

1 Coherence of Terrestrial Vertebrate Species Richness with 2 External Drivers Across Scales and Taxonomic Groups

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Abstract

Aim: Understanding connections between environment and biodiversity is crucial for conservation, identifying causes of ecosystem stress, and predicting population responses to changing environments. Explaining biodiversity requires an understanding of how species richness and environment co-vary across scales. Here, we identify scales and locations at which biodiversity is generated and correlates with environment.

Location: Full latitudinal range per continent.

Time period: Present-day.

Major taxa studied: Terrestrial vertebrates: all mammals, carnivorans, bats, songbirds, hummingbirds, amphibians.

Methods: We describe the use of wavelet power spectra, cross-power and coherence for identifying scale-dependent trends across Earth's surface. Spectra reveal scale- and location-dependent coherence between species richness and topography (E), mean annual precipitation (Pn), temperature (Tm) and annual temperature range (ΔT).

Results: > 97% of species richness of taxa studied is generated at large scales, i.e. wavelengths $\gtrsim 10^3$ km, with 30–69% generated at scales $\gtrsim 10^4$ km. At these scales, richness tends to be highly coherent and anti-correlated with E and ΔT , and positively correlated with Pn and Tm . Coherence between carnivoran richness and ΔT is low across scales, implying insensitivity to seasonal temperature variations. Conversely, amphibian richness is strongly anti-correlated with ΔT at large scales. At scales $\lesssim 10^3$ km, examined taxa, except carnivorans, show highest richness within the tropics. Terrestrial plateaux exhibit high coherence between carnivorans and E at scales $\sim 10^3$ km, consistent with contribution of large-scale tectonic processes to biodiversity. Results are similar across different continents and for global latitudinal averages. Spectral admittance permits derivation of rules-of-thumb relating long-wavelength environmental and species richness trends.

Main conclusions: Sensitivities of mammal, bird and amphibian populations to environment are highly scale-dependent. At large scales, carnivoran richness is largely independent of temperature and precipitation, whereas amphibian richness correlates strongly with precipitation and temperature, and anti-correlated with temperature range. These results pave the way for spectral-based calibration of models that predict biodiversity response to climate change scenarios.

1 Introduction

Biological diversity is critical to many basic human needs, including health, food, water and shelter. It also plays an important role in moderating physical and chemical processes in natural environments

63 (Balmford & Bond, 2005; Barrett *et al.*, 2011; Corenblit *et al.*, 2011; Fei *et al.*, 2014; Mori *et al.*, 2022).
64 Quantifying links between environment and biodiversity is crucial for understanding the response of
65 ecosystems to climatic and physiographic change, and for conservation efforts (Araújo & Rahbek, 2006;
66 Hampe & Petit, 2005; Norris *et al.*, 2013; Yasuhara *et al.*, 2020a). Many extrinsic processes postulated
67 to control biodiversity (e.g. climate) are rapidly changing; therefore quantifying the strength of rela-
68 tionships between them is a pressing concern (Nogués-Bravo *et al.*, 2018).

69

70 Environmental variables and species richness exhibit variance in space across a range of scales (e.g.
71 Belmaker & Jetz, 2011; Buckley *et al.*, 2012; Keil & Chase, 2019). However, it is unclear whether
72 coherence between variables is uniform across all scales Storch *et al.* (2007). In this study we test
73 the following five hypotheses. First, species richness is highly coherent with environmental variables
74 across all scales. That would imply a direct forcing of richness by external drivers regardless of scale.
75 It would give a basis for using theory developed to predict species richness at one scale (e.g. field sites)
76 to predict richness at all scales. Secondly, species richness is most coherent with external variability at
77 small scales, i.e. local changes in environment determine where species richness prospers. Thirdly, spe-
78 cies richness is most coherent with changes in environment at large scales, i.e. global scale variability
79 (e.g. large-scale climate change). Fourthly, the coherence of species richness with external variables
80 depends on taxon. In other words, taxa have unique responses to environment variables. Finally,
81 species richness does not directly depend on environment (i.e. coherence between species richness and
82 environmental variables is low). Instead, species richness depends upon other factors, namely biotic
83 interactions (prey-predator, competition), and/or historical contingencies.

84

85 Here, we test the five hypotheses by quantifying coherence between species richness of continental
86 vertebrate taxa and elevation, precipitation, temperature, and annual temperature range, which are
87 postulated to drive biodiversity (e.g. Antonelli *et al.*, 2018; Rahbek & Graves, 2001). As a starting
88 point, we focus on mapping coherence between contemporary biotic and environmental signals as a
89 function of scale and location, using wavelet spectral analyses. Many existing approaches, e.g. spa-
90 tial regression analyses, are unsuited to testing such hypotheses because of the challenges associated
91 with disentangling scale and location from biotic and environmental data to identify correlations. In-
92 stead, here, we use wavelet spectral analyses, which inherently disentangle scale-dependent effects, and
93 identify strength of correlation between variables at individual scales. Such analyses have been used to
94 identify scale-dependence of temporal biotic series, to filter spatial series and identify outliers, and to
95 investigate biodiversity on local ($\lesssim 500$ km) scales (Carl *et al.*, 2008, 2016; Dormann *et al.*, 2007; Keitt,
96 2007; Ma & Zhang, 2015; Roberts & Mannion, 2019). We acknowledge that other processes, including
97 species-species interactions, are also important for determining species richness (e.g. Chaudhary *et al.*,
98 2021; Yasuhara *et al.*, 2020b; Yasuhara & Deutsch, 2022). As such, we also present a preliminary

99 assessment of the coherence between species richness of different taxa, in Supporting Information. We
100 return to discuss the results of these tests in the context of the five hypotheses described above, in the
101 concluding section of this manuscript.

102

103 Identifying links between biodiversity and environment has recently become significantly more
104 tractable for three reasons. First, global patterns of species richness have been estimated with un-
105 precedented detail, from horizontal scales as broad as continents, to those as fine as ~ 10 km in
106 wavelength (e.g. Jenkins & Joppa, 2009; Jenkins *et al.*, 2013, 2020; Kass *et al.*, 2022; Marsh *et al.*,
107 2022). Second, values and variance of many environmental variables postulated to be responsible for
108 determining distributions of species are now available globally at even higher resolution (e.g. Kar-
109 ger *et al.*, 2017). Finally, wavelet spectral methods, which can identify the locations and scales at
110 which signals (e.g. spatial series of taxa) are generated, as well as coherence and phase differences
111 (offsets) between series such as species richness, topography and climate, are now established (see
112 Materials and Methods; Grinstead *et al.*, 2004; Torrence & Compo, 1998). Such analyses are key to
113 understanding how the changing global climate will affect the distribution of biodiversity across Earth.

114

115 2 Materials and Methods

116 2.1 Species Richness Data

117 Species richness is here defined as number of species of a given taxon within a 10×10 km square. We
118 use the grids compiled by Jenkins *et al.* (2013), which were generated by combining maps of species
119 distributions, and counting the number of overlapping polygons in a given cell. For birds, the species
120 richness data were calculated from breeding ranges compiled by BirdLife International NatureServe
121 (2011). For amphibians and mammals, the data were based on expert range maps generated by the
122 International Union for Conservation of Nature (2021).

123

124 Figure 1a–f shows species richness per 10×10 km cell for all mammals (Mammalia), carnivorans
125 (Carnivora), bats (Chiroptera), songbirds (Passeriformes), hummingbirds (Trochilidae), and amphi-
126 bians (Amphibia). In Supporting Information Figures 1–2, we show species richness and associated
127 analyses for other evaluated taxonomic groups (including Cetartiodactyla, Eulipotyphla, primates,
128 marsupials, rodents, parrots, and frogs). The taxa which we focus on in the main text of this study
129 have the greatest latitudinal coverage and are well-mapped, and they also cover a range of modes of life.
130 The data shown in Figure 1 reinforce well-known large-scale observations, e.g. the latitudinal diversity
131 gradient, but also contain evidence of significant complexity across scales of interest, here wavelengths

132 between 10–10⁴ km (e.g. Hillebrand, 2004; Willig *et al.*, 2003). We examine species richness trends in
133 this study, since it is a straightforward measure of diversity, and has been determined for a wide range
134 of taxa from fine scales up to near global scales. Here, we focus on terrestrial vertebrate taxa since
135 terrestrial surface environmental conditions are well-mapped, as is terrestrial vertebrate biodiversity.
136 Similar analysis is possible for marine taxa, invertebrates, plants etc., and for metrics other than spe-
137 cies richness, for example range sizes and trophic interactions.

138

139 Species richness varies as a function of the spatial range characteristics of a study, particularly
140 “grain”, i.e. piece-wise horizontal resolution within a study (Gaston, 2000; Palmer & White, 1994;
141 Willig *et al.*, 2003). By using a constant grain (i.e. “focus” or grid spacing) of 10 km, challenges
142 associated with comparing results generated using different grains are avoided (Willig *et al.*, 2003).
143 Here, scale-dependent trends are calculated as a function of “extent”, i.e. total width of study region,
144 rather than “grain”, i.e. width of each plot/grid cell within the study region *sensu* Palmer & White
145 (1994). Hurlbert & Jetz (2007) indicated that range map data might only be valid at wavelengths
146 > 100 km. In this study, we evaluate how short-wavelength uncertainties in species richness contribute
147 to uncertainties in calculated wavelet spectra by adding theoretical noise to transects before they are
148 transformed into the spectral domain (Supporting Information Figure 3; panels a–c show results of
149 adding white noise which has wavelengths between 10–100 km). Latitudinal transects through ter-
150 restrial vertebrate richness and environmental data are shown in Figure 2. We show data from the
151 Americas, where transects can be generated that encompass almost all of Earth’s latitudinal range
152 (Figures 1 & 2: A–A’). Transects through data for Australia (B–B’), Africa (C–C’), Eurasia (D–
153 D’) and global averages are shown in Supporting Information Figures 4–7.

154

155 2.2 Environmental Variable Data

156 Figures 1g–j and 2m, o, q and s show examples of maps and cross sections through elevation and
157 climatic data which we use, from the ETOPO1 and CHELSA datasets, respectively (Amante & Eakins,
158 2009; Karger *et al.*, 2017).

159

160 The global elevation grid ETOPO1 has a horizontal resolution of 1 arc-minute (Figure 1g; Amante
161 & Eakins, 2009). It is primarily generated from ~ 30 m resolution Shuttle Radar Topography Mis-
162 sion (SRTM30) data and includes interpolated coastlines and satellite altimetry (Jarvis *et al.*, 2008).
163 Amante & Eakins (2009) suggested a mean vertical error of ~ 10 metres for ETOPO1. We down-
164 sampled the data to a horizontal resolution of 10 km using Generic Mapping Tools to match resolution
165 of species richness grids (Wessel *et al.*, 2019).

166

167 Annual mean values for climatic data, from 1981–2010, were extracted from the Climatologies at
168 High Resolution for the Earth’s Land Surface Areas (CHELSA) dataset (Karger *et al.*, 2017). CHELSA
169 was generated by applying corrections to the ERA-Interim climatic reanalysis and has a horizontal
170 resolution of up to 30 arc-seconds (Dee *et al.*, 2011). Temperature data were corrected for elevation
171 above sea level and precipitation rates were corrected using wind direction, valley exposition and
172 boundary layers. Precipitation rate is weakly dependent on elevation. These values were successfully
173 benchmarked against alternative climatology data and models: WorldClim, TRMM, GPCP and GHCN
174 (Hijmans *et al.*, 2005; Goddard Earth Sciences Data and Information Services Center, 2017; Lawrimore
175 *et al.*, 2011; Schneider *et al.*, 2014). The data were down-sampled to 10 km prior to spectral analyses.

176 2.3 Continuous Wavelet Transform

177 Spatial series, x_n , of species richness or environmental variables were transformed into distance-
178 wavenumber space using continuous wavelet transforms (for practical guide, see Torrence & Compo,
179 1998). The transform convolves uniformly sampled spatial series with a mother wavelet, ψ . The
180 Morlet wavelet with dimensionless frequency $\omega_o = 6$ is used in this study, although other mother
181 wavelets are investigated in Supporting Information Figure 8. Use of different mother wavelets (Mor-
182 let, order $\omega_o = 4, 8$; Paul, order $m = 2, 4, 6$; derivative of Gaussian, order $m = 2, 4, 6$) does not
183 significantly change patterns of mapped power, and distance-averaged power shows similar trends to
184 the results presented here. The mother wavelet is scaled and translated along spatial series by n' to
185 reveal variations in amplitude as a function of scale, s , and position, x_n . Sampling interval $\delta t = 10$
186 km, $n = 0, 1 \dots N - 1$, where N is number of measurements. The wavelet transformation is

$$W_n(s) = \sum_{n'=0}^{N-1} x_n \psi^* \left[\frac{(n' - n)\delta t}{s} \right], \quad (1)$$

187 where $*$ denotes the complex conjugate. We use the `mlpy` Python module to transform the spatial
188 series (Albanese *et al.*, 2012), which is based on the methods summarized by Torrence & Compo
189 (1998). Scales were calculated using the approach described in Torrence & Compo (1998), such that
190 $s_j = s_o 2^{j\delta_j}$, where $j = 0, 1, \dots J$. The smallest scale, $s_o = 2\delta t$. A minimum grid spacing of 10 km
191 therefore yields a minimum scale for wavelet spectral analysis of ~ 20 km (Torrence & Compo, 1998).
192 In the example shown in Figure 2, $N = 1598$, $\delta_j = 0.1$ and $J = 96$, which yields a total of 97 scales
193 that range from ~ 20 to $\sim 15,521$ km. Spatial series were mirrored across the x (distance) and
194 y (dependent variable) axes to reduce edge effects (Roberts *et al.*, 2019). Inverse transforms were
195 reconstructed for each signal to quantify fidelity of transformed series (see Torrence & Compo, 1998).
196 Median difference between input signals and inverse transforms were always $\leq 0.9\%$. Signals can be
197 filtered by calculating inverse transforms at specific wavelengths. For example, Figure 2a, c, e, g, i and

198 k show inverse transforms of species richness spectra at scales $> 10^3$ km. Depending on taxonomic
199 group, these filtered signals fit the input species richness trends with mean differences of 4.4–25% (see
200 Figure 2 caption for mean raw differences in terms of species per pixel, spx). The same filtering process,
201 but including wavelengths $> 10^2$ km, yields mean differences of only 0.7–4.7%. The distance-averaged
202 power spectrum, which yields similar results to Fourier transformation, is given by

$$\phi(s) = \frac{1}{N} \sum_{x=0}^N |W_n(s)|^2. \quad (2)$$

203 In Figure 2, we plot rectified distance-averaged power $\phi_r = \phi(s)s^{-1}$ after Liu *et al.* (2007). We
204 calculate distance-averaged power within and outside of the tropics, but note that in those calcula-
205 tions, power was normalized by the proportion of the transect within/outside of the tropics respect-
206 ively. Therefore there is no bias in distance-averaged power if the transect has a greater distance
207 within/outside of tropical latitudes. A guide to scale-dependence and self-similarity of spatial series is
208 the color of spectral noise that they possess. For example, red (Brownian) noise occurs when $\phi \propto k^{-2}$,
209 where k is wavenumber or spatial frequency, proportional to $1/\text{wavelength}$, indicating self-similarity.
210 Pink noise occurs when $\phi \propto k^{-1}$, and white noise indicates that power is equal across all scales,
211 $\phi \propto 1$. Best-fitting spectral slopes for all variables and transects were identified using simple one- and
212 two-slope models after Roberts *et al.* (2019); see Supporting Information Figures 9–18.

213 2.4 Cross Wavelet Power & Wavelet Coherence

214 Cross wavelet power is calculated to identify signals in separate spatial series (e.g. amphibian richness
215 and precipitation) that have large amplitudes located at the same position in distance-wavenumber
216 space. To facilitate comparison, signals are normalized to zero mean and unit variance prior to trans-
217 formation. The normalized signals X and Y , are transformed to yield W^X and W^Y . Cross wavelet
218 power W^{XY} is calculated such that

$$W^{XY} = W^X W^{Y*}, \quad (3)$$

219 where $*$ denotes complex conjugation. Wavelet coherence, R_n^2 , is calculated to identify parts of signals
220 that are coherent, but not necessarily of common high amplitude, such that

$$R_n^2(s) = \frac{|S\{s^{-1}W_n^{XY}(s)\}|^2}{S\{s^{-1}|W_n^X(s)|^2\} \cdot S\{s^{-1}|W_n^Y(s)|^2\}}, \quad (4)$$

221 where s, n and $W_n(s)$ are as in Equation 1. S is an operator that smooths along distance and scale
222 (Grinsted *et al.*, 2004).

223

224 Signals with certain spectral distributions (e.g. red noise) can, by chance, correlate without true

225 interdependence. Therefore, it is important to calculate the coherence between each pair of signals,
226 and not simply their cross wavelet power. Torrence & Compo (1998), Grinsted *et al.* (2004) and others
227 have shown that coherence between signals above (assumed) background noise (i.e. spectral distribu-
228 tions) can be estimated by first calculating the coherence between large numbers of surrogate datasets
229 with the same autoregressive (AR) coefficients as the original data set. In this study, the minimum
230 bound for statistically significant coherence (above assumed background noise) per scale, for each tran-
231 sect, was calculated from cross wavelet power spectral analysis of 300 random signals. Guided by the
232 spectral content of actual biotic and environmental signals, we assumed that each theoretical random
233 signal has a red noise distribution, which was generated using the same autocorrelation coefficient as
234 the actual input signals. The surrogates also have the same length and number of measurements,
235 N , as the actual signals. The 90% significance limit for coherence, which was used to mask Figure
236 3, depends only on scale and not position, and was calculated using Monte Carlo methods with the
237 PyCWT Python module (Grinsted *et al.*, 2004; Krieger *et al.*, 2020).

238

239 The local phase difference (angular offset, $0 \leq a \leq 2\pi$) of two signals is given by the complex
240 argument of their cross wavelet transform, $\arg(W^{XY})$ (Grinsted *et al.*, 2004). Figure 3 indicates
241 phase difference as arrows measured from horizontal: in-phase, $a = 0$, \triangleright ; anti-phase, $a = \pi$, \triangleleft .
242 A working example for species richness and elevation, including continuous wavelet transformation,
243 cross wavelet power and wavelet coherence calculations, can be found at [https://doi.org/10.5281/](https://doi.org/10.5281/zenodo.XXXXXX)
244 [zenodo.XXXXXX](https://doi.org/10.5281/zenodo.XXXXXX).

245 3 Results and Discussion

246 3.1 Wavelet Transformation of Richness and Environmental Variables

247 Spectral analyses of vertebrate species richness and environmental variables are shown in Figures 2 and
248 3. Figure 2 shows that highest spectral power, $\phi (\propto z^2$, where z is signal amplitude), is concentrated
249 at largest scales for all taxa and environmental variables studied. Dependent on taxonomic group,
250 from 96% to almost 100% of power resides at wavelengths $> 10^3$ km. 30–69% of power resides at
251 wavelengths $\gtrsim 10^4$ km. These results reinforce the notion that species richness is dominated by long
252 wavelength, latitudinal, variability.

253

254 Species richness tends to have a pink noise spectrum (see Supporting Information Figures 9–18).
255 Thereby, shorter wavelength features in species richness signals tend to have the lowest amplitudes
256 and comprise relatively little (few %) of species richness signal at a particular location. Mammals
257 and bats are better characterized by red noise at long wavelengths. This result implies self-similarity

258 across scales, and that signal amplitudes decrease even more rapidly with decreasing wavelength than
259 for other taxa. At wavelengths $\gtrsim 10^3$ km, species richness power for amphibians is best characterized
260 as blue noise, i.e. $\phi \propto k^1$. This trend is not observed along the entire transect, but indicates that short
261 wavelength features can be increasingly important contributors to amphibian richness (see Figure 2f).
262 A single spectral slope akin to pink noise can adequately fit the amphibian richness spectrum (see e.g.
263 Supporting Information Figure 9f).

264

265 We address concerns regarding the accuracy of range maps in two ways. First, we show the results
266 of inverse wavelet transforms at scales > 1000 km alongside the results for transformation of the full
267 frequency content of the available data. Secondly, we explore the impact of inserting distributions
268 of theoretical, hitherto unknown, species in a suite of increasing severe tests for our results and con-
269 clusions. White noise was added to the amphibian transect in a systematic set of tests. These tests
270 examined changes in calculated spectra when noise with maximum amplitudes of 10%, 50% and 100%
271 of the standard deviation of the original signal's amplitude (in this case = 24 species per pixel) was
272 added to the transect prior to transformation. These tests included adding noise at wavelengths $\lesssim 100$,
273 $\lesssim 1000$ and $\lesssim 10,000$ km (Supporting Information Figure 3). As expected, these tests indicated that
274 spectral power is least likely to be well constrained at short wavelengths. Nonetheless, these tests
275 indicate that even high amplitude uniformly distributed noise does not significantly change the over-
276 all spectral characteristics of terrestrial species richness. Finally, we note that that Hurlbert & Jetz
277 (2007) suggest that species richness values estimated from range maps are likely to be overestimates at
278 short length scales compared to richness estimated from atlas data (e.g. their Figure 5). That result
279 indicates that range maps are more likely to generate higher power at short wavelengths than maps
280 derived from atlas data. Therefore, spectra derived from atlas data are likely to be even redder than
281 those obtained from range maps. In other words, our conclusion that species richness is dominated
282 by long wavelength structure is expected to be insensitive to the choice of range or atlas data. In
283 fact, atlas data is likely to more strongly emphasize the importance of long wavelength variability for
284 determining species richness. We suggest that these observations and results indicate that range maps
285 are useful for our purposes.

286

287 Although almost no power is concentrated below wavelengths of ~ 100 km for any of the taxa
288 examined here, there are regions of some wavelet transforms which show increased power in the range
289 ~ 300 – 1000 km. This deviation, away from a broadly monotonic decrease in power towards shorter
290 wavelengths, is driven principally by species richness within tropical latitudes, and is especially prom-
291 inent for songbirds, hummingbirds and amphibians (Figure 2h, j, l). Supporting Information Figure
292 19a–f shows that at wavelengths $\gtrsim 1000$ km, there is no notable difference between power in species
293 richness within or outside the tropics, across the Americas. However, at wavelengths $\lesssim 1000$ km, there

294 is significantly greater power for regions within the tropics. This trend arises because power spectral
295 slopes remain close to -2 at shorter wavelengths outside of the tropics (i.e. red noise; Supporting
296 Information Figure 19), before increasing to be closer to -1 (i.e. pink noise). We suggest that these
297 results are consistent with the concept that topography in tropical regions can generate higher spe-
298 cies richness towards the equator via the increased effectiveness of altitudinal variation in habitat at
299 isolating species either physically or physiologically (i.e. by being associated with variation in other
300 environmental variables such as temperature; Ghalambor *et al.*, 2006; Janzen, 1967). We find the
301 effect has a greatest impact on species richness power of hummingbirds and amphibians; the impact
302 on bats and songbirds richness appears to be more modest. Tropical increases in species richness of
303 carnivorans, and mammals more generally, are much more subdued (Supporting Information Figure
304 19).

305

306 Elevational transects exhibit red and pink noise spectral characteristics at wavelengths $\gtrsim 10^3$ km
307 and $\lesssim 10^3$ km, respectively, which we note is similar to distance-averaged power from wavelet trans-
308 forms of longitudinal river profiles and other topographic transects (Supporting Information Figures 9g,
309 9q, 10g, 10q; Roberts *et al.*, 2019; Wapenhans *et al.*, 2021). Precipitation rate, temperature and annual
310 temperature range can also be characterized as red and pink noise (Supporting Information Figures
311 9h–j, r–t & 10h–j, r–t). Similar results are obtained for latitudinal transects through Africa, Eurasia
312 and Australia, as well as across global, latitudinally-averaged sections (see Supporting Information
313 Figures 11–23).

314 **3.2 Coherence between Richness and Environment**

315 Visual inspection of Figure 2 indicates that there is strong, location- and scale-dependent, similarity
316 between the wavelet transforms of transects through species richness and environmental variables. To
317 quantify the strength of these relationships we calculate cross wavelet power, which identifies co-located
318 high amplitudes in the location-scale domain, and wavelet coherence (see Materials and Methods). Fig-
319 ure 3 shows results for carnivorans (which are similar to those for mammals generally), and amphibians
320 (which are similar to those for bats, songbirds and hummingbirds). See Supporting Information Figure
321 1–2 for analyses of those other taxa. Figure 3a shows cross wavelet power between species richness
322 of carnivorans along transect A—A' and elevation. Almost no short-wavelength ($< 10^3$ km) features
323 are coherent above a 90% confidence limit. These short wavelength regions contain almost no cross
324 wavelet power; 94% of all cross wavelet power is in the region of high coherence colored on Figure 3a,
325 which accounts for 30% of the location-scale domain. 79% of the area of the cross wavelet spectrum
326 that is significantly coherent resides at wavelengths $\gtrsim 10^3$ km. Distance-averaged cross wavelet power
327 for all parts of the power spectrum, not just those parts which are coherent above the 90% significance

328 threshold, is shown to the right of each panel, on a logarithmic scale. Full, unmasked, plots of cross
329 wavelet power are shown in Supporting Information Figure 24. Masked and unmasked cross power
330 plots for other transects and global latitudinal averages are shown in Supporting Information Figures
331 25–32. Distance-averaged cross wavelet power between all taxa and environmental variables studied is
332 shown in Supporting Information Figures 19–23, panels g–ad.

333

334 Cross wavelet power between amphibians and elevation is also highest at long wavelengths, al-
335 though overall there is a smaller proportion of the two signals that is coherent: 78% of the plot region
336 is masked in Figure 3e (gray regions). Only a small part of the cross wavelet transform for amphibians
337 and elevation is coherent below wavelengths of ~ 5000 km, and that part lies near the centre of the
338 transect, i.e. within the tropics. Distance-averaged power outside the tropics, plotted to the right
339 of Figure 3e as a gray curve, is an order of magnitude lower than within the tropics, especially at
340 wavelengths $\lesssim 3000$ km. This observation is in contrast to cross power between species richness of
341 carnivorans and elevation, where there is almost no difference between the results within and outside
342 of the tropics, across all scales. These results may indicate that carnivorans are less affected by “moun-
343 tain passes” (*sensu* Janzen, 1967) in the tropics, compared with amphibians (cf. Antonelli *et al.*, 2018;
344 Eronen *et al.*, 2015; Rahbek *et al.*, 2019; Rolland *et al.*, 2015). At wavelengths $\sim 10^3$ km, carnivoran
345 species richness is most coherent with elevation and mean annual temperature atop terrestrial plateaux
346 (e.g. Rocky-Mountains-Colorado Plateau and Altiplano, between 4000 – 7000 km and 13,000 – 14,000
347 km distance along transect A–A’, respectively; Figures 1–3). An obvious interpretation is the local
348 importance of tectonics for determining biodiversity (Antonelli *et al.*, 2018).

349

350 Statistically significant (above background red noise), coherent cross wavelet power between carni-
351 voran species richness, mean annual precipitation rate, temperature and annual temperature range is
352 shown in Figure 3b–d. Results for amphibians are shown in panels f–h. Cross power between amphi-
353 bian species richness and precipitation rate, temperature, and temperature range tends to be higher
354 within the tropics compared to outside the tropics at wavelengths $\lesssim 3000$ km (cf. grey, blue, black
355 lines in Figure 3). Those differences are absent or reduced for carnivorans. Furthermore, a smaller area
356 of the power spectra of these three climatic variables is significantly coherent with carnivoran richness,
357 compared to amphibian richness (cf. extent of gray masks in Figure 3). One likely interpretation of
358 these results is that carnivorans are less sensitive to changes in those variables than amphibians (e.g.
359 Rolland *et al.*, 2018). Calculated phase indicates long-wavelength anticorrelation between elevation
360 and species richness for both carnivorans and amphibians (left-pointing arrows in Figure 3a and e;
361 phase angle, $a = \pi$; see Materials and Methods). Highly coherent long-wavelength anticorrelation
362 between amphibian species richness and annual temperature range is also observed across the entire
363 transect. Highly coherent, long-wavelength cross power between precipitation rate or temperature,

364 and species richness of both carnivorans and amphibians, is in phase, i.e. there is positive correlation
365 at these scales. This result is in agreement with the idea that faster diversification rates contribute
366 to heightened species richness, since it suggests that both taxonomic groups benefit from increased
367 energy and higher productivity associated with greater availability of heat and water (cf. Allen *et al.*,
368 2006).

369

370 **3.3 Global and Local Species Richness and Environment**

371 These results for the Americas can be compared to transects from Australia, Eurasia and Africa. For
372 Australia, similar trends in power spectral slopes, distance-averaged power and cross wavelet power
373 are observed (Figure 1: B—B'; Supporting Information Figures 4, 11–12, 20, 25–26). However, for
374 Australia there is almost no difference in power or cross power between tropical regions and regions
375 outside the tropics. We note, however, that the Australian transect does not include the entirety of
376 the tropics; it only spans latitudes between 11.2°S–37.6°S. Signals are mostly coherent at wavelengths
377 $\gtrsim 10^3$ km, and the same pattern of correlation/anticorrelation is observed with climatic variables as
378 that recovered for the Americas (Supporting Information Figures 25–26). In Africa, songbirds and
379 amphibians have greater species richness power within the tropics but the differences outside of the
380 tropics are not as stark as for the Americas (Figure 1: C—C'; Supporting Information Figure 21a–f).
381 This result may reflect differences in Cenozoic paleoclimatic history between Africa and the Americas
382 (Hagen *et al.*, 2021). The greatest difference between cross power within and outside the tropics is for
383 precipitation rate, suggesting that water availability is a more important control species richness for all
384 African taxa studied here. Wavelet coherence indicates that, across Africa, carnivoran species richness
385 does not correlate with environmental variables, whereas species richness of amphibians is strongly
386 positively correlated with precipitation rate at long wavelengths, in agreement with the findings of
387 Buckley & Jetz (2007). Anticorrelation is observed between amphibian species richness and temper-
388 ature across Africa. Results for Eurasia are dominated by the presence of the Tibetan Plateau, and
389 the low proportion of the transect within tropical latitudes (Figure 1: D—D'; Supporting Information
390 Figures 6, 15–16, 22, 29–30). Similar trends to the Americas are observed, albeit with generally lower
391 cross power and coherence.

392

393 Mean terrestrial values of each variable across all latitudes globally were transformed into the
394 location-scale domain. Distance-averaged wavelet power spectra of the resulting transects have spec-
395 tral slopes between -2 and -1 (red to pink noise), reflecting the importance of long-wavelength trends.
396 Species richness power for all taxa, except Mammalia and Carnivora, is at least an order of magnitude
397 lower outside of tropical latitudes, at wavelengths $\lesssim 3000$ km, consistent with results obtained from

398 transforming the American transect (Figures 2 and 3). This result suggests that the increase in species
399 richness power at short wavelengths may be a global phenomenon, reflecting sensitivity of tropical
400 species to local climatic effects.

401

402 Figure 4a–f shows a summary of results for the transect across the Americas, A—A'. Panels a and
403 b show inverse wavelet transforms generated using only the the longest 20% of scales (wavelengths
404 $\gtrsim 3750$ km; approximately one quarter of the length of the transect), for amphibian and carnivoran
405 species richness respectively. Those taxa have significantly different modes of life and their wavelet
406 power, cross power and coherence with environment exhibit the greatest differences of any taxa stud-
407 ied. These low-pass filtered series account for ~ 41 and $\sim 11\%$ of species richness in terms of mean
408 difference to input, respectively. Figure 4c–f shows inverse wavelet transforms of environmental series
409 filtered in the same way. Coherence, R_n^2 , between the filtered series and amphibian (green) and carni-
410 voran (purple) species richness trends is annotated on each panel. These results for the Americas are
411 consistent with global averages, although coherence is generally lower for global results since regions
412 of common power may not be generated at the same longitudes (see Figure 4g–l).

413

414 3.4 Transfer Functions Between Environment and Species Richness

415 It is useful to derive ‘rules-of-thumb’ to estimate species richness from environmental variables. This
416 study indicates that such conversion schemes (transfer functions) are likely to be particularly useful
417 at large scales where most species richness appears to be determined, and where coherence with
418 environmental variables is highest. In the scale-distance domain, after wavelet transformation (see
419 Methods and Materials), species richness W of any given taxon X can be expressed:

$$W^X(x, s) = Z^{X\epsilon}(x, s) \cdot \epsilon(x, s) + \eta, \quad (5)$$

420 where x is distance, s is scale, and η is noise or contributions from variables that have not been
421 considered. $Z^{X\epsilon}$ refers to the admittance (transfer function) between a set of environmental variables ϵ ,
422 and richness in the scale-distance domain, W^X . For any individual variable Y within ϵ , the admittance
423 between W^X and W^Y can be expressed:

$$Z^{XY}(x, s) = \frac{W^X W^{Y*}}{W^Y W^{Y*}}, \quad (6)$$

424 where $*$ denotes complex conjugate. Thereby, a species richness signal X can be estimated from
425 calibrated admittance Z^{XY} , by convolving the inverse transform of Z^{XY} with Y (see Methods and
426 Materials; Torrence & Compo, 1998).

427

428 Table 1 and Figure 4 show estimates of admittance between environmental variables and amphibian
 429 and carnivoran richness series for the largest 25% of scales (wavelengths ≥ 3756 km). The mean of
 430 the transfer function between mean annual precipitation (i.e. W^Y) and amphibian richness (W^X) in
 431 the Americas, for example, is $12.7_{-3.2}^{+5.3}$ spx/m. These scales account for $\gtrsim 60\%$ of observed amphibian
 432 species richness. Stated uncertainties are distances to 1st and 3rd quartiles of admittance values, across
 433 all (x, s) space for $s \geq 3756$ km. Calculated admittance is most likely to be a reliable rule of thumb
 434 for converting environmental variable values into species richness when coherence is high. In Table
 435 1, bold values indicate values where coherence is > 0.5 . For example, consider the mean large-scale
 436 coherence between amphibian richness and precipitation, which is 0.76 ± 0.2 . This result, coupled
 437 with the associated high positive admittance value, is suggestive of amphibian richness dependence on
 438 large-scale precipitation patterns. Compare that result to the relatively low mean large-scale coher-
 439 ence between American carnivoran richness and annual temperature range, which is only 0.22 ± 0.2 ,
 440 reflecting their likely independence. A comparison with results generated using global mean species
 441 richness and environmental variables indicates that these rules are generally applicable (see Table 1
 442 & Figure 4). Such simple rules-of-thumb appear to provide a means to predict species richness from
 443 external variables at large ($\gtrsim 1000$ km) scales.

444

Table 1: Rules of thumb to convert values of environmental variables into species richness. E = elevation, Pn = mean annual precipitation, Tm = mean annual temperature, ΔT = annual temperature range (see Figures 1–2). Z = admittance (transfer function) between environmental variables and species richness at large scales (> 3756 km); subscripts AA & MA indicate American amphibians & mammals; subscripts AG & MG indicate global amphibians & mammals; units are spx/km, spx/m or spx/ $^{\circ}$ C, where spx = species per pixel. R_n^2 = coherence between environmental variable and species; bold = admittance and coherence values for species-variable relationships with mean $R_n^2 > 0.5$

Variable	Z_{AA}	R_n^2	Z_{MA}	R_n^2	Z_{AG}	R_n^2	Z_{MG}	R_n^2
E , km	$-7.6_{-2.5}^{+3.0}$	0.52 ± 0.3	$-1.7_{-1.1}^{+0.6}$	0.62 ± 0.2	$2.5_{-17.7}^{+15.5}$	0.14 ± 0.1	$-1.3_{-6.4}^{+5.5}$	0.13 ± 0.1
Pn , m	$12.7_{-3.2}^{+5.3}$	0.76 ± 0.2	$2.5_{-0.9}^{+2.0}$	0.51 ± 0.2	$14.0_{-7.5}^{+10.1}$	0.41 ± 0.2	$1.8_{-2.5}^{+2.2}$	0.21 ± 0.2
Tm , $^{\circ}$ C	$0.9_{-1.0}^{+0.8}$	0.31 ± 0.2	$0.2_{-0.3}^{+0.2}$	0.28 ± 0.2	$-0.4_{-2.3}^{+3.0}$	0.46 ± 0.2	$-0.3_{-0.5}^{+0.7}$	0.55 ± 0.3
ΔT , $^{\circ}$ C	$-0.4_{-0.5}^{+0.9}$	0.26 ± 0.3	$-0.1_{-0.1}^{+0.3}$	0.22 ± 0.2	$-0.7_{-0.7}^{+0.7}$	0.28 ± 0.2	$0.0_{-0.2}^{+0.1}$	0.29 ± 0.2

445 3.5 Drivers of Species Richness

446 In the introduction of this paper we described five hypotheses to test. First, we hypothesized that
 447 species richness is highly coherent with environmental variables across all scales. Given the spectral
 448 analyses we present, which indicate that most species richness does not have statistically significant co-
 449 herence with environmental variables at wavelengths $\lesssim 1000$ km at most latitudes, this hypothesis can
 450 be rejected. Our second hypothesis—species richness is most coherent with external variability at small

451 scales—is also thus rejected. Our third hypothesis—species richness is most coherent with changes in
452 environment at large scales—was found to be reasonable. Generally, our results indicate that species
453 richness is most coherent with environment at wavelengths $\gtrsim 10^3$ km, where highest species richness
454 power also resides. Our fourth hypothesis—that coherence of species richness with external variables
455 depends on taxonomic group—was also found to be reasonable. For example, amphibian richness is
456 found to be highly coherent with temperature range at scales > 5000 km, whereas carnivoran richness
457 has very low coherence at these scales (cf. Figure 3d and 3h). Furthermore, amphibian richness seems
458 to be more coherent with precipitation and temperature within the tropics, while carnivoran richness
459 is not. Our fifth and final hypothesis—that species richness does not directly depend on environment,
460 instead, species richness depends upon biotic interactions—requires modification. Species richness was
461 found to have both high and low coherence with environmental variables, depending on location, scale
462 and the environmental variable being considered. We have shown that wavelet transformation provides
463 a means to identify coherence in the space-frequency domains. Our preliminary assessment of species-
464 species interactions indicates that their coherence is also scale- and location-dependent (Supporting
465 Information Figure 2). However, we note that calculated coherence and cross power between species
466 tends to be lower than that between species and environmental variables. We note that historical
467 effects, i.e. speciation/extinction rates over geologic time, are not identified within this study, solely
468 modern correlations between variables, although long-term speciation/extinction rates may themselves
469 depend on environment (e.g. Skeels *et al.*, 2022). Nonetheless, we tentatively suggest that these res-
470 ults indicate that environment is more important in determining species richness than species-species
471 interactions.

472

473 **3.6 Implications for Macroecological Biodiversity Patterns**

474 A principal result of this study is that terrestrial species richness tends to be most coherent with
475 topography, precipitation and temperature at long wavelengths ($> 10^3$ km). These results indicate
476 that large-scale variation in tectonic and climatic processes play a governing role in generating the
477 latitudinal diversity gradient (Field *et al.*, 2009). However, our results also indicate that the distri-
478 bution of taxa, and their coherence and phase with environmental variables, is highly location- and
479 scale-dependent. For example, whereas carnivorans and amphibians are in phase and coherent with
480 mean annual precipitation and temperature at wavelengths $> 10^4$ km, that is not true at smaller scales
481 (i.e. shorter wavelengths). Significant deviations from the latitudinal diversity gradient indicate that
482 external variables such as elevation, climatic patterns and tectonic history, play important roles in
483 determining biodiversity at specific locations and scales (e.g. Archibald *et al.*, 2010, 2013; Hagen *et al.*,
484 2021; Jones *et al.*, 2022; Mannion *et al.*, 2014; Saupe, 2021; Song *et al.*, 2020; Yasuhara *et al.*, 2017).

485

486 Spectral analyses highlight the importance of the tropics for biodiversity, in particular for amphi-
487 bians, for which local changes in elevation and mean annual temperature (but not annual temperature
488 range) are highly coherent with species richness. These results are consistent with the idea that in-
489 creased resource availability in the tropics may generate higher primary productivity, supporting a
490 greater number of individuals within a given area (i.e. higher carrying capacity), and therefore a
491 greater number of different species (e.g. Fritz *et al.*, 2016; Gillman *et al.*, 2015; Hawkins *et al.*, 2003;
492 Kessler *et al.*, 2014). Our results support the suggestion that elevated topography at the tropics is
493 more likely to result in increased species diversity when compared to higher latitudes (Ghalambor
494 *et al.*, 2006; Janzen, 1967; Polato *et al.*, 2018). However, this trend is not uniformly observed across
495 taxa and for all continents. Species richness of carnivorans, for example, has no significant coherence
496 with elevation or temperature range in the tropics, which suggests that this group is largely unaf-
497 fected by the challenges posed by tropical mountain ranges. This might reflect the group's relatively
498 unusual biogeographical history and seemingly high dispersal ability, with carnivorans originating at
499 high latitudes and dispersing into the tropics, with net diversification rates comparable in tropical
500 and temperate regions (Rolland *et al.*, 2015). Power spectral slopes for such taxa are steeper (more
501 negative) at shorter wavelengths, whereas more environmentally-sensitive taxa, such as hummingbirds
502 and amphibians, have shallower spectral slopes at longer wavelengths within tropical latitudes.

503

504 Cross wavelet power and coherence indicate that species richness is decoupled from short wavelength
505 ($\lesssim 10^3$ km) changes in elevation, temperature, annual temperature range and precipitation at nearly
506 all locations, except for certain taxa within the tropics. Locally, uplifted topography can be highly co-
507 herent with species richness. Trends across the Americas are reflected in global, latitudinally-averaged,
508 transects and for other continents. In general, the species richness of taxa such as hummingbirds and
509 amphibians is strongly and positively correlated with precipitation rate and temperature, except in
510 Africa, where high temperatures may limit availability of water.

511

512 3.7 Conclusions

513 In summary, wavelet power spectral analysis provides insight into the coherence between species rich-
514 ness and environmental variables. Species richness is shown to vary as a function of location and scale.
515 Comparisons with topography, temperature and precipitation show that species richness tends to be
516 highly coherent with external forcing at large scales (wavelengths $> 10^4$ km). Phase difference between
517 signals reveals that species richness is in-phase with precipitation and temperature, and anti-phase with
518 elevation and annual temperature range, at these scales. However, these relationships are dependent

519 on scale and taxon. At smaller scales, richness of bats, songbirds, hummingbirds and amphibians tends
520 to be greatest in the tropics, where calculated coherence highlights the importance of topography and
521 temperature range for determining species richness. Carnivorans, in contrast, show little coherence
522 with environmental variables at these scales in the tropics. Instead, they are most coherent in the
523 vicinity of terrestrial plateaux, for example the Colorado Plateau and Altiplano. These observations
524 suggest that large scale ($> 10^3$ km) variations in environmental variables determine almost all of the
525 distribution of terrestrial vertebrates. Smaller scale ($\lesssim 10^3$ km) variation can play an important role
526 locally, particularly within the tropics. These results highlight the general importance of environ-
527 mental change at the scale of tens degrees of latitude for determining biodiversity. They also indicate
528 that changes at smaller scales are comparatively more important in the tropics for determining species
529 richness. Crucially, these results could be used to predict the changes in biodiversity that could arise
530 from different future Earth climate change scenarios.

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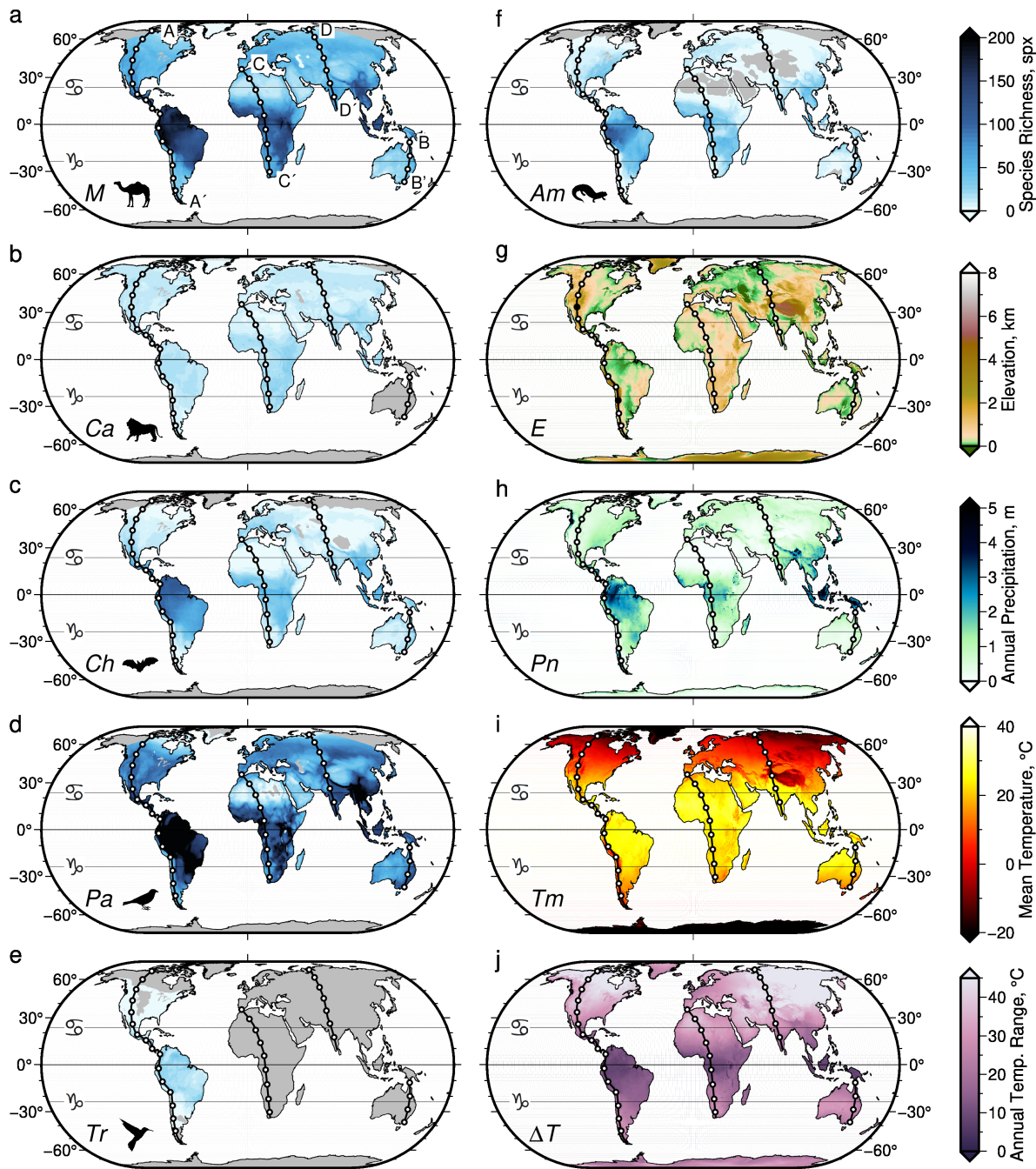


Figure 1: Global patterns of species richness and environment. (a) All Mammalia (*M*, mammals), (b) Carnivora (*Ca*, carnivorans), (c) Chiroptera (*Ch*, bats), (d) Passeriformes (*Pa*, songbirds), (e) Trochilidae (*Tr*, hummingbirds), (f) Amphibia (*Am*, amphibians); spx = species per 10×10 km pixel (Jenkins *et al.*, 2013); horizontal lines = Tropics of Cancer (northern), Capricorn (southern), and Equator; A—A' = transect through Americas investigated here; B—B', C—C', D—D' = transects investigated in Supporting Information. Global latitudinal mean transects also studied therein and in Figure 4. (g) Elevation (*E*) from ETOPO1 global model with horizontal resolution of 1 arc-minute (Amante & Eakins, 2009); filled circles on A—A' = Colorado Plateau/Mexican Highlands and Andean Altiplano. (h)–(j) Mean annual precipitation rate (*Pn*), temperature (*Tm*), and temperature range (ΔT) from 1981–2010 (Karger *et al.*, 2017).

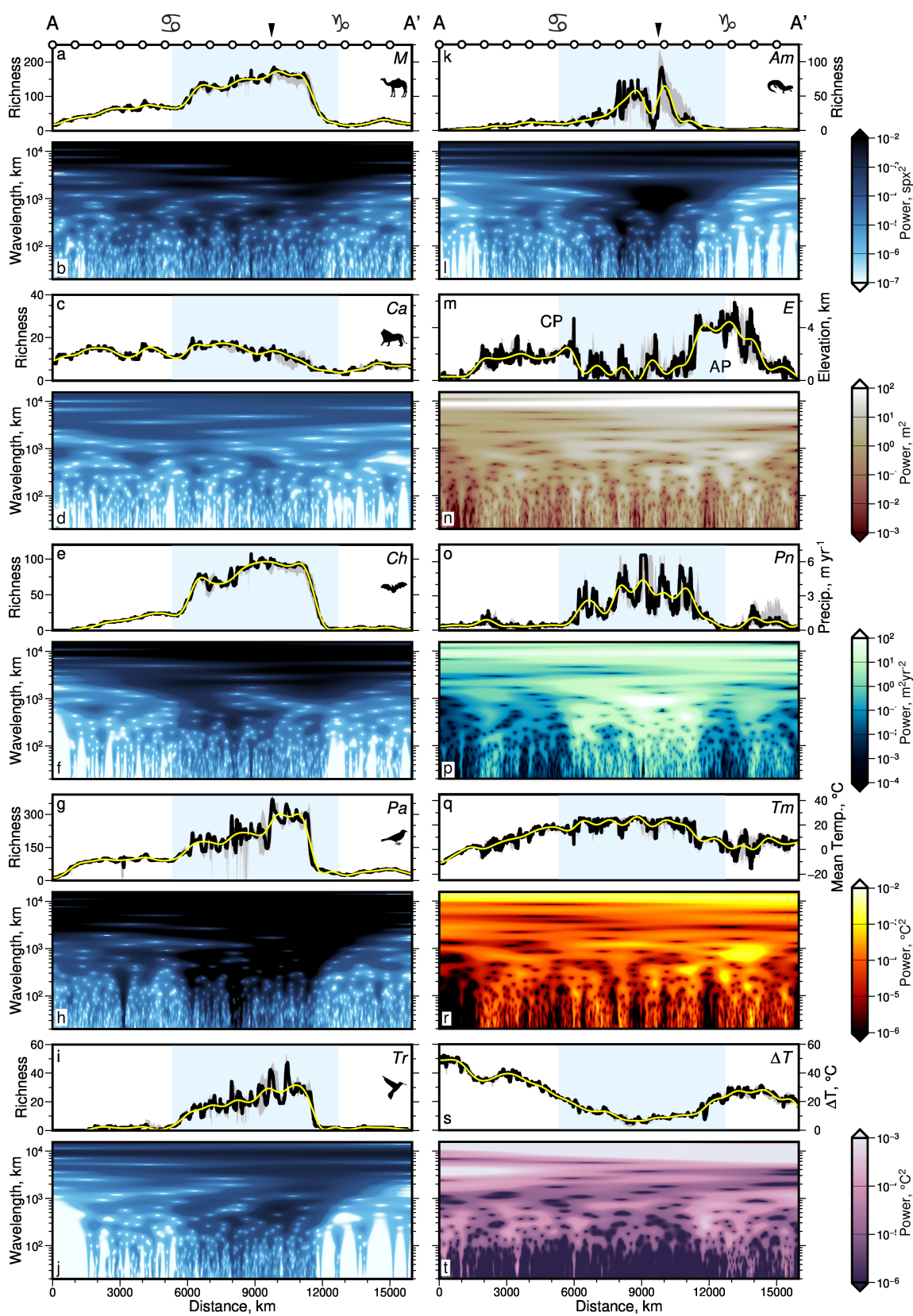


Figure 2

Figure 2 (preceding page): Wavelet transforms of species richness and environment. (a) Black line = species richness of Mammalia (M) along transect A—A'; gray bands = 100 km wide swaths centred on A—A'; blue bands = tropical latitudes; white circles are shown every 1000 km, see transect A—A' in Figure 1; black arrow and symbols above top axis = Equator and tropics as in Figure 1. Yellow line = inverse wavelet transform of signal, filtering to pass only wavelengths > 1000 km; mean difference to input signal = 3.0 ± 3.3 (1σ) spx. (b) Continuous wavelet transform of mammal richness spatial series (black line in panel a). Colors = rectified spectral power as a function of location and scale (wavelength); spx = species per pixel. (c)–(t) As (a)–(b) but for Carnivora (Ca), Chiroptera (Ch), Passeriformes (Pa), Trochilidae (Tr), Amphibia (Am), elevation (E), mean annual precipitation rate (Pn), temperature (Tm) and temperature range (ΔT) along transect A—A' (Amante & Eakins, 2009; Jenkins *et al.*, 2013; Karger *et al.*, 2017). Mean differences between signals and inverse transforms filtered to remove wavelengths < 1000 km = 0.7 ± 0.6 spx (Ca), 1.5 ± 2.1 spx (Ch), 11.6 ± 16.7 spx (Pa), 1.5 ± 2.5 spx (Tr), 2.9 ± 5.3 spx (Am), 0.36 ± 0.3 km (E), $0.35^{+0.5}_{-0.35}$ m/yr (Pn), 2.2 ± 2.2 °C (Tm), and 1.2 ± 1.1 °C (ΔT). See Supporting Information for results for transects B—B', C—C', D—D' and average global latitudinal transect. Note high spectral power concentrated at wavelengths $> 10^3$ km for all series. High species richness power (darker patches) at shorter wavelengths tends to be concentrated within the tropics.

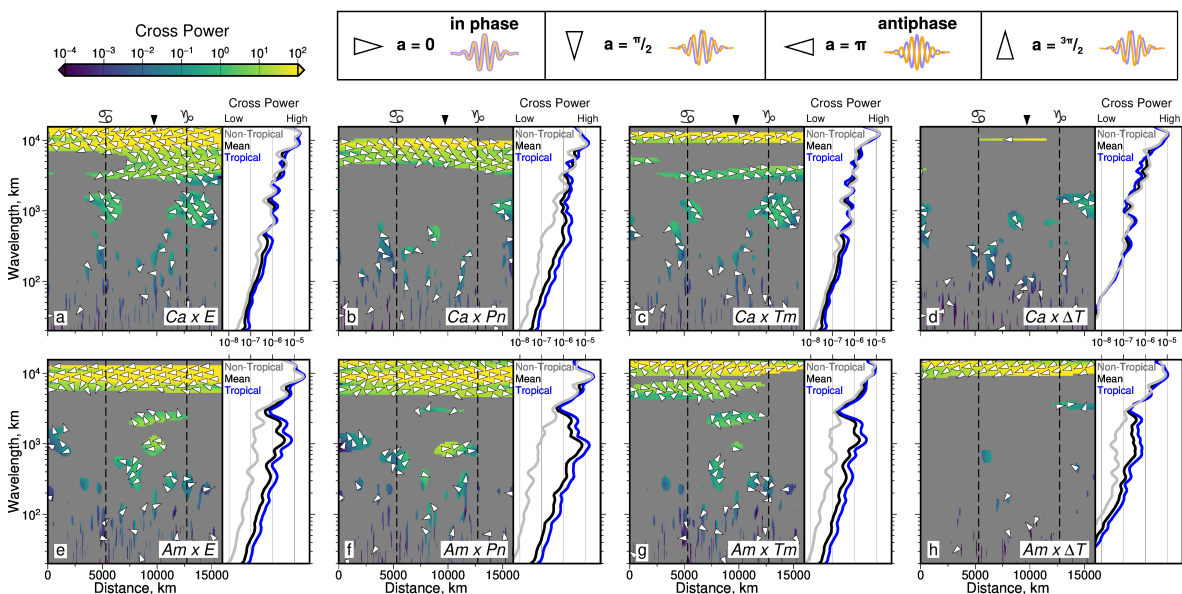


Figure 3: Coherence, cross power, and phase difference between species richness and environmental variables. (a) Comparison of Carnivora (Ca) and elevation (E) as a function of location and scale along transect A—A' (Figures 1–2). Colors = cross wavelet power; yellow = co-located large (positive or negative) amplitude signals. Gray masks regions with coherence below 90% significance level (see body text, Materials and Methods). Arrows = phase difference between spatial series: right/left pointing = in-phase/anti-phase (see guide above panels b–d). Black arrow and symbols above plot = Equator and tropics, as in Figure 1. Side panel: black/blue/gray lines = distance-averaged cross wavelet power of all/tropical/non-tropical latitudes (see Figure 2). High cross power = large co-located amplitudes in the two spatial series. (b)–(d) Comparison of Carnivora and mean annual precipitation rate (Pn), temperature (Tm) and annual temperature range (ΔT). (e)–(h) Comparison of amphibian species richness and same environmental variables as panels a–d. Statistically significant coherence is concentrated at wavelengths $> 10^3$ km, where species tend to be in- or anti-phase with environmental variables. The least statistically significant coherence is for Carnivora and temperature range (note gray mask across most of panel d).

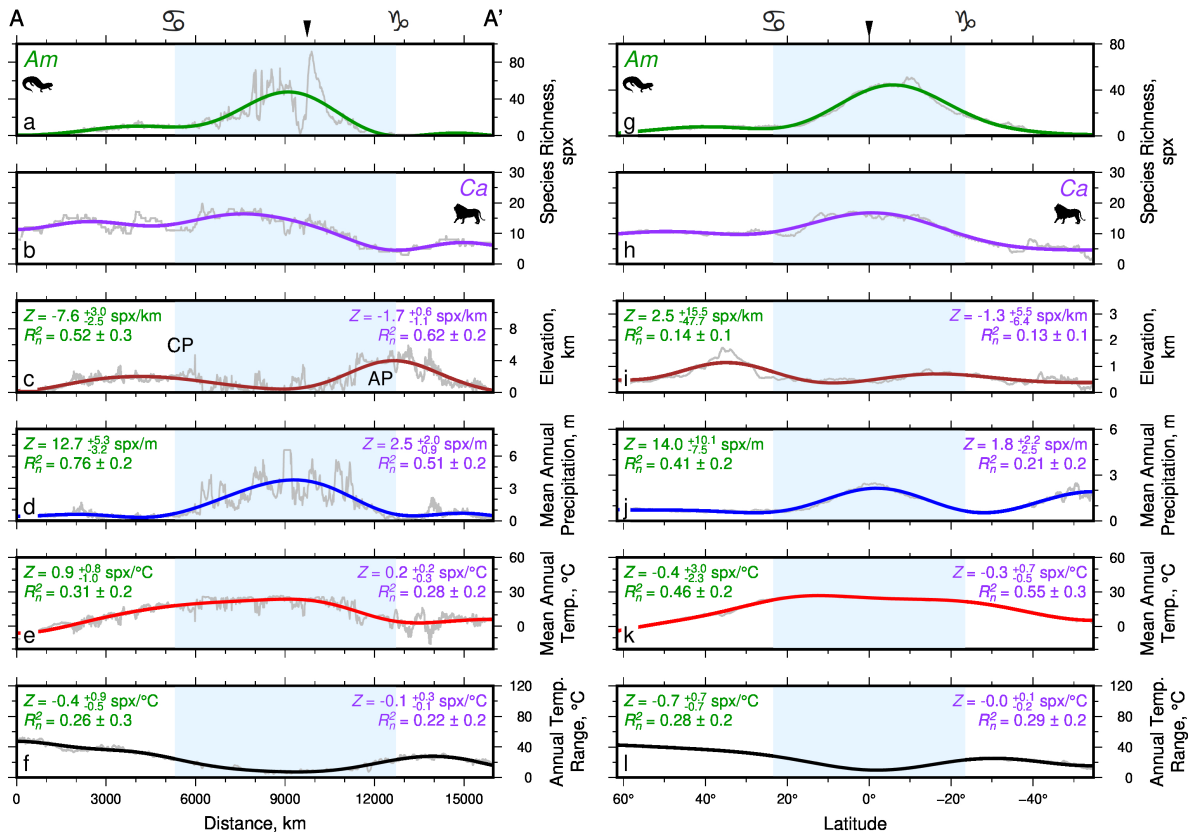


Figure 4: Summary of coherence between species richness and environmental variables at large scales. (a-f) Species richness and environmental variables along Americas transect A—A' (see Figures 1–2). Blue band = tropical latitudes; black arrow = Equator; symbols above x axis = Tropics of Cancer and Capricorn. (a) Amphibian species richness. Gray = full-resolution observed species richness trend (see Figure 2k). Green = inverse wavelet transform showing filtered amphibian species richness at wavelengths ≥ 3756 km (i.e. one quarter of transect length scale). Mean difference between gray and green lines = 4.4 spx. (b) As (a) but for carnivoran species richness. Mean difference between gray and green lines = 1.2 spx. (c)–(f) As (a)–(b) but for elevation, mean annual precipitation, mean temperature, and annual temperature range, respectively. Z , R_n^2 = mean admittance and coherence, respectively, between species richness and given environmental variable. Green text = admittance for amphibian species richness; purple text = admittance for carnivoran species richness. (g)–(l) As (a)–(f) but for mean global latitudinal transects.