Biophysical underpinning of astroglial physiology probed with realistic cell models

Leonid P. Savtchenko*[1], Lucie Bard*[1], Thomas P. Jensen[1], James P. Reynolds[1], Igor Kraev[2], Mikola Medvedev[2], Michael G. Stewart[3], Christian Henneberger[1,3,4], Dmitri A. Rusakov[1]

[1] UCL Institute of Neurology, University College London, UK;

[2] The Open University, Milton Keynes, UK

[3] Institute of Cellular Neurosciences, University of Bonn, Germany

[4] German Center of Neurodegenerative Diseases (DZNE), Bonn, Germany

* Equal contribution

Correspondence: Dmitri Rusakov (d.rusakov@ucl.ac.uk) or Leonid Savtchenko (skaalsa@ucl.ac.uk)

Electrically non-excitable astroglia take up neurotransmitters, buffer extracellular

K^+ and generate Ca^{2+} signals that release molecular regulators of neural circuitry.

The underlying machinery remains enigmatic, mainly because the nanoscopic,

sponge-like astrocyte morphology has been difficult to access experimentally or

explore theoretically. Here, we have systematically evaluated the multi-scale

morphology of protoplasmic astroglia to construct a realistic multi-

compartmental cell model that can be biophysically interrogated in NEURON

computational environment. This approach has been implemented as an

astrocyte-model builder ASTRO. As a proof of concept, we explored a

hippocampal astrocyte reconstructed *in silico* against a battery of physiological

and imaging experiments. This exploration has unveiled some basic features of

astroglial physiology inaccessible empirically, such as the characteristic length

of membrane voltage propagation, membrane effects of local glutamate

transport, spatiotemporal dynamics of intracellular K^+ redistribution, key Ca^{2+}

buffering properties, and some basic relationships between free Ca^{2+} dynamics

and experimental readout of fluorescent Ca^{2+} indicators.

Astroglia have emerged as an essential contributor to neural circuit signalling in the
brain. In addition to the well-established mechanisms of neurotransmitter uptake and
extracellular K^+ buffering, electrically passive astrocytes appear competent in
transducing, integrating and propagating physiological signals using intracellular Ca^{2+}
transients [1-3]. Astroglial Ca^{2+} signals range from local hotspots to global Ca^{2+} waves,
displaying a variety of dynamic ranges and time scales (reviewed in [4, 5]). Tri-
dimensional (3D) reconstructions of astroglia using electron microscopy (EM) have long
revealed a fine system of nanoscopic processes [6-10] that pervade the entire cell
expanse [11-13]. Deciphering the cellular cascades and the biophysical mechanisms that
shape Ca^{2+}-dependent signalling in this sponge-like system has been a challenge [14, 15].
A similar challenge pertains to our understanding of how the complex geometry of
astrocytes controls intracellular ionic homeostasis during rapid local influx of K^+, other
ions, or water molecules [16-23].

In contrast, cellular mechanisms underpinning neuronal physiology have been explored
and understood in great detail. This owes to the fact that rapid advances in patch-clamp
electrophysiology and cell imaging in neurons have been recapitulated and
mechanistically interpreted using realistic biophysical cell models *in silico*. The multi-
compartmental modelling of various neuronal types developed in the NEURON
computational environment [24, 25] has been highly instrumental in advancing our
knowledge about neural signal integration in the brain. There have also been numerous
attempts to simulate astroglial function, based predominantly on a reductionist
approach (recently reviewed in [26, 27]). Aimed to answer a specific scientific question,
such models would normally focus on kinetic reactions inside astroglia [28-30], between
astroglial and neuronal compartments [16, 31-33] or, alternatively, on the regulatory
influences of astroglial signalling in neuronal networks [34-36]. These studies have
provided some important insights into the biophysical basis of astroglial activity and its
adaptive roles. However, their scope would normally exclude complex 3D morphology
of astrocytes, intracellular concentration gradients and heterogeneous molecular fluxes,
uneven landscapes of cell membrane potential, or complex relationships between Ca^{2+}
kinetics and Ca^{2+}-sensitive fluorescence recordings. Thus, to date there have been no
attempts to integrate cellular functions of astrocyte on multiple levels, in one entity *in*
silico, with the realistic morphology playing a major role in shaping astroglial identity [14].

Our aim was therefore three-fold. First, to generate a modelling approach that

recapitulates fine astroglial morphology on multiple scales while retaining full

capabilities of biophysical simulations enabled by NEURON. We have therefore

developed (MATLAB- and NEURON-based) computational algorithms and software

that, firstly, use experimental data to recreate the space-filling architecture of

nanoscopic astroglial processes, and, secondly, make this cell architecture NEURON-

compatible. Our case study focused on the common type of hippocampal protoplasmic

astroglia, which has been the main subject of studies focusing on synaptic plasticity and

neuron-glia interactions [37-40]. We have combined patch-clamp electrophysiology, two-

photon excitation (2PE) imaging and 2PE spot-uncaging, fluorescence recovery from

photobleaching (FRAP), viral transduction of astroglia-targeted Ca^{2+} indicators *in vivo*,

and quantitative correlational 3D EM to systematically document the multi-scale

morphology and some key physiological traits of these cells. Based on these

experimental constrains, we have built a multi-compartmental 3D cell model that

recapitulates known features of hippocampal astrocytes and is fully integrated into the

NEURON environment. Correspondingly, we have equipped the latter with additional

functionalities relevant to astroglial studies, such as: generation of nanoscopic

processes that fill in tissue volume; options for incorporating endfoot and gap junctions;

options to arrange extracellular glutamate spot-uncaging and localised extracellular K^+

rises; menu for running 3D FRAP experiments; intracellular Ca^{2+} signal generation

options; settings to replicate conditions of fluorescence imaging.

Our second objective was to implement this modelling approach as a user-friendly,

interactive simulation instrument - cell model builder - capable of recreating and probing

various types of astroglia *in silico*. Thus, we have integrated our algorithms and

software as a modelling tool ASTRO, which enables an investigator to generate

morphological and functional astroglial features at various scales (current version at

https://github.com/LeonidSavtchenko/Astro).

Finally, as a proof of concept, we explored and probed the test-case model against

experimental observations, aiming to understand some important aspects of astroglial

physiology that cannot be accessed through direct experimental probing. In our case

study focusing on *stratum radiatum* astroglia, we have therefore evaluated key

electrogenic features of the cell membrane, basic aspects of intracellular K^+ dynamics,

the range of intracellular Ca^{2+} buffering capacity, and how the classical molecular

machinery of Ca^{2+} 'puffs' and 'sparks' could explain the common observations in

astroglia using fluorescence imaging of Ca^{2+}-sensitive indicators. The results suggest

that ASTRO could provide an important means for testing physiological hypotheses and

causal interpretation of experimental observations pertinent to astrocytes.

RESULTS

Reconstructing gross morphology of live astroglia: stem tree

The gross morphology of CA1 astrocytes has long been documented [11, 41, 42] pointing to

the overall cell radius of 30-50 μm, the somatic diameter of 7-15 μm, 4-9 primary

processes [11, 43], and a main-branch ramification index (RI, probability of bifurcation per

branch) of less than one [44]. To elucidate the structural detail of CA1 astrocytes *in situ*

we employed acute hippocampal slices. Individual astroglia were held in whole-cell,

loaded with the bright morphological tracer Alexa Fluor 594 (Methods), and imaged

using two-photon excitation (2PE; Fig. 1a,b), a procedure shown quantitatively to

faithfully reveal fine cell morphology [42].

Our modelling strategy was to start with the principal branch structure (cell 'stem tree')

which could be resolved in 2PE optical images (i.e., with a diameter above the

diffraction limit, 0.3-0.5 μm; Fig. 1a,b) and 3D-reconstructed using computer tools that

have been previously validated in neuronal studies. Here, we have employed two

complementary approaches. In the first approach (Methods), the cell was imaged in a *z*-

stack (Fig. 1a,b), individual images were corrected for the depth-dependent

fluorescence signal decrease, identifiable stem-tree branches were traced semi-

automatically and 3D-reconstructed using Simple Neurite Tracer (Fiji-ImageJ, NIH; Fig.

1c; Methods). This routine was similar to that of reconstructing 3D geometry of neuronal

processes, as described widely in the literature. To store the recorded structure data in

a NEURON-compatible format (Fig. 1d) we used Vaa3D (Allen Institute). Alternatively,

the entire 3D reconstruction procedure could be carried out using commercially

available Neurolucida (MBI), which provides a NEURON-compatible output.

The second approach was to build the stem tree of the 'typical' (or 'average') astrocyte

from the population under study. First, the stem tree structure was taken from a digital

library: in our case study we used a library neurogliaform cell (P32-DEV136, database

http://neuromorpho.org), which displays gross morphology similar to that of the

astrocyte (Fig. 1e, diagram). Next, the numbers and diameters of tree branches were

adjusted to match experimental measurements from an experimental sample (in our

case 13 astrocytes, 98 individual branches; Fig. 1e). The stem tree options also include

settings to construct and explore an endfoot structure (Supplementary ASTRO User

Guide, pp 20-21).

The endfoot

The endfoot surrounding blood vessels is a key morphological and functional feature of

most astrocytes [45, 46]. Because its morphology varies enormously among cells, it would

seem appropriate to use direct experimental 3D reconstructions (as in Fig. 1a-d) to

incorporate this feature into the cell architecture. ASTRO provides a separate menu for

constructing the endfoot, connecting it to the 'main' arbour, and populating it with the

nanoscopic processes (Supplementary ASTRO User Guide, pp 20-21). All nanoscopic

process structures and biophysical mechanisms available in the present model builder

(see sections below) could be incorporated into the endfoot. However, as it represents

a self-contained, highly specialised cell compartment, it will require a separate study to

fully develop its realistic biophysical machinery and incorporate it in accord with

experimental observations. Simulation tests in the present study focus therefore on the

'main' astrocyte arbour, which is thought to play a key role in astroglia-neuron

interactions [4, 5, 14, 47]. Nonetheless, parts of the main arbour may include processes that

surround small blood vessels.

Nanoscopic astroglial processes: experimental measurements

Once the cell stem tree was ready and NEURON-formatted, it was imported by ASTRO

for the next step of morphological reconstruction: stochastic generation of nanoscopic

processes. Whilst the stem tree branches can be traced in the light microscope, the

bulk of astroglial morphology comprises irregularly shaped, nanoscopic processes that

occur throughout the cell expanse and appear as a blur in optical images (Fig. 1a,b). 3D

EM reconstructions have been used to detail the spatial organisation of these

processes [10, 12, 13, 42]. To systematically quantify such structures, we employed a

correlational-EM approach based on our earlier 3D EM studies of astroglia [10, 48].

Individual astrocytes were patched whole-cell in acute brain slices and, after cell filling

with biocytine and subsequent DAB conversion, traced and reconstructed in a

comprehensive serial-section survey (Methods; Supplementary Fig. 1). A recent EM

study has suggested that chemical fixation via heart perfusion might lead to 30-35%

volumetric tissue shrinkage and concomitant changes in fine astroglial morphology [49].

To minimise such effects, we used rapid fixation of thin acute slices by submersion,

which was earlier shown to cause much smaller hippocampal tissue shrinkage (linear

shrinkage ~5%) [50] preserving the extracellular space volume fraction of ~0.12 (reviewed

in [51]) which was close to ~0.15 seen under rapid cryofixation [49].

Obtaining 3D reconstruction of the entire astrocyte at nanoscopic resolution is

technically difficult and it may not necessarily represent the 'typical' cell. We therefore

focused on the characteristic morphological features of 3D-reconstructed fragments

sampled from multiple individual astrocytes from the CA1 population, as detailed earlier

[10] (Fig. 2a). The aim of this exploration was to extract key morphometric characteristics

of the nanoscopic astrocyte processes in order to replicate such 'typical' processes in

the cell model.

We therefore developed algorithms and MATLAB-based software (integrated in

ASTRO) that enabled us to sample and store small astrocyte processes from our sets

of 3D-reconstructed astroglial fragments (Fig. 2a,b; Supplementary ASTRO User Guide

p. 12). The 3D EM reconstruction tools used here have been detailed earlier [10]; ASTRO

can use other text-formatted files containing 3D coordinates of the cell membrane

'mesh', which is a common outcome of EM 3D reconstruction programs. This procedure

aimed at building a database of astroglial nanoscopic structures: each process would

comprise a *z*-stack (arbitrarily long) of serial 60 nm sections, with each section

represented by reference points scattered on its 'polygonal' perimeter (Fig. 2c;

Supplementary ASTRO User Guide pp. 12-13).

Nanoscopic astroglial processes: NEURON-compatible transformation

Our cell modelling approach aimed to take advantage of the NEURON computational

environment enabling multiple mechanisms of cellular biophysics. However, NEURON-

built cell models employ cylindrical compartments designed to follow the shape of

neuronal dendrites or axons. Because astroglial processes have much less regular

shapes (Fig. 2a-c), we carried out a separate investigation to establish whether and

how cylindrical NEURON compartments can be used to recapitulate geometry and

biophysical properties of realistic astrocyte processes.

To attempt the transition from real (3D EM-based) to NEURON-compatible shapes, we

transformed individual 'polygonal' z-stacks of 3D-reconstructed processes into z-stacks

of cylindrical (disk-shaped) slabs (Fig. 2d). Here, the adjacent polygonal sections with

cross-section areas S_i and S_{i+1} which had an intersection area of T_i (Fig. 2d, left and

middle), were represented by two 'main' cylinder slabs with base areas S_i and S_{i+1}

(termed 'leaves') plus a 'transitional' cylinder slab (termed 'stalk') with the base area T_i

(Fig. 2d, right). This transformation approximately preserved the 'diffusion bottleneck'

and the surface-volume relationship of the original shape. Applying this rule

systematically transformed 3D-reconstructed nanoscopic processes into the NEURON-

compatible, cylinder-based shapes, with the volume and surface areas similar to those

of the original shapes (Fig. 2e).

Next, we employed Monte Carlo simulations to systematically compare two key

biophysical traits between the original and the NEURON-compatible shapes, the

intracellular diffusion transfer rate and the dynamic electrical impedance. These tests

involved 'injecting' 3000 Brownian particles into one end of the shape and counting their

flux rate at the other end, with or without an electric field (Fig. 2f; example in

Supplementary Movie 1). The mathematical algorithms involved were detailed, tested

and validated against experimental data in our previous studies [52-54]. In the majority of

cases, our tests showed remarkable similarity between the two shapes in their dynamic

electrical and diffusion conductance properties (Fig. 2g). In the remaining cases, minor

stochastic variation of the cylindrical compartment diameters achieved a similar match.

After we accumulated a representative population of cylindrical compartment radii from

the 3D EM measurements (example in Supplementary Fig. 2), we also found that the

biophysical match between original and transformed shapes remained when the

cylindrical compartments were shuffled randomly with respect to the original shape

(example in Supplementary Fig. 3). Thus, the population of nanoscopic astroglial

processes could be represented in a NEURON-compatible format based on the

statistical properties of 3D-reconstructed cell fragments. The sampling and

transformation procedure for nanoscopic processes, as described above, has been

integrated in ASTRO for investigators to explore, monitor, and validate (Supplementary

ASTRO User Guide, pp. 11-18).

These measurements provided the data set for the astroglial stem tree (Fig. 1d,e) to be

populated with nanoscopic processes in accord with their experimentally obtained

statistics. However, this stochastically simulated 'morphogenesis' had to be further

constrained by the experimentally observed tissue-filling properties of astroglia, as

described below.

Tissue volume fraction and astrocyte surface-to-volume ratios

The tissue volume fraction (VF) occupied by astroglia in the hippocampal neuropil has

been estimated to range between 5-10% [8-10, 55]. VF is a key descriptor of astrocyte

morphology because astrocytes occupy adjacent tissue domains, with very little overlap

[11, 56], while their fine processes fill the local volume in a sponge-like manner [10, 13, 42]. The

distribution of astroglial VF in local tissue thus provides an important guide to the

astrocyte architecture.

2PE microscopy provides a straightforward tool to monitor VF of live astroglia *in situ*

because the fluorescent signal it collects represents emission from within a thin (~1 μm)

layer at the focal plane (Supplementary Fig. 4a). Thus, the signal intensity produced by

the dye-filled astroglia scales with the local VF occupied by one or more astroglial

processes (Supplementary Fig. 4b). At the same time, the 6-10 μm wide astrocyte

soma imaged in the same 1 μm focal plane will correspond to 100% VF occupied by the

dye-filled cytosol (at least in its brightest part which is least populated with cell

organelles; Supplementary Fig. 4c) [10, 14]. Thus the local- to-somatic fluorescence

intensity ratio provides direct readout of astroglial VF (Fig. 3a). We therefore first

established the typical distribution of VF within individual cell arbours, as a function of

the distance from the soma, for CA1 astrocytes (Fig. 3b; blue dots, mean ± SEM for 13

cells). Next, we adjusted the computer-generated nanoscopic morphology to match this

VF distribution (Fig. 3b, red dots).

The latter is achieved by systematically adjusting two key structural parameters in the

model using ASTRO menus. The first is the average size of nanoscopic processes,

determined as the number of leaves per process (up to 100; Supplementary ASTRO

User Guide, Fig. 16, p. 19). The second parameter is the 'seed' density for nanoscopic

processes to be planted on the stem-tree branches, which normally ranges between

one to three (SeedNumber variable; Supplementary ASTRO User Guide, p. 20).

Another morphological feature essential to the biophysical identity of cell processes is

their surface-to-volume ratio (SVR): this value determines how the transmembrane

molecular fluxes are converted into intracellular concentration changes. Careful

stereological analyses of hippocampal astroglial processes using 3D quantitative EM

arrive consistently at the SVR in the 15-25 µm^{-1} range [8, 41, 55, 57, 58]. The astrocyte

fragments sampled in our experiments (Fig. 2a-c) had SVRs fully consistent with this

range, and the modelled nano-shapes (Fig. 2f, Supplementary Fig. 3b) faithfully

reproduced this SVR range across the entire simulated cell morphology. The average

SVR could be further fine-tuned in ASTRO, by adjusting the distribution of 'leaves' and

'stalks' in stochastically simulated nanoscopic protrusions (Supplementary ASTRO User

Guide, pp. 21-23).

Thus, by systematically generating nanoscopic processes on the 'typical' stem tree (Fig.

1e) we arrived at what appears to be a realistic, NEURON-compatible CA1 astrocyte

geometry comprising 35000-45000 individual compartments (Fig. 3c; see

Supplementary Table 1 for the summary of morphological features pertinent to the

present case study of CA1 astroglia). Our final test for the realistic nature of the

astrocyte *in silico* was to see if its architecture has intracellular diffusion connectivity

similar to that in live astroglia. To address this we carried out a series of experiments

and simulations based on fluorescence recovery after photobleaching (FRAP), as

outlined below.

Internal connectivity of astroglia: experiment *versus* model

FRAP protocols have long been used to gauge mobility of fluorescent molecules in live

cells: the speed with which fluorescence recovers in a laser-bleached cell region

reflects diffusion speed of the fluorescent molecules under study [59]. Notably, a 2PE

FRAP technique applied to water-soluble indicator Alexa Fluor 594 was used to monitor

diffusion transport through the ultrathin necks of neuronal dendritic spines [60]. We sought

to apply a similar method to the bulk of nanoscopic astrocyte processes using line-scan

FRAP: by bleaching the dye molecules within a thin cylindrical volume across the

astrocyte tree, this approach should gauge average diffusion connectivity within the

adjacent structures (Fig. 3d; Methods), as shown previously [61].

Here, we first confirmed that the intracellular fluorescence showed full recovery well

before a 60 s interval (Supplementary Fig. 5a,b), thus indicating no concomitant signal

from what could be an immobile dye fraction. Furthermore, 60 s after the first cycle, the

second FRAP cycle faithfully reproduced the time course of photobleaching and

recovery displayed during the preceding FRAP cycle (Supplementary Fig. 5c). This

pointed to the stability of FRAP experiments in the current settings.

We next simulated this experiment in the astrocyte model (Fig. 3e; Supplementary

Movie 2) and employed an 'alternating' FRAP protocol (with laser shutter repeatedly on-

off), to compare simulation outcome with experimental observations. We found that by

adjusting one free parameter in the model, the rate of Alexa line-scan photobleaching, it

was possible to achieve a good match between experimental and simulated data in

individual cells (example in Fig. 3f). The summary (n = 10 recorded cells) also showed a

good correspondence between recorded and simulated data for the FRAP rates at

variable distances from the soma (Fig. 3g). These observations indicated that the

simulated cell morphology was a good representation of the inner connectivity of real-

life CA1 astrocytes.

Passive electrical properties of astrocytes

Specific membrane conductance G_m is an important determinant of electrical signal

propagation in non-excitable cells, such as protoplasmic astroglia. First, we measured

input resistance of live astroglia in whole-cell mode: the average value (mean ± SEM:

2.66 ± 0.31 MΩ, n = 15; Fig. 4a) was readily consistent with the previous

measurements [37, 38]. Next, we pulled outside-out patches, estimated the patch

membrane area (using the classical voltage-step method [62], Methods) and recorded

patch resistance: these measurements gave us the average astrocyte G_m (mean ±

SEM: 0.69 ± 0.18 mS/cm^2; Fig. 4b). Finally, we asked whether this experimental value

of G_m (Fig. 4b) could be reproduced by an astrocyte model constrained by two other

independent experimental parameters, (steady-state) input resistance R_i and total

membrane area S_{mem}, in accord with Ohm's law $G_m = (S_{mem}R_i)^{-1}$. Because modelled

nanoscopic processes are generated stochastically, we have obtained a small

representative sample of the astrocyte in question, by repeating the stochastic

reconstruction of tissue-filling nanoscopic processes, in accord with the morphological

constraints (Fig. 4c). In the sample, the cell volume (resulting from stochastic

simulations) varied well within the range of 3000-4800 μm^3 characteristic of CA1

astroglia [11, 55]. In each cell in the sample, the free parameter G_m was adjusted until the

cell R_i matched the measured value (Fig. 4a). This test produced the average G_m value

of 0.78 ± 0.04 mS/cm^2 (Fig. 4c), which was indistinguishable from the experimental G_m

value (Fig. 4b) or the earlier empirical measurements [18]. Thus we have confirmed that

our model reproduces 'correct' experimental values of G_m without artificially imposing it:

in other words, stochastically generated nanoscopic processes replicate realistic

membrane properties of astrocytes.

To understand how far a membrane voltage/current signal propagates in electrically

leaky astroglial membranes, we simulated a simple voltage-clamp experiment. This

showed that large currents induce very modest astrocyte membrane depolarisation,

which propagates with an electrical 3D-space constant of 30-60 μm depending on the

signal frequency (Fig. 4d, left). Intriguingly, a membrane voltage signal generated at the

periphery is attenuated with a smaller space constant of 25-30 μm (Fig. 4d, right). This

asymmetry (reported previously in nerve cells [63]) arises due to the location-dependent

distribution of membrane conductance near the signal cite implying that similar signals

arising at different astrocyte loci could have different membrane effects.

Because astrocyte membranes are enriched in potassium channels, in particular $K_{ir}4.1$

type [17-19], it would seem important to assess their contribution to the voltage signal

propagation. However, the overall unit conductance of $K_{ir}4.1$ in astrocytes is either

compatible with or lower than the membrane current leak due to the other conductance

contributors (including ion channels, exchangers, gap junctions and semi-channels) [16,

64]. This suggests that $K_{ir}4.1$ per se has little effect on the passive voltage spread profile,

even though these channels play a major role in maintaining the astrocyte resting

membrane potential throughout the astrocyte [65]. For the sake of simplicity, in our

baseline test model the multiple membrane mechanisms that hold astroglial resting

membrane potential near -83...-85 mV were represented by a non-specific channel

current. In these conditions, when we added evenly distributed $K_{ir}4.1$, the effect on

voltage propagation was significant only when the overall $K_{ir}4.1$ conductance exceeded

substantially its expected physiological range (Supplementary Fig. 6a). In physiological

circumstances, this scenario could be affected by changes in extracellular K^+

concentration ($[K^+]_{out}$), and by the poorly understood contributing role of other ion

channels and exchangers in the astrocyte membrane (see sections below).

Membrane voltage-current landscape generated by local glutamate uptake

We next asked how the current generated by glial glutamate transporters GLT-1 in

response to local glutamate release (a common physiological scenario) alters

membrane potential across the cell. First, we recorded from a CA1 astrocyte (acute

hippocampal slices) the current response to two-photon spot-uncaging of glutamate (20

ms) at variable distances from the soma (Fig. 4e, traces). To mimic this experiment in

the model, we implemented the GLT-1 transporter kinetics [66, 67] and 'scattered' the

transporters uniformly on the cell surface at an effective density of $\sim 10^4$ μm^{-2} [68]. In these

simulations, glutamate was applied within small spherical areas (radius ~3 μm, duration

~20 ms) at quasi-random distances from the soma *in silico*. An excellent match

between the modelled somatic current and whole-cell recording data could be obtained

(Fig. 4e, plot) by adjusting one free model parameter, the amount / peak concentration

of released glutamate. The model unveiled the dynamic landscape of astrocyte

membrane voltage, under somatic voltage clamp (to mimic experimental conditions),

which varied only within ~0.2 mV across the entire cell (Fig. 4f; Supplementary Movie

3).

To reiterate on the electric properties of astroglia, these data show quantitatively how a

local current hotspot in the electrically leaky astroglial membrane stays highly localised,

with little effect on the voltage landscape (Fig. 4f). Because of the low baseline

membrane resistance, this scenario remained only weakly affected by adding the

voltage-dependent $K_{ir}4.1$ conductance [16, 64] unless the latter exceeded its physiological

range (Supplementary Fig. 6b-d). Again, this test dissects the role of $K_{ir}4.1$ in otherwise

unchanged cell conditions (e.g., resting membrane potential stays unchanged as we

artificially add $K_{ir}4.1$) whereas removal of $K_{ir}4.1$ in real cells could lead to complex

physiological effects [65, 69]. In any case, these tests suggest that local glutamate

transporter currents on their own cannot significantly depolarise astrocyte membrane

away from the region of active transport.

Potassium uptake and redistribution inside astroglia

Rapid uptake of the neural activity-induced excess of extracellular K^+ and its

intracellular redistribution are essential functions of brain astroglia [18-20, 46, 70]. Several

elegant biophysical models have shed light on the kinetics and the tissue-wide spread

of this process [16, 46, 71]. Modelling with ASTRO should help to understand the detailed

spatiotemporal dynamics of K^+ redistribution inside individual astroglial compartments,

on the microscopic scale.

As a proof of concept, we simulated a physiological scenario in which extracellular $[K^+]$

($[K^+]_{out}$) was elevated, from resting 3 mM to 10 mM (level often associated intense local

excitatory activity [72]), inside a 20 µm wide spherical region of the astrocyte-containing

tissue, for 2 seconds (Fig. 5a). The cell model was populated with $K_{ir}4.1$ channels with

the kinetics described earlier [16], and unit conductance of 0.1 mS/cm^2. Upon $[K^+]_{out}$

elevation, these channels were activated homogeneously inside the 20 µm area (peak

current 0.01 mA/cm^2; not shown). This prompted local K^+ entry, leading to a local

increase in $[K^+]_{in}$, which dissipated inside the cell over several seconds after $[K^+]_{out}$

returned to 3 mM (Fig. 4b). During the period of elevated $[K^+]_{out}$, the membrane

potential landscapes reflected very slight depolarisation generated by $K_{ir}4.1$ (Fig. 5c),

not dissimilar to the slight depolarisation arising from local glutamate uptake (Fig. 4f).

These data provided a quantitative illustration of extracellular K^+ buffering and its

intracellular redistribution, suggesting that the latter is likely to be controlled by local K^+

efflux, in particular through $K_{ir}4.1$. Again, although $K_{ir}4.1$ appear to represent a major,

potential-forming membrane astrocyte conductance [16, 64, 65], and further experimental

studies are required to understand whether the effect of other membrane mechanisms

has any significant contribution to K^+ buffering.

To further dissect the contributing role of active K^+ extrusion mechanisms, such as

pumps and ion exchangers, we carried out similar tests (K^+ within a ~15 µm spherical

area, over 200 ms at a current density of 1 mA/cm^2), but with the active removal of

intracellular K^+ by a first-order pump (this may also reflect gap-junction escape) and no

contribution from $K_{ir}4.1$ channels (Supplementary Fig. 7). Again, the model captured the

dynamic spatial landscape of $[K^+]_{in}$, with and without any extrusion mechanisms,

confirming that local K^+ efflux could efficiently limit the spatial spread of $[K^+]_{in}$ elevations

(Supplementary Fig. 7).

Gap junctions and hemichannels

Gap junctions comprising of adjoined connexin proteins connect neighbouring astroglia: this enables electrical current leak and the diffusional flow of molecules across the tissue-wide astroglial syncytium [40, 73]. In addition, astroglial membranes are enriched in connexin hemichannels permitting transfer of large molecules between the cytoplasm and the extracellular medium [74]. To understand their contribution to the membrane conductance in our tests, we blocked them with carbenoxolon (CBX, bath applied 50 µM). This increased cell input resistance by ~30% (Supplementary Fig. 8), consistent with the previous measurements in CA1 astrocytes [75].

To enable theoretical probing of these features, we have incorporated basic gap junction options in the ASTRO menu, both as electric conductance and as a diffusion channel (Methods; Supplementary ASTRO User Guide, pp. 18, 26-28). For the sake of clarity, however, simulation examples in the present study considered gap junctions and hemichannels as a constituent contributor to membrane conductance. Dissecting their precise functional roles will require further dedicated experimental studies aiming to quantify their biophysical properties as well as their 3D distribution across the astrocyte tree [76].

Astroglial calcium waves: probing the impact of calcium buffering

Astrocytes have long been known to generate non-dissipative Ca^{2+} waves communicating physiological signals within and among cells [1, 2]. The underlying cellular mechanisms are thought to engage a variety of Ca^{2+} stores, channels, and pumps involving endoplasmic reticulum and mitochondria (reviewed in [5, 15, 77]). The key molecular machinery here appears to engage Ca^{2+}-dependent Ca^{2+} release from Ca^{2+} stores controlled by inositol 1,4,5-trisphosphate (IP_3) receptors and possibly ryanodine receptors, involving a highly non-linear, sometimes bell-shaped, dependence between their activity and local Ca^{2+} concentration [78-81]. More recently, significant Ca^{2+} transients have been detected in astroglia lacking IP_3 receptors [82-84], although cell-engulfing, 'global' waves seem to involve IP_3 action (reviewed in [4]). In all such cases, a critical parameter that controls intracellular Ca^{2+} signal propagation is the Ca^{2+} buffering capacity of the cytosol [85], the feature explored in detail in nerve cells [86].

The biophysical mechanisms of Ca^{2+} oscillations and Ca^{2+} diffusion and buffering are

enabled in the NEURON environment by default, based on the kinetics of IP_3 signalling

which has been well explored (Methods). In addition, ASTRO package incorporates two

alternative IP_3-dependent Ca^{2+} mechanisms described earlier in theoretical studies of

astroglial Ca^{2+} signalling [26, 28, 87]. In the default configuration, four key parameters

control intracellular Ca^{2+} wave dynamics: resting (or average) IP_3 concentration C_{IP3},

resting Ca^{2+} level $[Ca^{2+}]_0$, concentration $[B]$ of endogenous Ca^{2+} buffer, and buffer

affinity (dissociation constant) K_D. Most studies constrain C_{IP3} within 0.5-3 µM [88-90]. As

for $[Ca^{2+}]_0$, we were able to measure it directly using time-resolved fluorescence

microscopy pointing to the range of 50-100 nM [91]. In contrast, $[B]$ and K_D of

endogenous Ca^{2+} buffers can vary widely across cell types and thus remain unknown

yet essential determinants of intracellular Ca^{2+} dynamics in astroglia.

Similar to the biophysical models of other cell types [88-90], we found that Ca^{2+} waves in

the modelled astroglia could be generated over a wide (physiologically plausible) range

of the aforementioned parameters. To dissect and illustrate the impact of Ca^{2+} buffering

on Ca^{2+} wave propagation *per se*, we compared the wave dynamics in modelled cells

with and without a relatively small amount (10 µM) of mobile Ca^{2+} buffer added (Fig. 6a;

see legend for parameter detail; Supplementary Movie 4). It appeared that adding the

buffer reduced the speed and amplitude of $[Ca^{2+}]$ waves significantly (Fig. 6a).

Intriguingly, this also implies that increased Ca^{2+} buffering in these conditions leads to

longer periods of elevated intracellular $[Ca^{2+}]$. Again, an accurate interpretation of such

results requires a dedicated experimental study.

Astroglial calcium waves: assessing cell calcium-buffering capacity *in vivo*

Because there is an ongoing debate on whether astroglial Ca^{2+} waves seen in acute

slices or in culture are fully physiological [92], we sought to document such waves in live

animals, taking advantage of our earlier protocols of astroglial Ca^{2+} monitoring *in vivo*

[91]. However, imaging intracellular Ca^{2+} waves in the hippocampus *in vivo* involves

mechanically invasive methods the impact of which on astroglial function is not fully

understood. We therefore imaged somatosensory cortex astroglia (accessible with

intact brain surface), which appear to show some remarkable similarities in their major

morphological features and territorial volumes, across ages, with hippocampal

astrocytes [93].

We asked therefore whether [B] and K_D for cortical astroglia could be assessed by

matching the Ca^{2+} wave dynamics documented *in situ* with the modelled outcome. To

reduce the concomitants of *ex vivo* preparations and to take advantage of our earlier

measurements of astroglial Ca^{2+} *in vivo* [91], we recorded spontaneous activity of

astrocytes in the rat somatosensory cortex using a virus-transfected Ca^{2+} indicator

expressed under a GFAP promoter. To minimise the spatiotemporal filtering effects of

free-diffusing Ca^{2+} indicators we used plasma-membrane tethered GCaMP6f, with a ~50

ms Ca^{2+}-sensitive fluorescence response time [94]. The animals were anaesthetized, to

limit bursts of sensory input-evoked regional Ca^{2+} rises that could be mistaken for self-

propagating Ca^{2+} waves [95]. Gross morphology of astrocytes was monitored in a red

channel using bolus-loading of sulforhodamine 101 (SR101) via a pressurized patch-

pipette (Fig. 6b), as described previously [91].

In baseline conditions, we could readily detect spontaneous Ca^{2+} waves that appear to

engulf entire individual cells (spreading over 10-20 μm in the XY plane with an average

frequency of ~5.4 waves a minute over a 160 x 160 μm ROI; Fig. 6c, top panels;

Supplementary Movie 5), in good correspondence with previous observations [96]. The

time first derivative of the fluorescence transients could reveal the spreading wave

fronts of these signals (Fig. 6c, bottom panels). Thus, only the waves spreading in a

clear centrifugal manner (as in Fig. 6c) were selected, to ensure that they represent

single-cell regenerative events rather than signals evoked by synchronous external

influences.

Our measurements indicated that such waves propagated with an average radial speed

of 3.94 ± 0.16 μm/s (mean ± SEM, n = 54 events). Intriguingly, this speed appears

significantly lower than that of the stimulus-induced astroglial Ca^{2+} waves in brain slices

or in culture (15-25 μm/s, reviewed in [97]). The relatively fast speed of Ca^{2+} waves in

slices or in culture could be because the exogenous stimulus, such as agonist bath or

puff application, could prompt a synchronous receptor response over the entire cell

expanse. Similarly, in the awake animals, we could detect local waves resembling those

under anaesthesia, albeit at a lower frequency (and apparently lower magnitude), in

between large synchronous Ca^{2+} elevations engulfing entire regions (Supplementary

Movie 6). Again, the possible overlap of such prevalent synchronous events and

'genuine' one-cell waves made the anesthetised animals a preferred choice to assess

Ca^{2+} buffering properties of astroglia.

In our tests, intracellular Ca^{2+} waves reported by the GCaMP6f fluorescence could be

readily reproduced in an astrocyte model in response to an instantaneous Ca^{2+} rise (5

µM for 0.1 ms) near the soma (Fig. 6d). Although we could observe simulated waves

over a wide range of buffering parameters, the experimentally documented propagation

speed corresponded to only certain values of $[B]$ and K_D (Fig. 6e): for the plausible K_D

range of 0.2, 0.5, 1 and 2 µM the corresponding $[B]$ were 200, 260, 340, and 510 µM,

respectively. The empirical relationship between these $[B]$ and K_D values was almost

perfectly linear (in µM): $[B]$ = 170(1 + K_D). This simple formula captures the buffering

properties of cortical astrocytes, also suggesting the theoretical lower-limit Ca^{2+} buffer

concentration of ~170 µM (ignoring the residual buffering effect of membrane-bound

GCaMP6f).

Decoding astroglial Ca^{2+} signals: from recorded fluorescence to the underlying

Ca^{2+} dynamics

Historically, the relatively slow, global Ca^{2+} elevations (as in Fig. 6c) have been the

most prominent indicator of astroglial activity. Recent advances in high-sensitivity Ca^{2+}

imaging have revealed faster and more local Ca^{2+} signals which are prevalent in small

processes and probably engage compartmentalised and diverse cell functions

(reviewed in [4, 15, 98, 99]). In most cases, such observations rely on high-affinity Ca^{2+}

indicators, thus providing only a crude reference to the underlying Ca^{2+} dynamics [14]. To

translate recorded fluorescence into the intracellular Ca^{2+} dynamics, one has to model

in silico Ca^{2+} entry, diffusion, and the buffering by endogenous proteins and Ca^{2+}

indicators, as shown in numerous studies of nerve and muscle cells (reviewed in [81, 100-

102]).

We therefore asked to what degree Ca^{2+} imaging data commonly recorded in astroglia

in situ shed light on the underlying intracellular Ca^{2+} dynamics. First, we imaged CA1

astrocytes loaded whole-cell with Fluo-4 in acute hippocampal slices: these cells

showed robust spontaneous Ca^{2+} activity consistent with previously reported data [12, 38,

103] (Fig. 7a,b; Supplementary Movie 7).

While the precise origin of astrocyte Ca^{2+} hotspots is still poorly understood, there has

been a large body of work exploring the basic machinery of Ca^{2+} oscillations for other

cell types [80, 81, 104]. The key features emerging from these studies are the occurrence of

detectable microscopic Ca^{2+} domains (0.5-5 μm apart [80, 105]) represented by clusters of

stochastically activated Ca^{2+} channels (such as IP_3 or ryanodine receptors), with the

agonist-dependent mean opening time of 2-20 ms, peak amplitude of ~5 pA, and an

inter-opening interval varying between zero (overlap, multi-channel opening) and ~50

ms [106-108]. Stochastic activation of this system produces local Ca^{2+} 'sparks' (associated

with ryanodine receptors clusters) or 'puffs' (associated with IP_3-receptor clusters) [80],

typically with a puff event frequency of 0.1-2 Hz [30, 109], leading, in certain conditions, to

formation of global Ca^{2+} activity [104]. Genetic deletion of astroglial IP_3 receptors has

unveiled other important sources of spontaneous Ca^{2+} activity, such as mitochondria

and Ca^{2+} channels [82-84]. However, the exact kinetic properties of such Ca^{2+} sources

remain to be established.

In our example, in keeping with the basic paradigm, we have scattered local clusters of

Ca^{2+} channels (IP_3 receptor type) along several branches of the modelled astrocyte, 1-5

μm apart (Fig. 7c), within a ~20 μm area of interest: Ca^{2+} activity outside the area was

of little consequence because of the comprehensive diffusional dissipation in the

absence of global Ca^{2+} events. The simulated astrocyte was 'filled' with 150 μM freely

diffusing Fluo-4 (k_{on} = 600 mM^{-1}ms^{-1}, k_{off} = 21 ms^{-1}) and the endogenous buffer in

accord with our estimation (Fig. 6d, we used 200 μM, K_D = 0.2 μM; other combinations

consistent with Fig. 6d data produced similar results).

Next, we explored simulated channel openings, within and among Ca^{2+} channel

clusters, over the plausible range of their characteristic frequencies (see above;

illustrated in Fig. 7d). Astrocyte Ca^{2+} fluorescence was represented by the

concentration of Ca-bound Fluo-4 and recorded at four arbitrarily selected points in the

area of interest (Fig. 7c; Supplementary Movie 8). We found that at the average interval

between channel openings of ~3 ms within clusters, and ~7 s among clusters, the

simulated fluorescence signals (Fig. 7e) were similar to the experimental recordings

(Fig. 7b). The intracellular Ca^{2+} dynamics underlying these signals was readily available

in the model (Fig. 7f). Clearly, obtaining accurate data on the intracellular Ca^{2+}

dynamics would require systematic experimental probing involving varied Ca^{2+} buffering

conditions and a molecular dissection of the Ca^{2+} cascades involved. However, the

example shown here indicates that the classical molecular machinery of Ca^{2+} signalling

long explored in other cell types [80, 81, 104-106, 108] could generate intracellular Ca^{2+} activity

consistent with Ca^{2+} imaging data collected in astrocytes. Our further exploration of IP_3

arrangement within local astrocyte branches revealed a complex relationship between

inter-cluster distances, spontaneous Ca^{2+} activity and its fluorescent-indicator readout

(see below).

588 Variations in cell features: probing an impact on astroglial function

One of the advantages of having a realistic cell model is the possibility to dissect an

impact of specific cellular features or physiological environment (which might be

inaccessible for experimental manipulation) on cell's behaviour. This strategy could also

reveal whether certain feature combinations make the modelled cell unstable. To

illustrate such an approach, we carried out a selection of example tests in which some

key functional traits of modelled cells were monitored against changes in model

parameters.

First, to see how strongly gross morphology of the astrocyte could influence its

biophysics, we compared two different modelled cells, one with the stem tree

reconstructed from an *in situ* experiment (Fig. 1a-d), and the 'typical CA1 astrocyte',

with the tree branches adjusted to match the average features of CA1 astrocytes (Fig.

1e). The two modelled cells featured different stem trees but were populated with

nanoscopic processes based on the same volume fraction, size, and shape

distributions (as illustrated in Figs 2-3). Simulations revealed only subtle differences

between the cells in their passive voltage spread, input resistance, or Ca^{2+} wave

generation (Supplementary Fig. 9). In a similar context, simulating astrocyte 'swelling'

by ~20% (by evenly increasing the width of cell processes throughout) had only

moderate consequences (Supplementary Fig. 10). These examples suggest a relatively

narrow range of effects arising from morphological variations *per se*, when all other

features remain unchanged.

We next asked how the effective intracellular diffusion coefficient - for instance, for

molecules dialysed in whole-cell mode - would affect the molecule equilibration time in

the astrocyte. Our simulations mimicking whole-cell dialysis through the soma provide a

quantitative illustration of how reducing diffusivity from 0.6 to 0.05 μm^2/ms (values

characterising obstacle-free diffusion of small ions and 2-3kDa molecules, such as

Alexa 488-dextran, respectively) slows down dye equilibration (Supplementary Fig. 11).

These examples, however, should be further tested against experimental data, mainly

because larger molecules tend to undergo significant steric and viscous hindrance

inside small cell compartments (hence diffuse disproportionately slower) [110]. Also, these

example data ignore the possible diffusion escape of the smaller molecules through gap

junctions: this feature could be added through the ASTRO menu (Supplementary

ASTRO User Guide, pp. 27-28) once the gap junction sink parameters have been

experimentally constrained.

Finally, we used simulation settings described in Fig. 7 to ask how the clustering of the

IP_3-dependent Ca^{2+}-signalling mechanisms affects local Ca^{2+} activity in astrocyte

branches. It appears that spreading this signalling mechanism into individual (equally

spaced) clusters, with the same total amount of the IP_3 activity, facilitates triggering of

significant de novo Ca^{2+} events, which could feature prominently in fluorescent

recordings (Supplementary Fig. 12).

DISCUSSION

Realistic astrocyte models: rationale

The present study sought to create a simulation tool ASTRO that would allow scientists

to explore and test mechanistic hypotheses pertinent to complex astroglial physiology,

on the scale from nanoscopic processes to the entire cell expanse. Biophysical cell

models with realistic morphology have been important for a better understanding of

neural function yet there have hitherto been no similar tools available to study astroglia.

We therefore aimed at filling this knowledge gap.

In particular, we tried to show specific examples where the ASTRO built model could

evaluate how fast molecular diffusion exchange occurs within the astrocyte; whether

and how local electric events can control voltage-sensitive membrane mechanisms

across the astrocyte; how fast ions such as K^+ could enter the cell through known

channels and how fast they could equilibrate inside it; what key cellular parameters

control astrocyte Ca^{2+} waves; how one should interpret Ca^{2+} imaging recordings.

Building astrocyte models with ASTRO

The task of recreating astroglial morphology *in silico* included three steps. The first step

was to construct a 'stem tree' based on experimental documentation of main astrocyte

branches that are readily identifiable in the microscope. This procedure is similar to the

commonly used 3D reconstructions of fluorescently labelled nerve cells using *z*-stacks

of their optical sections. The second step, which was the key methodological challenge

here, was to recreate the complex morphology of numerous nanoscopic astroglial

processes that pervade the synaptic neuropil. To address this, we developed algorithms

and computational tools with the aim, firstly, to document and quantify such processes

using an empirical 3D EM database, and secondly, to transform the recorded shapes

into NEURON-compatible, cylinder-compartment-based shapes, with matched

biophysical properties. The latter match was to be verified using specific Monte Carlo

simulation tests (for diffusion and electrodiffusion) incorporated in the software. The

recorded shapes provided all the key statistics characterising the shapes of nanoscopic

astroglial processes in the model.

Thus, the third step was to populate the modelled stem tree with the stochastically

generated nanoscopic processes, in accord with their morphometric characteristics

obtained as outlined above. The essential experimental constraint here was the tissue

volume fraction occupied by astroglial processes, which we and others could measure

directly using either 2PE microscopy or 3D EM. Because neighbouring astrocytes

display a negligible overlap in the tissue volume, this quantity faithfully reflects the

space-filling properties of individual astroglia. We therefore stochastically generated

individual processes on dendritic branches until their bulk, within the cell expanse,

matched the empirically measured astroglial tissue volume fraction. This procedure

completed the architecture of the modelled astrocyte: the model could now be explored

using NEURON simulation environment which we equipped with several additional

functions specific to astroglial exploration.

In our case study, we obtained and incorporated detailed morphological data on

hippocampal astroglia in area CA1, recreated the 'typical' cell *in silico,* and explored it

against experimental observations involving patch-clamp electrophysiology and 2PE

imaging *in situ*. This investigation has shed light on some features of astroglial

physiology that have not been attainable in experimental measurements. Our

simulations predict that membrane local glutamate uptake or K^+ intake via $K_{ir}4.1$

generate only very small membrane depolarisation across the astrocyte. It appears that

transient rises of extracellular [K^+] lead to relatively small relative changes of intra-

astroglial [K^+], which dissipate relatively quickly, within one cell, helped by efficient local

K^+ efflux through $K_{ir}4.1$ channels. Our tests illustrated that relatively small changes in

the Ca^{2+} buffering properties of astroglia might lead to significant changes in the spread

of regenerative intracellular Ca^{2+} signals. The modelling also suggested that the

classical molecular mechanisms of rapid Ca^{2+} sparks and hotspots described in other

cell types could be consistent with the experimental observations of the apparently slow

Ca^{2+} activity documented using common Ca^{2+}-sensitive fluorescent indicators.

687   Exploring astrocyte function and astroglia-synapse exchange

In broad terms, the astrocyte modelling approach presented here serves several

general purposes. Firstly, to recapitulate complex astrocyte morphology at multiple

scales, thus providing a first structurally realistic representation of these cells, as well as

the possibility to explore the effects of morphological changes on cell function.

Secondly, to assess whether experimental results obtained in astroglia, their

interpretation or the underlying hypotheses are biophysically plausible, by reproducing

them in the realistic model. Thirdly, to understand microscopic spatiotemporal profiles of

ion currents, molecular fluxes, and other events that cannot be monitored or registered

experimentally. Finally, to predict the relationships between specific cellular features

(morphology, Ca^{2+} buffering, channel current density, molecular transport, etc.) and

physiological phenotype registered experimentally.

ASTRO has direct access to the full library of biophysical mechanisms pertinent to

synaptic and non-synaptic receptors enabled by NEURON. Therefore, arbitrary patterns

of synaptic influences throughout the astrocyte tree could be simulated. In addition,

excitatory synaptic function could be mimicked using the glutamate uncaging menu in

ASTRO. Clearly, the investigator is supposed to incorporate a specific, experimentally

constrained kinetic mechanism of the synaptically-induced receptor action of interest,

be this IP_3 release, Ca_{2+} entry, K^+ fluxes, etc. Judging by the numerous examples from

neuronal physiology, each such quest will require a separate, dedicated study

combining experimental tests with theoretical probing in the model.

Thus, it is important to stress that our main aim here was not to present one fixed

astrocyte model. Instead, we sought to create an interactive modelling tool ASTRO that

would enable researchers to test biophysical causality of their experimental

observations in astroglial types of interest. The examples presented here provide

illustrations of how such tasks could be accomplished. In words, ASTRO is not a

'frozen' tool: as new investigatory tasks emerge, it is to be upgraded and equipped with

further modelling instruments. The current version of ASTRO and its detailed user's

manual are accessible for download and exploration at

https://github.com/LeonidSavtchenko/Astro.

AUTHORS CONTRIBUTIONS

D.A.R., L.P.S. and C.H. conceived the study; L.P.S. developed and implemented the

modelling approach and its computing environment; C.H. designed and carried out

morphometric studies *ex vivo*; C.H. and L.B. carried out patch-clamp and imaging

experiments and analyses *ex vivo*; T.J. carried out some imaging experiments *ex vivo*,

J.P.R. carried out imaging experiments *in vivo*; I.K., M.M., and M.G.S. designed and

carried our quantitative EM studies; D.A.R. narrated the study, carried out 3D cell

reconstructions, analyses and wrote the paper which was subsequently contributed to

by all the authors.

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 FIGURE LEGEND

Figure 1. Reconstructing astroglial stem tree *in silico*.

a, A characteristic image of CA1 astroglia, whole-cell load with Alexa Fluor 594 (λ_x^{2p} =
800 nm), single optical section (*stratum radiatum*, depth of ~100 μm).

b, Cell as in panel **a** shown as a full *z*-stack projection.

c, Stem tree of astroglia shown in **a-b**, separated and reconstructed in 3D using
NeuroTrace (Fiji ImageJ, NIH); 2D view of a 3D image.

d, Astrocyte stem tree shown in panel **c** quantified, loaded and displayed in NEURON
format using Vaa3D (Allen Institute); thin 'buds' indicate initial seeds for 'planting,
nanoscopic protrusions at a certain longitudinal density; 2D view.

e, Diagram, 'typical' astrocyte stem tree built by modifying a library neurogliaform cell
(2D view); plot, matching the branch diameters in the model (red) and in recorded
astroglia (blue; n = 13 cells including 98 dendrites); solid lines, the best-fit dependence
(power low, $y = a \cdot x^b$) for the corresponding data scatters.

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Figure 2. Documenting and reconstructing *in silico* nanoscopic astroglial
protrusions.

a, 3D EM reconstruction of an astrocytic fragment (green) and adjacent dendritic
spines (grey) with postsynaptic densities (red), from serial sections in area CA1; surface
rendering applied for clarity, as in [10]; dotted square, fragment of interest.

b, Fragment shown in **a** but with the surface point scatter delimiting the shape; false
colour scale, z-depth as indicated; dotted square, fragment of interest (nanoscopic
astroglial process) as in **a.**

c, Astroglial process highlighted in **a-b** shown as a stack of serial polygonal sections
(60 nm thick, to follow EM serial sections) delimited by the surface points.

d, Transformation of the adjacent EM sections (*left;* grey polygonal slabs with base
areas S_i and S_{i+1}) with intersection area T_i (*middle*; green polygons) into NEURON-
compatible two main (grey, 'leaves') and one transitional (green, 'stalk') cylindrical slabs
with the corresponding base areas (*right*).

e, Transformation of the 3D EM reconstructed process fragment (top panel) into
NEURON-compatible cylinder section stacks (*bottom* panel); side view. Individual
sections (blue, *top*) were transformed into 'main' cylinders of the same base area (blue,
bottom). Green segments (*top*) depict adjacent surfaces between sections, which are
represented by green 'transitional' cylinders (*bottom*), as detailed in **d.**

f, A characteristic example of a 3D-EM reconstructed astroglial process represented by
serial polygonal sections (*top*) and its representation by serial cylindrical compartments
(*bottom*). Scattered dots illustrate a snapshot of the Monte Carlo simulation test

(monitored live in ASTRO; Supplementary Movie 1) in which Brownian particles are
injected into the bottom of the 3D structure, and their arrival time at the top is
registered, to compare molecular diffusivity (no electric field) and electrodynamic
properties ($2.5 \cdot 10^3$ V/m electric field in the z-direction, one electron charge $e = 1.6 \cdot 10^{-19}$
C per particle applied) between the two shapes.

g, The outcome of Monte Carlo tests for the two shapes shown in **e**, for the molecular
diffusion flux (top) and ion current (bottom), measured at the top exit of the shapes (as
in **e**), upon injection of the Brownian particles into the bottom entry (as in **e**); blue and
green, 3D-EM-reconstructed and NEURON-compatible shapes, respectively. See [53]
regarding details of electrodiffusion simulations.

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Figure 3. NEURON-based astrocyte model: determining volumetric quantities.

a, Image panel, a characteristic astrocyte in area CA1 (Alexa Fluor 594, $\lambda_x^{2p} = 800$ nm)
seen in a single 2PE optical section (~1 µm thick) crossing the cell soma; dotted lines,
sampling fluorescence intensity profiles reporting the astrocyte tissue volume fraction
(VF). Graph, VF profiles (fluorescence local/soma ratio) obtained along the dotted lines
i and *ii* in the image, as indicated.

b, Matching modelled (red) and experimental (blue; data from n = 13 astrocytes) VF
values (ordinate) sampled at different distances from the soma (abscissa).

c, A complete NEURON-generated astrocyte model (z-projection), with main branches
depicted in orange (partly obscured by smaller processes) and nanoscopic protrusions
(schematic depiction) in purple. Note that tortuous processes of real-life astroglia are
represented by biophysically equivalent 'straightened' processes in the model.

d, An example of astroglia as in **a**; dotted line, line scan position to measure internal
diffusion connectivity (using Alexa Fluor 594 photobleaching); patch pipette is seen.

e, A snapshot of a photobleaching experiment *in silico* showing the intracellular Alexa
concentration dynamics in a modelled astrocyte; arrows, photobleaching line
positioning; false colour scale, intracellular Alexa concentration, as indicated
(Supplementary Movie 2).

f, Matching the modelled (red) and the experimental (blue) time course of intracellular
Alexa Flour fluorescence during a photobleaching experiment as shown in **d-e**, one-cell
example (CA1 area, *stratum radiatum* astrocyte). Grey segments indicate laser shutter-
on when fluorescence recovery occurs (red).

g, Statistical summary of photobleaching experiments (n = 10 astrocytes) and related
simulations, as depicted in **d-e**, comparing experimental (blue) and simulated (red)
data.

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1101 **Figure 4. Electrogenic properties of protoplasmic astroglia.**

a, Traces, a characteristic current-voltage recording of CA1 astroglia; graph, input resistance (bar, mean ± SEM; dots, individual cell data; n = 15).

b, Specific membrane conductance G_m measured in excised whole-cell (outside-out) patches of CA1 astrocytes (bar, mean ± SEM; dots, individual cell data; n = 5).

c, Dots, G_m values obtained by stochastically generating astrocyte models within the empirical range of cell volumes (abscissa) and input resistance matching data shown in **a**; dotted line and grey shade, mean ± SEM for the sample shown; note that NEURON-model astrocyte surface area accounts for both sides and bases of individual cylindrical compartments (Methods).

d, Membrane space constant estimated using a full astrocyte model for centrifugal (left panels) and centripetal (right panels) voltage signal propagation. Cell shape diagrams: V_m landscape snapshots generated by local application (shown by arrow) of a sine voltage signal (amplitude +5 mV). Graphs: signal amplitude attenuation at various signal frequencies, as indicated, for centrifugal and centripetal cases, as indicated.

e, Traces, example of whole-cell recordings (blue) in response to spot-uncaging of glutamate (λ_u^{2p} = 720 nm, 20 ms duration), at two distances from the astrocyte soma, as indicated; red lines, simulated whole-cell current in the corresponding model arrangement (~5 µm -wide glutamate application; GLT-1 kinetics [67]; GLT-1 surface density 10^4 µm^{-2} as estimated [68]). Plot, a summary of glutamate uncaging experiments (blue open dots, n = 8 cells / 90 uncaging spots) and uncaging tests simulated in the model (red solid dots, n = 39).

f, Model snapshot 5 ms post glutamate spot-uncaging depicting the cell membrane current (j, left) voltage (V_m; right) landscape (example in Supplementary Movie 3); false colour scale.

Figure 5. Simulated dynamics of intracellular potassium and membrane voltage triggered by extracellular potassium rise.

a, Cell shape diagrams, time series snapshots of the cell shape (3D-reconstruction reconstruction shown in Fig. 1a-d) illustrating a spherical 20 µm wide area within which extracellular $[K^+]_{out}$ was elevated form baseline 3 mM to 10 mM, for 2 seconds (onset at t = 0), as indicated; $K_{ir}4.1$ channels were evenly distributed with unit conductance of 0.1 mS/cm^2 generating peak current density (in the region with $[K^+]_{out}$ = 10 mM) of 0.01 mA/cm^2. The $K_{ir}4.1$ kinetics were incorporated in NEURON, in accord with [16], as:

$$I_{Kir} = G_{K_0}^* [K^+]_{out}^{1/2} (V_A - V_{KA} - V_{A1}) \left(1 + \exp\left(\frac{V_A - V_{KA} - V_{A2}}{V_{A3}} \right) \right)^{-1} + I_{LA} \text{ where } G_{K_0}^* \text{ is the effective}$$

conductance factor, V_{KA} is the Nernst astrocyte K$^+$ potential, V_A astrocyte membrane potential, K_0 is $[K^+]_{out}$, V_{A1} an equilibrium parameter (sets I_{Kir} to 0 at -80 mV), V_{A2} and V_{A3} are constants calibrated by the I-V curve, and I_{LA} residual leak current.

b, Cell shape diagrams, snapshots illustrating the spatiotemporal dynamics of internal $[K^+]_{in}$ in the test shown in **a;** false colour scale, as indicated.

c, Snapshots illustrating the spatiotemporal dynamics of the membrane voltage in the test shown in **a;** false colour scale, as indicated.

Figure 6. Ca^{2+} waves and Ca^{2+} buffering capacity of astrocytes.

a, Testing the basic effects of mobile Ca^{2+} buffers on Ca^{2+} wave generation in astroglia; cell diagrams, snapshots of $[Ca^{2+}]$ landscape at time points after wave generation, with and without Ca^{2+} buffer, as indicated (note that some branches could obstruct full 3D view); graphs, evolution of the corresponding $[Ca^{2+}]$ profiles (zero Distance, soma centre), as indicated. Relevant model parameters: Ca^{2+} diffusion coefficient, 0.3 μm^2/ms; immobile/endogenous Ca^{2+} buffer concentration, 200 µM (K_f = 1000 mM^{-1} ms^{-1}; K_D = 20 ms^{-1}); mobile Ca^{2+} buffer concentration, 10 µM (K_f = 600 mM^{-1} m^{-1}; K_D = 0.5 ms^{-1}; diffusion coefficient 0.05 μm^2/ms); Ca^{2+} pump activation threshold, 50 nM; Ca^{2+} pump flux density, 20 µM/ms; basal IP_3 concentration, 0.8 µM; IP_3 concentration upon release, 5 µM ; IP_3 release onset time, 1 s (further detail in Supplementary ASTRO User Guide, Supplementary Movie 4).

b, 2PE image of the rat somatosensory cortex *in vivo* (~100 µm deep); a single optical section shown with bulk-loaded sulforhodamine 101 (SR101, red channel) to label astroglial structures, as detailed earlier [91]; *AB*, examples of the astrocyte cell body; *AE*, examples of astrocyte endoot processes surrounding blood vessels; see Supplementary Movies 1 and 2.

c, Region of interest (as in **b**) shown in the $[Ca^{2+}]$-sensitive GGaMP6f (green) channel. *Top*, snapshot sequence (Supplementary Movie 4) depicting generation and spread of an intracellular Ca^{2+} wave (dotted circle); *bottom*, same sequence shown as the time derivative of fluorescence intensity (over 50 ms interval), to highlight the Ca^{2+} wave front; false colour scale. See Supplementary Movie 5 for awake-animal recording example.

d, Cell shape diagrams, snapshots of the simulated Ca^{2+} wave spreading with the speed that matches experimental observations; false colour scale (*C*, concentration). Plot, the average intracellular $[Ca^{2+}]$ profile depicting the corresponding (centrifugal) Ca^{2+} wave propagation; δx illustrates the wave front speed measurement (distance travelled over 0.5 s).

e, Summary of simulations depicted in **c**, to determine the cytosol Ca^{2+} buffering capacity (combination of the buffer affinity K_d and its concentration) that would explain the observed Ca^{2+} wave speed; $[IP_3]$, assumed intracellular concentration of IP_3 [88-90]; dotted line, average experimental speed of centrifugal Ca^{2+} waves in astroglia *in vivo*, determined as in (**a-b**) (n = 54 events in approximately 20 individual cells).

Figure 7. Deciphering intracellular free [Ca^{2+}] dynamics from fluorescence Ca^{2+} imaging in astroglia *in situ*.

a, Example, astrocyte (CA1 area, acute hippocampal slice; Fluo 4 channel, λ_x^{2P} = 800 nm) held in whole-cell, with regions of interest for Ca^{2+} monitoring (circles, ROIs 1-7; Supplementary Movie 7).

b, Time course of Ca^{2+} sensitive fluorescence (Fluo-4 channel) recorded in ROIs 1-7 as in **a**, over 100 s; same colours correspond to ROIs on the same branch (ROIs 2-3 and 4-5).

c, An astrocyte model with localised Ca^{2+}-puff sources (orange dots) and four recording points (arrows, 1-4); dotted oval, region for analyses: cell area outside has a negligible effect of the Ca^{2+} sources as shown (Supplementary Movie 8).

d, Example of channel-like local Ca^{2+} entry activity generated by a single localized Ca^{2+} source, in accord with the known biophysical properties of cellular Ca^{2+} sparks and hotspots [30, 104, 109].

e, Time course of simulated Fluo-4 fluorescence (150 µM 'added') in ROIs 1-4 shown in **c**: it has statistical properties similar to those recorded *in situ* (**b**); shaded area, time window for higher temporal resolution (see **f**); same line colours correspond to ROIs on the same cell branch.

f, *Right*, simulated intracellular [Ca^{2+}] dynamics underlying Fluo-4 fluorescence shown in **e**; *left*, trace fragments on the expanded time scale (shaded area in **e**), as indicated; the fragments correspond to the period of relatively high [Ca^{2+}].

 METHODS

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Experimental methods: Electrophysiology *ex vivo*

Acute hippocampal transverse slices (350 μm thick) were prepared from P21-28
Sprague-Dawley rats, in full compliance with the national guidelines, the European
Communities Council Directive of 24 November 1986 and the European Directive
2010/63/EU on the Protection of Animals used for Scientific Purposes. Slices were
prepared in an ice-cold slicing solution containing (in mM): NaCl 60, sucrose 105,
NaHCO3 26, KCl 2.5, NaH2PO4 1.25, MgCl2 7, CaCl2 0.5, glucose 11, ascorbic acid
1.3 and sodium pyruvate 3 (osmolarity 300–310 mOsM), stored in the slicing solution at
34 ºC for 15 min and transferred for storage in an extracellular solution containing (in
mM): NaCl 125, NaHCO3 26, KCl 2.5, NaH2PO4 1.25, MgSO4 1.3, CaCl2 2 and
glucose 16 (osmolarity 300-305 mOsm). All solutions were continuously bubbled with
95% O2/5% CO2. Slices were allowed to rest for at least 60 min before recordings
started.

Whole-cell patch-clamp recordings of *stratum radiatum* astroglia were performed in a
submersion-type recording chamber. Slices were superfused with an extracellular
solution containing (in mM): NaCl 125, NaHCO3 26, KCl 2.5, NaH2PO4 1.25, MgSO4
1.3, CaCl2 2 and glucose 16 (osmolarity 300-305 mOsm), continuously bubbled with
95% O2/5% CO2. Whole-cell recordings were obtained with patch pipettes (3-5 MΩ)
with an intracellular solution containing (in mM): KCH3O3S 135, HEPES 10, Tris-
phosphocreatine 10, MgCl2 4, Na2ATP 4, Na3GTP 0.4 (pH adjusted to 7.2 with KOH,
osmolarity 290–295 mOsM). The cell impermeable Ca2+ indicator OGB-1 (200 μM
unless indicated otherwise; Invitrogen O6806) was added to the internal solution. CA1
pyramidal cells were held at -70mV. Protoplasmic astrocytes located in the stratum
radiatum were identified by their small soma size, low resting potential (< -80 mV) and
low input resistance (<10MΩ). Astrocytes were held in voltage clamp at their resting
potential or in current clamp.

In some experiments, whole-cell patches were excised by pulling gently the patch
pipette attached to the cell body until the patch was completely detached from the
processes and its membrane sealed. Estimation of the patch capacitance c was carried
out using a classical voltage-step method in which (a) a brief voltage step ΔV is applied,
(b) the area under the transient capacitance current after the end of the voltage step is
measured, giving electric charge Q, and (c) membrane patch capacitance is estimated
as $c = Q / \Delta V.$ Thus the patch area is evaluated from the ratio c / c_m where specific
membrane capacitance $c_m = 1$ μF / cm^2 is a common characteristic of astroglial
membranes [18]. The steady-state current response to the voltage step was used to
calculate the patch conductance, which was then normalized to the membrane area to
obtain G$_m$.

1243

Experimental methods: 2PE imaging, uncaging, and FRAP *ex vivo*

Astrocytes were filled via whole-cell patch clamp with 40-100 µM Alexa Fluor 594 for
15-20 min. We used an Olympus FV1000 imaging system optically linked to a
femtosecond pulse Ti-sapphire MaiTai laser (Newport Spectra-physics). Cells were
imaged using an Olympus XLPlan N 25x water immersion objective. Fluorescence
recordings were obtained in line-scan mode (500Hz, line placed through the astrocyte
arbour and across the soma) at λ= 800 nm at an increased laser power of 15-20 mW
under the objective to induce substantial bleaching of Alexa Fluor 594. Fluorescence
was collected for 750-1000 ms, then a mechanical shutter was placed in front of the
laser beam for 1-2 s to allow fluorescence recovery.

We used a combined two-photon uncaging and imaging microscope (Olympus, FV-
1000MPE) powered by two Ti:Sapphire pulsed lasers (Chameleon, Coherent, tuned to
720 nm for uncaging and MaiTai, Spectra Physics, tuned to 800 nm for imaging). Cells
were imaged using an Olympus XLPlan N 25x water immersion objective. The intensity
of the imaging and uncaging laser beams under the objective was set to 5 mW and 12-
17 mW, respectively.

To record spontaneous Ca2+ transients in frame scan mode, 200 µM Fluo-4
(Invitrogen) and 100 µM Alexa Fluor 594 (Invitrogen) were added to the intracellular
solution. 350-500 µm^2 fields of view where imaged within the arbour of the patched
astrocyte and the fluorescence emitted by Alexa Fluor 594 and Fluo-4 was collected at
a rate of 3-5 Hz for 2-3 min. Time-dependent fluorescence transients were expressed
as ΔG/R where G corresponds to the background-subtracted Fluo-4 fluorescence and
R to the background-subtracted Alexa Fluor 594 fluorescence.

For MNI-glutamate uncaging, astrocytes were loaded with 100 µM Alexa Fluor 594 as a
morphological marker. Astrocytes were held in voltage-clamp mode at their resting
membrane potential (typically between -80 and -90 mV). The MNI-glutamate (12.5 mM)
was either puffed within the tissue from a glass pipette placed above the patched cell,
or added to the bath at 2.5 mM. Glutamate was uncaged for 20 ms at different
distances from the soma (5-25 µm).

1273

1274 Experimental methods: 3D reconstruction of live astrocyte stem tree

A *stratum radiatum* astrocyte was held in whole-cell mode, with Alexa Fluor 594 added
to the intracellular solution (see above; excitation at λ_x^{2p} = 800 nm). A z-stack of 2PE
images was collected using 100 x 100 µm (512 x 512 pixel) individual frames containing
the entire visible astrocyte structure, with a 0.5 µm z-step over 61 µm in depth. The
image stack was stored (8-bit tiff format), individual images were corrected for the
depth-dependent, quasi-exponential fluorescence signal decrease (Fiji Image-Adjust-
Bleach Correction, plugin by Kota Miura 2014: 10.5281/zenodo.30769). Fluorescence
background was subtracted (Fiji Image-Process), identifiable cell branches (> 0.3-0.5
µm in diameter) were traced semi-automatically in individual 2D optical sections and
reconstructed in 3D using Neurite Tracer (Fiji Plugins-Segmentation-Simple Neurite
Tracer; by Mark Longhair and Tiago Ferreira, MRC and Janelia Campus;

http://imagej.net/Simple_Neurite_Tracer; default segmentation sigma, 0.196). The data
sets representing diameters of tubular compartments and their 3D co-ordinates (pairs of
end points) were stores in SWC format. The Vaa3D software (Allen Institute,
http://www.alleninstitute.org/what-we-do/brain-science/research/products-tools/vaa3d/)
was used to convert these data sets into NEURON compatible files providing 3D
structure of the astroglia stem-tree (with tubular compartments representing individual
cylindrical compartments).

1293

Experimental methods: Astroglia-targeted expression of GCaMP6f *in vivo*

Animal procedures were conducted in accordance with the European Commission
Directive (86/609/ EEC) and the United Kingdom Home Office (Scientific Procedures)
Act (1986). Young male C57BL/6 mice (2 – 3 weeks of age) were anaesthetised using
isoflurane (5% induction, 1.5 - 2.5% v/v). Subcutaneous analgesic (buprenorphine, 60
μg kg^{-1}) was administered and the animal was secured in a stereotaxic frame (David
Kopf Instruments, CA, USA) and kept warm on a heating blanket. The scalp was
shaved and disinfected using three washes of topical chlorhexidine. Upon loss of pedal
withdrawal reflexes, a small midline incision was made to expose the skull. A
craniotomy of approximately 1 - 2 mm diameter was performed over the right
somatosensory cortical region using a high-speed hand drill (Proxxon, Föhren,
Germany). Stereotactic coordinates were +0.1 mm on the anterioposterior axis relative
to bregma, and 2 mm lateral to midline. Once exposed, a warmed aCSF variant (cortex
buffer, in mM; 125 NaCl, 2.5 KCl, 10 HEPES, 10 glucose, 2 CaCl$_2$, 2 MgSO$_4$) was
applied to the skull and cortical surface throughout the procedure.

AAV5 *GfaABC1D-LckGCaMP6f* (Penn Vector Core, PA, USA) was pressure injected
into the somatosensory cortex using a pulled glass micropipette stereotactically guided
to a depth of 0.6 mm beneath the pial surface, at a rate of approximately 1 nL sec^{-1}. A
given injection bolus contained between 0.25 and 0.5 x 10^{10} genomic copies, in a
volume not exceeding 500 nL. After injection, pipettes were left in place for 5 minutes
before retraction. The scalp was sutured with absorbable 7-0 sutures (Ethicon Endo-
Surgery GmbH, Norderstedt, Germany) and the animal was left to recover in a heated
chamber. Meloxicam (subcutaneous, 1 mg kg^{-1}) was administered once daily for up to
two days following surgery. After a 4 - 6 week AAV incubation period, animals were
prepared for multiphoton imaging through a cranial window implantation as described
below.

1320

Experimental methods: Two-photon excitation imaging of astroglia *in vivo*

Following viral transduction of LckGCaMP6f as above, male C57BL/6 mice (7-9 weeks
of age) were prepared for cranial window implantation and 2PE microscopy. Animals
were anesthetized using fentanyl, midazolam and medetomidine (i.p., 0.05, 5 and 0.5
mg kg^{-1}, respectively). Adequate anesthesia was ensured by continuously checking for
the loss of pedal withdrawal reflexes and anaesthesia was supplemented appropriately

throughout the procedure (typically 10-20 % of the original dose per hour). Body
temperature was maintained at 37.0 ± 0.5 °C using a feedback rectal thermometer and
heating blanket. The animal was secured in a stereotaxic frame and a craniotomy of
approximately 2.5 mm diameter was carried out over the right somatosensory cortex,
centred 0.2 mm caudal to bregma and approximately 2.5 mm laterally from the midline.
Once exposed, the cortical surface was continuously superfused with warmed aCSF (in
mM; 125 NaCl, 2.5 KCl, 26 $NaHCO_3$, 1.25 Na_2HPO_4,18 Glucose, 2 $CaCl_2$, 2 $MgSO_4$;
saturated with 95% O_2 / 5% CO_2, pH 7.4). Cortical astrocytes were labeled using
multicell bolus loading of sulforhodamine 101 (SR101, 5μM). SR101 (in cortex buffer
vehicle) was pressure-injected through a pulled glass micropipette targeted to 2 - 3
injection sites within the transduced region, comprising a total volume of 500nL. The
cortical surface was covered with 1% agarose and a glass coverslip was placed on top.
Using tissue adhesive (Dermafuse, Vet-Tech Solutions, UK), the coverslip was partially
secured and a custom-built headplate fixed to the skull. A single cranial-mounted screw
was inserted over the contralateral hemisphere and the entire assembly was then
secured using dental cement. During imaging, the headplate was used to secure the
animal under the objective on a custom-built stage.

In these experiments, two-photon excitation was carried out using a Newport-
Spectraphysics Ti:sapphire MaiTai laser pulsing at 80 MHz, and an Olympus FV1000
with XLPlan N 25x water immersion objective (NA 1.05). Acquisitions were carried out
using a wavelength of 920 nm and the mean laser power under the objective was kept
at 20 – 35 mW. Cortical astrocytes were readily identified through SR101 labeling and
verified for GCaMP6f expression by frame-scanning for calcium transient activity.
Recordings were made at a depth between 50 and 250 μm from the cortical surface. XY
time series (at 0.5 – 2 Hz with a pixel dwell time of 0.5 - 4 μs and pixel size of 0.248 –
1.59 μm) were taken in identified regions to measure spontaneous calcium activity.

1353

Experimental methods: Fast fixation and DAB staining of recorded astrocytes

In a subset of experiments we loaded an astrocyte with biocytin, and after the
experiment the slices were rapidly fixed (by submersion) with 1.25% glutaraldehyde and
2.5% paraformaldehyde in 0.1 M PB (phosphate buffer, pH 7.4), to be kept overnight,
submerged in 10% sucrose in PB for 10 min and then in 20 % sucrose in PB for 30 min.
The slices were consequentially freeze-thaw in liquid freon and liquid nitrogen for 3 sec
each to gently crack intracellular membranes and embedded in 1% low gelling
temperature agarose in PB (Sigma-Aldrich, USA). Embedded slices were sectioned at
50 μm on a vibrating microtome (VT1000; Leica, Milton Keynes, UK). 50 μm sections
were incubated in 1% H_2O_2 in PB for 20 min to eliminate blood background, washed
with 0.1 M TBS (tris buffer saline, pH 7.4) and incubated with ABC solution
(VECTASTAIN ABC, Vector laboratories, USA) for 30 min at room temperature. Next
section were washed with 0.1M TB (tris buffer, pH 7.4), pre-incubated with DAB (3,3'-
Diaminobenzidine tablets - Sigma-Aldrich, USA) solution (10 mg DAB tablet + 40 ml TB)
for 30 min at room temperature in dark and finally incubated with DAB+ H2O2 solution

(5 µl of 33% H_2O_2 + 25 ml of DAB solution) for 10-20 min at room temperature in dark.
The DAB stained sections was washed in PB, post-fixed in 2% osmium tetroxide and
further processing and embedding protocols were essentially similar to those reported
previously (Medvedev et al., 2010). Briefly, the tissue was dehydrated in graded
aqueous solutions of ethanol (40 - 100%) followed by 3 times in 100% acetone,
embedded into a mixture of 50% epoxy resin (Epon 812 ⁄ Araldite M) and 50% acetone
for 30 min at room temperature, embedded in pure epoxy resin, and polymerized
overnight at 80°C. Sections in blocks were coded and all further analyses were carried
out blind as to the experimental status of the tissue.

1378

Experimental methods: 3D electron microscopy

Serial sections (60–70 nm thick) were cut with a Diatome diamond knife as detailed and
illustrated earlier [111, 112], and systematically collected using Pioloform-coated slot copper
grids (each series consist of up to 100 serial sections). Sections were counterstained
with 4% uranyl acetate, followed by lead citrate. Finally sections were imaged in *stratum*
radiatum area of CA1 (hippocampus) using AMT XR60 12 megapixel camera in JEOL
1400 electron microscope. Serial sections were aligned as JPEG images using SEM
align 1.26b (software available from http://synapses.clm.utexas.edu/). 3D
reconstructions of DAB stained astrocyte fragments and the adjoined to stained
astrocytes dendritic spines (that host clearly identifiable excitatory synapses) were
performed in Trace 1.6b software (http://synapses.clm.utexas.edu/). 3D reconstructions
of selected astrocytic segments and dendritic spines were imported to 3D-Studio-Max 8
software for rendering of the reconstructed structures.

1392

Statistics summary

The present study contained no longitudinal or multifactorial experimental designs. In
electrophysiological or imaging experiments the main source of biological variance in
were either individual cells or individual preparations (the latter in case of field
measurements in acute slices), as indicated. In accord with established practice, in the
ex vivo tests we routinely used one cell per slice per animal, which thus constituted
equivalent statistical units in the context of sampling, unless indicated otherwise.
Statistical hypotheses pertinent to mean comparisons were tested using a standard
two-tailed t-test, unless the sample showed a significant deviation from Normality, in
which case non-parametric tests were used as indicated. The null-hypothesis rejection-
level was set at $\alpha = 0.05$, and the statistical power was monitored to ensure that that the
sample size and variance were adequate to detect a mean difference (in two-sample
comparisons) of 10-15% or less.

1406

Astrocyte model: generating 'invisible' nanoscopic morphology

Nanoscopic processes of the astrocyte model are generated in a probabilistic manner
based on the sample statistics from 3D EM reconstructions (Fig. 2). The total cell
surface area S_{tot} represented by the cylinder-based shape approximations (Fig. 2d,e),
consists of the (lateral) surface areas of all cylinder-compartment sides S_{lat} added to the
surface areas of 'main' cylinder bases S_M (blue in Fig. 2d, bottom) minus the surface
areas of 'transitional' cylinder bases S_T (green in Fig. 2d, bottom). In our case study,
computations indicated that $S_T = 0.20 S_M$ throughout modelling: thus, the formula $S_{tot} =$
$S_{lat} + 0.8 S_M$ was applied. In the generated cell models, the S/V ratios were ranging from
~7 μm^{-1} near the soma to an average of ~22 μm^{-1} in the bulk of the cell arbour, in
accord with the empirical observations.

1418

1419 Astrocyte model: transporter/channel kinetics and diffusion-reaction
1420 mechanisms

Models built with ASTRO can incorporate many dozens of NEURON-enabled channel
and transporter kinetic mechanisms that have been tested and validated in numerous
studies combining experiments and simulations [24]. The formal descriptions of the
respective algorithms could be found using an extensive NEURON database at
https://senselab.med.yale.edu/modeldb/ which also contains references and links to the
original studies and the mathematical formulism involved. Upon ASTRO installation on
the host computer, these mechanisms could also be inspected in the respective *.mod
files in the 'neuronsims' directory or, alternatively, online here
https://github.com/LeonidSavtchenko/Astro/tree/master/neuronSims.

Several channel current and diffusion-reaction mechanisms have been written
specifically for the present model. The kinetics of glutamate transporter GLT-1 involving
glutamate and ion fluxes has been incorporated in accordance with [113] (description in
the GluTrans.mod file). The $K_{ir}4.1$ potassium current has been incorporated in
accordance with [16] (Kir4.mod; Fig. 5 legend), intracellular K^+ diffusion was incorporated
as longitudinal diffusion (no radial rings) using a built-in Ca^{2+} diffusion algorithm
described in the next section (potassium.mod), the FRAP mechanism incorporated the
same algorithm plus a reaction-diffusion step (FRAP.mod), and K^+ extrusion was
modelled as a first-order pump (kpump.mod; Supplementary Fig. xx legend). Gap
junction mechanisms were enabled either as a (zero-order) current leak (gap.mod) or
as a diffuse escape (gapCa.mod). Throughout these mechanisms, the respective
kinetic parameters can be set using the relevant NEURON-enabled ASTRO menus, as
described in the User Guide
(https://github.com/LeonidSavtchenko/Astro/blob/master/ASTRO_User_Guide.pdf).

1444

1445 Astrocyte model: Ca^{2+} homeostasis and diffusion

ASTRO simulation algorithms enabling intracellular Ca^{2+} homeostasis and diffusion
(including that among adjacent compartments of unequal size) are detailed in Chapter 9
of the NEURON Book [24] (also here https://www.neuron.yale.edu/neuron/docs), and can

be found in the modified cadifus.mod file in the model installation. In brief, Ca^{2+} diffuses
freely whereas buffer-bound Ca^{2+} (which has much lower diffusivity) is considered
stationary, for the sake of simplicity. In individual cylindrical cell compartments, radial
diffusion occurs through four concentric shells surrounding a cylindrical central core,
and longitudinal diffusion is calculated using fluxes between the corresponding
concentric compartments adjusted for the cross-section areas. The longitudinal and
radial diffusion coefficient for Ca^{2+} was set to 0.3 μm^2 ms^{-1}, the basal level was set to 50
nM, and IP_3 concentration at 0.

In addition to free diffusion, Ca^{2+} homeostasis mechanisms included the SERCA pump,
SERCA channel and SERCA leak, the endogenous (stationary) and exogenous (Ca^{2+}
indicator) mobile buffers, and a plasma membrane Ca^{2+} pump with the threshold
mechanism (cadifus.mod). The kinetics of buffers can be modified using NEURON-
enabled menus. The mechanistic details of Ca^{2+} SERCA pump were as described
earlier [87, 114]. The current model implementation assumes that IP_3 is distributed
uniformly across cell compartments, i.e. that diffusion equilibration of IP_3 is fast
compared to Ca^{2+} concentration transients in space or time.

1465

Modelling with ASTRO: On-line access and installation

Detailed information on the installation and running of ASTRO can be found in the
User's Manual (Supplementary Material; online download at
https://github.com/LeonidSavtchenko/Astro/blob/master/Manual). The current version of
ASTRO can also be downloaded directly from
https://github.com/LeonidSavtchenko/Astro. The (regularly updated) User Guide can be
downloaded from the same location or found (current version) in the Supplementary
Information (User_Guide.doc).

In brief, running ASTRO without full-scale simulations of intracellular Ca^{2+} dynamics
requires the host computer to have MATLAB (2012 version or later,
https://uk.mathworks.com/products/matlab.html) and NEURON (7.2 or later,
https://neuron.yale.edu/neuron/download) installed under Windows 7 or Windows 10.

Simulating full intracellular Ca^{2+} dynamics is highly resource-consuming and normally
requires an additional Worker computer / cluster operating under Linux, with
preinstalled NEURON (https://neuron.yale.edu/neuron/download/compile_linux) and
MPI whereas the Host computer will require MATLAB (2013 version or later), NEURON
(7.0 or later), and access to the Internet. In house, the Linux version with the parallel
computations including intracellular diffusion simulation was routinely run using a 12-
node in-house computer cluster [115], taking advantage of the computational optimization
routines developed by us earlier for compartmental models and Monte Carlo
simulations [52, 53, 115].

1487

1488

Figure 1

a 3D EM

1 μm³

b 3D EM *in silico*

Z (μm)

Y (μm)

X (μm)

c nano-process

500 nm

d 3D EM → *intersection* → NEURON

e 3D EM

NEURON-ready

f 3D EM

NEURON-ready

g 3D EM vs N-ready

Diffusion

Diffusion flux (mol ms⁻¹)

Time (ms)

Current

Current, 10³ e ms⁻¹

Time (ms)

Figure 2

Figure 3

Figure 4

Figure 5

Figure 6

Figure 7