernel (44) for $\alpha < 2$:

$$\frac{\psi(x \gg \overline{x})}{1 - \psi(0)} \approx \frac{\alpha^2 \Gamma(\alpha/2)^2}{2^{3-\alpha}\pi^2} \sin\left(\frac{\pi\alpha}{2}\right) \frac{D_\alpha}{\rho\mu^2} x^{-\alpha-2}. \quad (55)$$

Plugging in (48) for $\psi(0)$ lets us solve for ψ:

$$\psi(x \gg \overline{x}) \approx \frac{\alpha^2 \Gamma(\alpha/2)/[2^{3-\alpha}\pi\Gamma(1-\alpha/2)]}{1 + \Gamma(1-\alpha/2)/[2^{2+\alpha/2}\pi\rho D_\alpha \delta^{2-\alpha}]} \frac{D_\alpha x^{-\alpha-2}}{\rho\mu^2}. \quad$$

$$\quad (frac)$$

When $\alpha = 2$, we instead recover classic expression for two dimensional diffusive motion at large distances (Barton *et al.* 2002):

$$\frac{\psi(x \gg \overline{x})}{1 - \psi(0)} \approx \frac{1}{4\rho D_2} \frac{\exp(-x/\overline{x})}{\sqrt{2\pi x/\overline{x}}}. \quad (56)$$

Coalescence time distribution

In this section we will find asymptotic expressions for the coalescence time distribution. As stated in the Results, intuitively we can think of the probability of identity ψ as measuring the probability of the pair of lineages coalescing $\lesssim 1/(2\mu)$ generations ago. We can make this statement more rigorous using the Hardy-Littlewood tauberian theorem connecting the long (short) time probability of coalescence to the small (large) mutation rate limit of ψ. It states that a function $f(t)$ has the limiting behavior $f(t) \to \frac{1}{\Gamma(\beta)}t^{\beta-1}L(t)$ as $t \to \infty$ ($t \to 0$), where L is a slowly varying function and $\beta > 0$, if and only if its Laplace transform $\widetilde{f}(2\mu)$ has the limiting behavior $\widetilde{f}(2\mu) \to (2\mu)^{-\beta}L(1/(2\mu))$ as $\mu \to 0$ ($\mu \to \infty$) (Feller (1971), XIII.5, Theorem 4).

Recent times

First we will consider the limit of recent times, $t \to 0$ / $\mu \to \infty$. For pairs sampled within coalescence range, $x \lesssim \delta$, by definition the coalescence time distribution approaches $p(t|x) \sim 1/(\rho\delta^d)$, up to numerical factors that depend on the details of the coalescence kernel. Here d is the dimensionality of the range, $d = 1$ or 2. For pairs sampled well outside coalescence range, $x \gg \delta$, we can assume that $x \gg \overline{x}$ as well, since $\overline{x} = (D_\alpha/\mu)^{1/\alpha} \to 0$ as $\mu \to \infty$. We can also assume that $1 - \psi(0) \to 1$ is independent of μ to leading order. (For $\alpha < d$ our expressions for $\psi(0)$ (38) and (48) are also independent of μ and non-zero, but these are only valid when $\overline{x} \gg \delta$, i.e., when μ is not arbitrarily large.) We can therefore apply the tauberian theorem to (30) and (55) to obtain:

$$p(t \ll x^\alpha/D_\alpha|x) \approx 2d^\alpha \left(\frac{\Gamma(1+\alpha/d)}{\pi}\right)^d \sin\left(\frac{\pi\alpha}{2}\right) \frac{D_\alpha t}{\rho x^{\alpha+d}} \quad (57)$$

$$\approx \frac{1}{\rho}K(x|t).$$

Our heuristic derivation in the Results section essentially proceeded in the opposite direction, starting from $p(t \ll x^\alpha/D_\alpha|x) \approx \frac{1}{\rho}K(x|t)$ and then deriving $\psi(x)$ from that. (57) is thus essentially just a restatement of our expressions for the tail of ψ, and its accuracy can be seen from the same simulation results shown in Fig. 4 and Fig. 6.

Long times

While there is a single unified expression for p in the $t \to 0$ limit, corresponding to the single expression for ψ in the $x \to \infty$ limit, for the opposite limit, $t \to \infty$ / $\mu \to 0$, we must treat different values of α separately, just as we did for ψ at small x. We verify our results with simulations, shown in Fig. 8.

For $\alpha < d$, we can simply take the inverse Laplace and Fourier transforms of (25) to find p, because $1 - \psi(0)$ is independent of μ to leading order. Since we are concerned with times long compared to the time for the lineages to traverse the coalescence zone, $t \gg \delta^\alpha/D_\alpha$, the normal factor in (25) can be neglected and $p(t|x)$ is simply given by the inverse Laplace transform of (24):

$$p(t \gg \delta^\alpha/D_\alpha|x) \approx \frac{1 - \psi(0)}{\rho}K(x|t) \quad (58)$$

$$\approx \frac{1 - \psi(0)}{\alpha(2\pi)^{d-1}\rho} \frac{\Gamma(d/\alpha)}{(2D_\alpha t)^{d/\alpha}} \text{ for } t \gg \frac{x^\alpha}{D_\alpha}. \quad (59)$$

Integrating (59) yields the cumulative distribution for $t \gg$

Figure 8 The distribution of coalescence times has a power law tail. Points show simulation results. Dashed magenta curves show the asymptotic predictions (in order of increasing α) (60), (61), (63), and (65). Time is scaled to dimensionless units. See Simulation Methods section for D_α values. We show statistics based on the cumulative distribution $P(t)$ rather than the density $p(t)$ because simulation estimates for the latter are very noisy. **Top left: for $\alpha < 1$, the distribution of coalescence times is proportional to the probability of lineages being nearby, $K(0|t) \propto t^{1-1/\alpha}$.** Plot shows $P(\infty) - P(t)$ rather than $1 - P(t)$ because lineages can disperse infinitely far away from each other and avoid coalescing entirely, i.e., $P(\infty) < 1$. We use the simulated value of $P(t = 10^6)$ to approximate $P(\infty)$. This empirical value deviates from the continuous-time prediction (7) by $\approx 30\%$ due to differences in the amount of coalescence in the first few generations (see "Breakdown of models at small scales"). **Top right: the distribution of coalescence times has a logarithmic tail for $\alpha = 1$.** In this marginal case, lineages do eventually coalesce even in infinite ranges, but can take extremely long to do so. **Bottom left: for $1 < \alpha < 2$, the distribution of coalescence times decays more quickly than the probability of lineages being nearby.** The coalescence time distribution has a power-law tail, $p(t|x) \propto t^{1/\alpha-2}$. This deviation from the scaling of the dispersal kernel at long times is due to the high probability of previous coalescence events. **Bottom right: for $\alpha > 2$, the coalescence time distribution may approach the diffusive limit.** The scaling of $1 - P$ appears to be close to that of the diffusive prediction, (65), but there is at least a difference in prefactor, perhaps again due to different probabilities of coalescence at very recent times.

x^α/D_α:

$$P\left(t \gg \frac{x^\alpha}{D_\alpha}\,\middle|\, x \gg \delta\right) \approx P(\infty|x) - \frac{\Gamma(d/\alpha)(1-\psi(0))(2D_\alpha)^{-d/\alpha}}{(d-\alpha)(2\pi)^{d-1}\rho t^{d/\alpha-1}}, \tag{60}$$

where $P(\infty|x) = \lim_{\mu\to0}\psi(x)$ is given by (7), (37), (48), or (52), depending on x and d.

For the marginal case $\alpha = d$, we can use a slightly different statement of the tauberian theorem that applies for $\beta = 0$ (Feller (1971), XIII.5, Theorem 2) to convert (38), (40), (49), and (54) to expressions for the cumulative distribution P:

$$P(t \gg x/D_d|x) \approx \begin{cases} \left[1 + \frac{2d\pi\rho D_d}{\ln(2D_d t/\delta)}\right]^{-1} & \text{for } x \ll \delta \\ \frac{\ln(2D_d t/x)}{2d\pi\rho D_d + \ln(2D_d t/\delta)} & \text{for } x \gg \delta. \end{cases} \tag{61}$$

We can then differentiate to find the density p:

$$p\left(t \gg \frac{x}{D_d}\,\middle|\, x\right) \approx \frac{2d\pi\rho D_d}{t\left[2d\pi\rho D_d + \ln(2D_d t/\delta)\right]^2}$$
$$\times \begin{cases} 1 & \text{for } x \ll \delta \\ 1 + \frac{\ln(x/\delta)}{2d\pi\rho D_d} & \text{for } x \gg \delta. \end{cases} \tag{62}$$

Note that in two dimensions, $\alpha = d$ represents the diffusive limit, and we expect these expressions for the marginal case to hold for all $\alpha \geq 2$.

For $d = 1 < \alpha \leq 2$, since the leading behavior of the cumulative distribution is trivial, $\lim_{\mu\to0}\psi(x) = \lim_{t\to\infty}P(t|x) = 1$, we must instead consider the complementary cumulative distribution, $\overline{P}(t|x) \equiv 1 - P(t|x)$. Its Laplace transform is:

$$\widetilde{\overline{P}}(2\mu) = \frac{1}{2\mu} - \widetilde{P}(2\mu|x)$$
$$= \frac{1}{2\mu}\left[1 - \psi(x)\right].$$

We can now apply the tauberian theorem to $\overline{P}$ and $\widetilde{\overline{P}}$. Since we are taking the $\mu \to 0$ limit, we have $\overline{x} \to \infty$, and we need only consider $\psi(x \ll \overline{x})$. Inspecting (33) and (35), we see that they have the limit:

$$\frac{1}{2\mu}\left[P(\infty|x) - \psi(x)\right] \to \alpha\sin(\pi/\alpha)\rho\overline{x} \text{ as } \mu \to 0.$$

Since $\overline{x} = (D_\alpha/\mu)^{1/\alpha}$, $\overline{P}$ has the limit:

$$\overline{P}(t \gg x^\alpha/D_\alpha|x) \approx \alpha\sin\left(\frac{\pi}{\alpha}\right)\frac{\rho(2D_\alpha)^{1/\alpha}}{t^{1-1/\alpha}}. \tag{63}$$

Differentiating (63) yields the density $p(t|x)$:

$$p(t \gg x^\alpha/D_\alpha|x) \approx (\alpha-1)\sin\left(\frac{\pi}{\alpha}\right)\frac{\rho(2D_\alpha)^{1/\alpha}}{t^{2-1/\alpha}}, \tag{64}$$

in agreement with Janakiraman (2017)'s Eq. 19.

For $\alpha = 2$, (63) and (64) simplify to the classic diffusive results:

$$\overline{P}(t \gg x^2/D|x) \approx 2\rho\sqrt{2D/t}, \tag{65}$$

$$p(t \gg x^2/D|x) \approx \rho\sqrt{2D/t^3}. \tag{66}$$

For $\alpha > 2$, we expect the coalescence rate $p(t|x)/\overline{P}(t|x)$ to behave similarly at long times, since the dispersal approaches a diffusion. But the distribution P may be different, due to differences in the probability of early coalescence (Fig. 8, bottom right).

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Breakdown of models at small scales

Great care must be taken in defining coalescent models in continuous space in order to guarantee that they have a consistent forward-time biological interpretation (Felsenstein 1975; Barton *et al.* 2010). We have not done this, and therefore the microscopic behavior of our models does not correspond to any biological population. However, the behavior at large scales (time long compared to one generation, distance long compared to the coalescence scale δ and the typical single-generation dispersal distance $c \equiv (2D_\alpha)^{1/\alpha}$) should still be realistic (Barton *et al.* 2002). Forien (2019) has recently given an example of a biologically consistent forward-time, continuous-space model that we believe would be described by our results on macroscopic scales. We also believe that our results would describe a stepping-stone model of discrete demes of size $\sim \rho\delta^d$ separated by distance $\sim \delta$.

The key place in which the microscopic details matter even for large distances and long times is the factor $1 - \psi(0)$ which appears in many of our expressions. As discussed above, for $\alpha < d$ even here the microscopic details are not necessarily important, but for $\alpha \geq d$ they are. Practically speaking, this quantity would typically have to simply be measured in a population or else treated as a fitting parameter when matching the large-scale predictions to data.

At a microscopic level, we expect that our continuous-time analytic model should deviate from discrete-time models such as the one we use in our simulations. As shown in Fig. 9, this becomes apparent for $\alpha < 1$ in one dimension (or more generally, $\alpha < d$). The two differ at scales smaller than the typical single-generation dispersal distance, $x < c = (2D_\alpha)^{1/\alpha}$, when this scale is large compared to the coalescence scale, $c \gg \delta$. In continuous time, nearby pairs with $x \ll c$ would be able to coalesce at times smaller than a single generation, $t \ll 1$. But in discrete time no pairs can coalesce until $t = 1$, by which time the dispersal kernel $K(x|1)$ is roughly flat out to $x \lesssim c$, and probability of identity thus becomes approximately constant for $x \lesssim c$. (For $\alpha \geq d$, the continuous-time model already predicts that ψ should be changing slowly at $x \ll \overline{x}$, and therefore we do not expect a disagreement with the discrete-time model.) Recall that our discrete-time model assumes no coalescence at $t = 0$ even for lineages starting at $x < \delta$; if we were to change this, ψ would discontinuously jump up to a second, higher plateau for $x < \delta$.

We can estimate the discrete-time value of $\psi(x \ll c)$ from a heuristic argument, at least when $\psi \ll 1$. In the absence of coalescence, the probability of the lineages being within coalescence range of each other in generation $t \geq 1$ is $\approx (2\delta)K(x|t) \approx (2\delta)K(0|t)$. For $\psi \ll 1$, including the possibility of coalescence will only slightly decrease this probability. Given that the lineages are in coalesce range, they coalesce with probability $1/(2\delta\rho)$. So in any one generation the probability of coalescence is $\approx K(0|t)/\rho$ and we can find ψ by summing over all generations:

$$\psi(x \ll c) \approx \sum_{t=1}^{\infty}\frac{K(0|t)}{\rho}$$
$$= \frac{\Gamma(1/\alpha)\zeta(1/\alpha)}{\alpha}\frac{1}{\rho c}, \tag{67}$$

where ζ is the Riemann zeta function. Fig. 9 shows that (67) accurately describes the simulations.

Figure 9 For very long-range dispersal, $\alpha < 1$, continuous-time and discrete-time models differ at short distances.
Scaled probability of identity ψ as a function of distance x for $\alpha = 0.5$, $\delta = 0.5$, and $\rho = 100$. Points show discrete-time simulation results. For the continuous-time model, the black curve shows the result of numerically integrating (27), while the dashed red and magenta lines show the asymptotic approximations (6) and (7), respectively. The continuous-time model predicts that ψ should only plateau within the coalescence distance δ, but for distance between δ and the typical single-generation dispersal distance c, the change in ψ is driven by the probability of coalescing at $0 < t \ll 1$. In the discrete-time model, these lineages have to wait until $t = 1$ to coalesce, leading to a lower, broader plateau, given by (67) (dashed green line). This discrepancy only exists for $\delta < x \ll c$, i.e., if $c < \delta$ then the discrete-time and continuous-time models agree (blue points).

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Appendix: Alternative derivations of the probability of identity by descent

Starting from a recursion equation

(25) can also be derived starting from a recursion equation for ψ requiring that it remain constant over an infinitesimal timestep dt (Malécot 1975; Barton *et al.* 2002):

$$\psi(x) = \frac{dt}{\rho}(1 - \psi(x))\mathcal{N}(x) + e^{-2\mu dt} \int dy \, \psi(x-y) K(y|dt). \tag{68}$$

(68) is saying that at equilibrium the local increase in identity due to coalescence (first term) must be balanced by the loss of identity due to mutation and the spreading of identity due to dispersal (both included in the second term).

Taking the spatial Fourier transform $\mathcal{F}\{\cdot\}$ of (68) simplifies the second term at the expense of complicating the first:

$$\widehat{\psi}(k) = \frac{dt}{\rho}\mathcal{F}\{(1-\psi(x))\mathcal{N}(x)\}(k) + e^{-2\mu dt}\widehat{\psi}(k)\widehat{K}(k|dt). \tag{}$$

Solving for $\widehat{\psi}$ gives:

$$\widehat{\psi}(k) = \frac{\mathcal{F}\{(1-\psi(x))\mathcal{N}(x)\}(k)}{\rho}\frac{dt}{1 - e^{-2\mu dt}\widehat{K}(k|dt)}$$

$$= \frac{\mathcal{F}\{(1-\psi(x))\mathcal{N}(x)\}(k)}{\rho}\sum_{j=0}^{\infty} dt \, e^{-2\mu j dt}\widehat{K}(k|dt)^j, \tag{69}$$

where in the second line we can take the Taylor series expansion because $e^{-2\mu dt} < 1$ and $\widehat{K}(k|dt) \leq 1$ since it is a characteristic function. Assuming dispersal is Markovian, we can simplify (69) by noting that $\widehat{K}(k|dt)^j = \widehat{K}(k|jdt)$, i.e., the distribution after time jdt is just the convolution of j dispersal steps of time dt each. Using this, we can convert the sum into an integral to find

(25):

$$\widehat{\psi}(k) = \frac{\mathcal{F}\{(1-\psi(x))\mathcal{N}(x)\}(k)}{\rho}\widehat{\widetilde{K}}(k,2\mu)$$

$$\approx \frac{1-\psi(0)}{\rho}\widehat{\mathcal{N}}(k)\widehat{\widetilde{K}}(k,2\mu),$$

where in the second line we have used the same ansatz that $1-\psi(x)\approx 1-\psi(0)$ for $|x|\lesssim\delta$ that we used in the main text.

Fractional diffusion equation

For Lévy flight dispersal, (26) can also be derived using a fractional diffusion equation. When X_t follows a diffusion, (17) can be written as a Feynman-Kac (diffusion) equation for ψ (Barton *et al.* 2002; Allman and Weissman 2018). For $\alpha < 2$, this generalizes to a fractional differential equation:

$$0 = 2D_\alpha\left(\frac{\partial^2}{\partial x^2}\right)^{\alpha/2}\psi(x) - 2\mu\psi(x) + \frac{1}{\rho}\mathcal{N}(x)\left(1-\psi(x)\right), \quad (70)$$

where $\left(\frac{\partial^2}{\partial x^2}\right)^{\alpha/2}$ is a Riesz fractional derivative, defined by its Fourier transform $\mathcal{F}\left\{\left(\frac{\partial^2}{\partial x^2}\right)^{\alpha/2}f\right\}(k) = -\,|k|^\alpha\,\mathcal{F}\{f\}(k)$ (Metzler *et al.* 2009; Carmi *et al.* 2010; Janakiraman 2017). It is therefore simpler to consider the Fourier transform of (70), which is equivalent to (26):

$$0 = -(2D_\alpha|k|^\alpha + 2\mu)\widehat{\psi}(k) + \frac{1}{\rho}\mathcal{F}\left\{\mathcal{N}(x)(1-\psi(x))\right\}(k)$$

$$\approx -(2D_\alpha|k|^\alpha + 2\mu)\widehat{\psi}(k) + \frac{1-\psi(0)}{\rho}\widehat{\mathcal{N}}(k). \quad (71)$$

For all $\alpha < 2$, the solution for ψ in two dimensions can be also written as a fractional differential equation (Chen *et al.* 2012):

$$0 = 2D_\alpha\left(\frac{\partial^2}{\partial x^2} + \frac{1}{x}\frac{\partial}{\partial x}\right)^{\frac{\alpha}{2}}\psi(x) - 2\mu\psi(x) + \frac{1}{\rho}\mathcal{N}(x)\left(1-\psi(x)\right), \quad (72)$$

where $\left(\frac{\partial^2}{\partial x^2} + \frac{1}{x}\frac{\partial}{\partial x}\right)^{\frac{\alpha}{2}}$ is a fractional Laplacian, defined by its Fourier transform $\mathcal{F}\left\{\left(\frac{\partial^2}{\partial x^2} + \frac{1}{x}\frac{\partial}{\partial x}\right)^{\alpha/2}\right\}(k) = -\,|k|^\alpha\,\mathcal{F}\{f\}(k)$ (Kwaśnicki 2017; Lischke *et al.* 2020). Note that the rotational symmetry of the problem allows us to write the Laplacian in terms of just the radial coordinate x, and ignore the angular coordinate. The two-dimensional Fourier transform of (72) has exactly the same form as (71), although again the interpretation is different as k is now the radial coordinate in two-dimensional k-space.