1 Pleistocene Mammal Population Fluctuation Patterns Inferred by Their Genomes

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15 Summary

16	Paleoclimate fluctuations critically affect paleoecological systems and influence mammal
17	populations, even resulting in population differentiation [1]. Historical effective population
18	size (N_e) can reflect these influences [2, 3]. Dozens of recent studies have investigated
19	the relationship between variations in N_e values of one or a small number of mammalian
20	species, inferred from genomic data, and fluctuations in paleoclimate [4-7]. However,
21	there lacks an integrated and comprehensive study on the relationship between the
22	fluctuations in paleoclimate and variations in N_e values inferred from genome sequencing
23	data of a wide range of mammals. To investigate patterns in mammalian N_e values during
24	the the Pleistocene, we gathered whole genome sequencing data of 60 mammals from
25	35 species distributed across Afro-Eurasia and the Americas, then inferred their N_e
26	curves using the Pairwise Sequentially Markovian Coalescent (PSMC) method; 30
27	mammalian N_e curves almost simultaneously started to contract at the turning point of the
28	Middle Pleistocene Transition (MPT); then the population of seven mammals started to
29	expand at the turning point of the Middle Brunhes Event (MBE), while the contraction of
30	other mammals' populations was prolonged to the later different time periods. Eight
31	mammals experienced a severe population contraction around the Last Glaciation
32	Maximum, as some aves did [8], while four potential ruminant beneficiaries showed an
33	expanding population. Sus scrofa and Bos taurus experienced an internal population
34	differentiation in the MPT. To conclude, the phenomenon that critical paleoclimate events

- 35 facilitated contemporaneous animal population fluctuations in the paleoecological system
- 36 is showed by our N_e curve analysis.

38 Keywords: PSMC, mammal Ne, PCA, MPT

40 **RESULTS AND DISCUSSION**

41

42 Effective population size of mammals during the Pleistocene

- By utilizing the nonuniform distribution of single nucleotide variants (SNVs) in the genome,
 the Pairwise Sequentially Markovian Coalescent (PSMC) method can reconstruct the
 historical effective population size (*N_e*) from any single individual of the population [4]. In
 this work, we applied the PSMC method to obtain the *N_e* curves of 35 mammal species in
 Afro-Eurasia and the Americas (Figure S1). The *N_e* curves estimated from the genome
- 48 data span the whole Pleistocene, yielding us that were sufficient to conduct the
- 49 subsequent analysis.
- 50

51 Climates contributed to the mammal N_e curves

- 52 *Ne* fluctuation could result from a combined effect of genetics, climate, locations, or other
- 53 factors [8]. To understand the degree that climate contributed to the N_e curves, we
- 54 extracted the principal components (PCs) from 60 N_e curves by the principal component
- analysis (PCA). PC1 and PC2 contributed variance ratios of 0.508 and 0.273,
- 56 respectively (Figure 1A). PC1 accounts for most of the variance; however, after
- 57 hierarchical clustering, no signals could be detected to support that PC1 is correlated with
- 58 genetic background (Figure S2). This finding could suggest that the complex
- 59 paleoclimate impact on the mammals. We clustered these mammals hierarchically with

60	PC2 (Figure 1B) and discovered that in many mammals (Ursus, Ponginae, Homo, Canis,
61	some subgroups in Rhinopithecus and some herbivorous mammals), their genetically
62	close subgroups were clustered in the same branch, with the exception of Bos and Sus.
63	Eurasian and African Bos clusters as well as Asian and European Sus clusters are
64	separated by a clear boundary, reflecting a long separation time of their populations that
65	will be elaborated in the following section. Hence, PC2 could be interpreted as the
66	integrated effect of the inherent genetic characteristics and geographical locations on the
67	mammals. In general, close locations imply similar climate conditions; therefore,
68	fluctuations of mammal N_e curves can reflect the climates of the mammal ancestors
69	confronting to some extent.
70	
70 71	<i>N</i> _e curve extrema clustered into seven clusters
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 71 72 73 74 75 	N_{e} curve extrema are consistent with significant climate events in the Pleistocene. These extrema (Table S1) are centered at seven time points with an average silhouette coefficient of 0.64,an acceptable cluster indicator. All seven clusters coincide with turning points of important climate events, as elaborated below.
 71 72 73 74 75 76 	N _e curve extrema are consistent with significant climate events in the Pleistocene. These extrema (Table S1) are centered at seven time points with an average silhouette coefficient of 0.64,an acceptable cluster indicator. All seven clusters coincide with turning points of important climate events, as elaborated below. The earliest cluster centered at 0.97±0.021 mya, with 40 extrema, coincided with

80	roxellana, Rhinopithecus strykeri, Ursus arctos, Ursus maritimus, Pan troglodytes verus
81	and Gorilla gorilla diehli; and ten N_e curves displayed a bottleneck, including Homo and
82	five species from Equus. Accordingly, we categorized the mammals into three classes
83	(Table S1): MPT Maxima Class, MPT Minima Class, and MPT Neutral Class (mammals
84	whose N_e curves are without the first extremum in the MPT class).
85	The contraction of the MPT Maxima Class around 0.97 ± 0.021 mya may result from
86	the Jaramillo geomagnetic reversal (1.07 mya~0.99 mya) [10] or changes in solar
87	insolation after the coeval $\delta^{13}C_{max-III}$ event of 1.0~0.95 mya (Marine isotope stage (MIS)
88	29–25) [11, 12]. The climate became too harsh at a large scale for mammals to survive [9]
89	and led to massive mammal extinction or population contraction events such as
90	warm-steppic mammals being replaced with cold-steppic mammals in European
91	Villafrancian fauna [13] and the southern-migration mammal wreck event of
92	Gongwangling fauna [14]. The Plio-Pleistocene contributed substantially to current
93	biomes [15] and MP fossils of 73.3% mammals in this class, including Ailuropoda [16],
94	Bos [17], Capra [18], Equus [19], Rhinopithecus [20] and Sus [21] are majorly distributed
95	in the Palearctic, according to the fossil database (Figure S3A, B, D, E, K and L).
96	Therefore, the population peak of the MPT Maxima Class in the MPT Cluster suggests
97	that the ancient populations of MP mammals in the Palearctic could have experienced
98	severe contractions around the turning point of the MPT.

99	The MPT Maxima Class population started to contract from the MPT Cluster center,
100	and their first bottlenecks after that are distributed in five clusters (Figure 3A). The MPT
101	Maxima Class, including African Pan troglodytes verus and Bos, wild-spread Equus and
102	European Sus Scrofa, reached the first bottleneck at 0.43±0.018 mya, which is referred to
103	as the MBE (Middle Brunhes Event) Bottleneck for 0.43 mya is happens to be the turning
104	point of the MBE (MIS13-MIS11) [22]. Eight mammals, including Capara, Asian Sus
105	scrofa, Ovis aries, and Bos indicus, reached the first bottleneck of their population at
106	approximately 0.32±0.01 mya, the time which was known during the Mindel-Riss
107	interglacial [23] (MRI bottleneck). These mammals, except the Bos indicus, were sampled
108	at a relatively high latitude. In the Mindel-Riss interglacial, the climate was similar to the
109	current climate. For four mammals including Ursus arctos, Ursus maritimus and two Bos
110	taurus individuals, their ancestral N_e values synchronously reached the first bottleneck at
111	0.22±0.010 mya in the middle of the Riss glaciation, which is a relatively weak MIS 7
112	period with a cool interglaciation climate [24] (MRG Bottleneck). For nine mammals from
113	the MPT Maxima Class, including Ovis ammon polii, Ailuropoda melanoleuca, Bos
114	grunniens, Rhinopithecus roxellana and Sus celebensis as well as three Bos taurus
115	individuals and one Sus scrofa individual, their ancestral N_e values started to show an
116	expansion of their populations until 0.12±0.008 mya in the warmest, last interglaciation,
117	the Eemian interglaciation [25] (EI Bottleneck). Two species, Gorilla gorilla diehli and
118	Rhinopithecus strykeri, reached the first bottleneck until 0.04 ± 0.004 mya (the most recent

cluster center). These mammal populations contracted almost simultaneously but startedto expand at different time period.

121 Climate conditions could prompt mammal population migration and differentiation. 122 As a result, the population of migrated mammals may have started to expand, while 123 mammals that did not migrate sustained contraction. In the MPT Maxima Class, the 124 mammals that experienced the MBE Bottleneck likely lived in low latitude regions, and 125 those that experienced the MRI Bottleneck usually resided at relatively high latitudes with 126 the exception of Bos and Sus scrofa, which will be elaborated in the following section. 127 However, there was no obvious association between population fluctuations and habitats of the mammals that experienced the MGR and EI bottlenecks. The fact that mammals in 128 129 the MPT Maxima Class spread in different latitudes, adapted themselves to different 130 environments and developed different habits, but shared five asynchronous expansion 131 times, could result from combined factors of environment in different latitudes and the capacity for mammalian adaptations. 132

We also noticed that ten mammals of the MPT Neutral Class (Figure 3C) reached a bottleneck while three of them reached a peak, at 0.67±0.030 mya, or the end of the MPT. At that time, the planet held a high ice volume but was relatively warm [24, 27]. The mammals that underwent a bottleneck around 0.67±0.030 mya included *Pongo pygmaeus*, *Ursus americanus Macaca mulatta lasiota*, *Macaca thibetana* and *Canis*. Their populations began to shrink separately before 1 mya, indicating that the population

139	contraction impetus differed for mammals from the MPT Maxima Class. The mammals
140	with a peak at that time included Odocoileus virginianus texanus, Gorilla gorilla gorilla and
141	Gorilla beringei graueri.
142	
143	Differentiation in mammal populations
144	Sus scrofa and Bos populations experienced internal population differentiation between
145	0.97 mya and 0.43 mya in the middle Pleistocene transmission.
146	Sus scrofa belongs to the MPT Maxima Class (Table S1); thereafter, the
147	population of European Sus scrofa experienced a prolonged contraction until the MBE
148	Bottleneck, while the population of Asian Sus scrofa continued contracting until the MRG
149	Bottleneck (Figure 4A). The middle Pleistocene fossils of Asian Sus scrofa were only
150	distributed in relatively low latitude areas in Asia (Figure S3L). We also noticed that some
151	Asian mammals, such as Capra aegagrus, Capra hircus [18], and Ailuropoda
152	melanoleuca [16], also started their population expansion at the MRG Bottleneck, while
153	the European Sus scrofa population started to expand at the turning point of the MBE,
154	which is the same time as the Equus caballus population did [19]. The long-term
155	population contraction of Eurasian Sus scrofa, following a different fluctuation trend in the
156	MBE Bottleneck, indicates a differentiation in Eurasian Sus scrofa during the MPT and
157	MBE. Kijas estimated that Eurasian Sus scrofa maternity divergence time is ~0.9 mya
158	according to the mtDNA genome sequence [32]; with synonymous and nonsynonymous

159	nucleotide substitutions of two main domains of the D-loop region in the cytochrome B
160	gene (CytB) (1,847 bp), Alves estimated the Eurasian Sus scrofa divergence time to be
161	~0.6 mya [33]. Based on interpopulational distances for the mtDNA CytB sequence
162	(1,140 bp), another estimated divergence time is ~0.5 mya [33, 34]. With phylogenetic
163	analysis based on pairwise genetic distance of the mitotypes of Eurasian Sus scrofa,
164	using an evolutionary rate of 2%, their divergence time was estimated as 0.28 mya;
165	according to mtDNA D-loop sequence divergence time analysis, the Eurasian Sus scrofa
166	divergence time is ~58 kya [35]. Our N_e curves analysis support that Eurasian Sus scrofa
167	could have differentiated in 0.9~0.5 mya (during the long, harsh MPT [9]). The
168	estimations that the divergence time of Eurasian Sus scrofa is ~0.28 mya or ~58 kya
169	could have resulted from the inadequate sequence lengths used for the respective
169 170	could have resulted from the inadequate sequence lengths used for the respective estimations.
170	estimations.
170 171	estimations. We noticed that for all the <i>Bos</i> individuals, their ancestral populations peaked at
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170 171 172 173	estimations. We noticed that for all the <i>Bos</i> individuals, their ancestral populations peaked at the MPT Maxima Class as <i>Sus</i> did, while they experienced their subsequent bottlenecks at different clusters (Figure 3A). The first population bottleneck of African <i>Bos indicus</i> and
170 171 172 173 174	estimations. We noticed that for all the <i>Bos</i> individuals, their ancestral populations peaked at the MPT Maxima Class as <i>Sus</i> did, while they experienced their subsequent bottlenecks at different clusters (Figure 3A). The first population bottleneck of African <i>Bos indicus</i> and one African <i>Bos taurus</i> located in the MBE bottleneck while Eurasian <i>Bos taurus</i> reached
 170 171 172 173 174 175 	estimations. We noticed that for all the <i>Bos</i> individuals, their ancestral populations peaked at the MPT Maxima Class as <i>Sus</i> did, while they experienced their subsequent bottlenecks at different clusters (Figure 3A). The first population bottleneck of African <i>Bos indicus</i> and one African <i>Bos taurus</i> located in the MBE bottleneck while Eurasian <i>Bos taurus</i> reached the first bottleneck at later different cluster (Figure 4B), this phenomenon indicates that

179	The first <i>Bos</i> fossil in Ethiopia originated in ~0.6 mya, indicating a migration of <i>Bos</i> [36].
180	which is also consistent with our estimated Bos taurus differentiation event, and because
181	of their potential ancestral population originated in the Palearctic, we could infer that their
182	migration direction is from Eurasia to Africa. Consequently, the harsh MPT period could
183	also result in the migration and differentiation of Bos taurus.
184	
185	The fluctuation of mammal N_e in the Late Pleistocene (LP) period
186	The N_e curves of many mammals fluctuated during the LP period. Around the time of the
187	LGM and Younger Dryas Event (YDE) (12.6~11.6 kya) [28, 29], the ancestral populations
188	of seven mammal species, including Equus hemionus onager, Macaca nemestrina,
189	Equus quagga boehmi, Canis lupus, Odocoileus virginianus texanus, and Equus caballus,
190	displayed a sudden contraction (Figure S4A). Furthermore, the populations of four
191	species, including Capra aegagrus, Gorilla gorilla diehli and non-African Homo sapiens,
192	experienced a sharp increase (Figure S4B).
193	The aforementioned fluctuation was accompanied by a megafauna extinction
194	during which many large mammals extricated [37-40]. The time of coincidence of the two
195	events indicates that the causes of the megafauna extinction event could also have
196	intensively influenced the ancestors who experienced the aforementioned decline event
197	of the extant mammals. The LGM and the YDE could have contributed to the ancient
198	mammal N_e decline event as well as the megafauna extinction event [37]. Many Aves

199	populations also displayed such fluctuations at the same time, as indicated by their N_e
200	curves [8]. However, neither the LGM nor the YDE was solely sufficient to extinguish
201	megafauna [41], nor could they be the only substantial causes for the N_e decline event.
202	Hence, there could be other factors contributing to these two events.
203	Homo sapiens ancestors could also have contributed to the N_e fluctuations. In our
204	results, the N_e curves of Canis lupus familiaris and Canis lupus started their divergences
205	~0.1 mya, corroborating that the estimated divergence time of the early dog and wolf
206	could be as early as 135 kya [42]. The N_e of the ancestor of Canis lupus experienced a
207	sudden, sharp decline event around the YDE, while the ancestor of Canis lupus familiaris
208	escaped such a decline (Figure 4C). According to recent research, Canis lupus familiaris
209	was domesticated in China in ~33 kya [43], in Central Asia in ~12.5 kya or in Europe in
210	~15 kya [44]. The domestication events occurred around a population contraction event
211	of Canis lupus, the ancestor of which lived without domestication by Homo sapiens.
212	Therefore, humans could be related to the expansion of the Canis lupus familiaris
213	ancestral population. In the last glaciation, Homo sapiens started to spread across the
214	world from Africa with brilliant hunting ability in ~60 kya [45]. Encountering these clever
215	hunters, megafauna seldom understood how to survive. Many mammal N_e decline events
216	also happened after Homo sapiens spread, according to our results (Figure S4A). The
217	Homo sapiens spreading and occupancies agree with the time and trace for the extinction
218	of mammoths [37, 46, 47]. This hints that Homo sapiens could have played an essential

219	role in modern mammal ancestral N_e contractions in the LP. Furthermore, the sharp
220	expansion of Bos indicus-2, Capra aegagrus and Gorilla gorilla diehli after the LGM could
221	also be attributed to mammalian ancestor population LP contractions and megafauna
222	extinctions. Their population contraction event vacated ecological niches; thereafter,
223	hostile environment adaptive ruminants such as Bos and Capra could have been the
224	potential beneficiaries during this harsh period, and the domestication of Homo sapiens
225	could also have facilitated their population expansion after the YDE [48].
226	
227	DISCUSSION
228	Considering earth as a system, paleoclimate fluctuations affect the rise and fall of the
229	population of earth's creatures in the paleoecological system. N_e curves inferred by
230	PSMC analysis offer a chance to investigate N_e dynamics in ancient times, and fossil
231	distributions can verify these results. The influences of the MPT, MBE, LGM and other
232	paleoclimate events on the paleoecosystem are obviously reflected in the mammal N_e
233	curves. Furthermore, the analysis of these mammal N_e curves reveals preliminary
234	impacts of climate fluctuations on the biosphere. In addition, the mammal N_e decline
235	events in harsh stages alert us to pay attention to the dangers that could be caused by
236	dramatic climate change in the future.
237	

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275 AUTHOR CONTRIBUTIONS

- S.C.L. supervised the research and revised the manuscript; Y.L.L. collected, analyzed
- 277 data and prepared the manuscript draft; B.L. participated in the literature review and

- 278 performed some analysis; X.X.P. collected some datasets; Z.J.X provided assistance;
- 279 Q.S supervised the research; S.B.L revised the manuscripts.

281 DECLARATION OF INTERESTS

282 The authors declare no competing interests.

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411 FIGURE LEGENDS

412	Figure 1. PC analysis of mammal demographic trajectories. (A) PC1 and PC2 plots.
413	Different colors or shapes represent different mammal classes. (B) Clusters obtained
414	using PC2 of N_e curves. Colors represent mammal classes. Bos and Sus are marked by
415	triangles and pentagrams, respectively. Mammals in light gray shading share a close
416	genetic relationship and cluster closely when using PC2 of N_e curves. Mammals in dark
417	gray shadows share a relatively close genetic relationship, live in similar habitats, and
418	cluster closely when using PC2 of N_e curves. The scale bar represents the Euclidean
419	distance between samples.
420	Figure 2. Sankey diagram of mammal extrema. Mammals in the same column possess
421	the same extremum. Sample names are in Table S1.
422	Figure 3. Sankey diagram of three classes. (A) MPT Maxima Class. (B) MPT Minima
423	Class. (C) MPT Neutral Class. Sample names are also in Table S1.
424	Figure 4. three mammal Ne curves. (A)Ne curves of Sus scrofa and Sus celebensis;
425	(B) N_e curves of Bos taurus, Bos indicus and Bos grunniens; (C) N_e curves of Canis lupus
426	familiaris and Canis lupus-1.

429 STAR★METHODS

430

431 KEY RESOURCE TABLE

	0011005	
REAGENT or RESOURCES	SOURCE	IDENTIFIER
Bos indicus-1	[49]	SRR3225526
Bos indicus-2	[49]	SRR3546787
Bos indicus-3	[49]	SRR3694652
Bos taurus-1	[49]	SRR3656921
Bos taurus-2	[49]	SRR3694478
Bos taurus-3	[50]	SRR934409
Bos taurus-4	[50]	SRR934402
Bos taurus-5	[51]	SRR3497611
Bos taurus-6	[51]	SRR3499149
Bos grunniens-1	[51]	SRR361224, SRR361227
Capra aegagrus-1	[52]	SRR1576690
Capra hircus-1	[53]	SRR488786, SRR488791, SRR488812, SRR488787,
	[]	SRR488805, SRR488813
Ovis aries-1	[54]	SRR1217914, SRR1217915, SRR1217916,
		SRR1217917, SRR1217939, SRR1217905,
		SRR1217906, SRR1217907, SRR1217908,
		SRR1217909, SRR1217937, SRR1217910, SRR1217911, SRR1217912, SRR1217913,
		SRR1217911, SRR1217912, SRR1217913, SRR1217919, SRR1217918, SRR1217919,
		SRR1217920, SRR1217921, SRR1217940
Ovis ammon polii-1	[55]	SRR5753922
, Equus caballus-1	[56]	SRR1055837
, Equus caballus-2	[56]	SRR1046151
, Equus asinus-1	[56]	SRR873443, SRR873444, SRR873445
, Equus quagga	[57]	ERR660704, ERR660705, ERR660706, ERR660707,
boehmi-1	[-]	ERR660708, ERR660709, ERR660710, ERR660711,
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Equus grevyi-1	[57]	ERR669470, ERR669471, ERR669472, ERR669473,
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Equus zebra	[57]	ERR657893, ERR657894, ERR657895, ERR657896,
nartmannae-1		ERR657897, ERR657898, ERR657899, ERR657900,
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texanus-1 Canis lupus	[59]	SRR4069815 SRR2827578
familiaris-1 Canis lupus familiaris-2	[59]	SRR2827588
Canis lupus familiaris-3	[59]	SRR2827594
Canis lupus familiaris-4	[59]	SRR2827563
Canis lupus familiaris-5	[60]	SRR1138369
Canis lupus-1	[59]	SRR2827600
Macaca mulatta lasiota-1	[61]	SRR278719, SRR278720, SRR278721, SRR278722, SRR278723, SRR278724, SRR278725, SRR278726, SRR278727, SRR278728, SRR278729, SRR278730, SRR278731, SRR278732, SRR278733, SRR278734, SRR278735, SRR278736, SRR278737, SRR278738, SRR278739, SRR278740
Macaca nemestrina-1		SRR1698391, SRR1698394
Macaca thibetana-1	[62]	SRR1024051
Rhinopithecus roxellana-1	[63]	SRR2017664
Rhinopithecus avunculus-1	[63]	SRR1588565
Rhinopithecus strykeri-1	[63]	SRR1588563
Rhinopithecus brelichi-1	[63]	SRR1588562
Sus scrofa-1	[64]	ERR173181
Sus scrofa-2	[64]	ERR173202

Sus scrofa-3	[64]	ERR173217
Sus scrofa-4	[64]	ERR173222
Sus scrofa-5	[64]	ERR173176
Sus scrofa-6	[64]	SRR652446, SRR652448, SRR652449, SRR652450
Sus celebensis-1	[64]	ERR173210
Pongo pygmaeus-1	[65]	SRR748012, SRR748013, SRR748014
Pan troglodytes verus-1	[65]	SRR748081, SRR748082, SRR748083, SRR748084
Pan troglodytes schweinfurthii-1	[65]	SRR726412, SRR726413, SRR726415, SRR726416
Pan paniscus-1	[65]	SRR740831, SRR740832, SRR740835
Gorilla gorilla gorilla-1	[65]	SRR748177, SRR748178
Gorilla gorilla diehli-1	[65]	SRR748109, SRR748110, SRR748111, SRR748112
Gorilla beringei graueri-1	[65]	SRR747961, SRR747962
Homo sapiens-1	[66]	1000genomes
Homo sapiens-2	[66]	1000genomes
Homo sapiens-3	[66]	1000genomes
Homo sapiens-4	[67]	ERZ324532
Homo sapiens-5	[67]	ERZ324535
Ailuropoda	[68]	SRR2716726, SRR2717569, SRR2717570,
melanoleuca-1 Ursus maritimus-1	[7]	SRR2717571 SRR518683
Ursus arctos-1	[7]	SRR510003 SRR518712, SRR518713
Ursus americanus-1	[7]	SRR510712, SRR510713 SRR518723
Ailuropoda	[7] http://fossil	13293, 36729, 36730, 36868, 38649, 40341, 51100,
, maropoda	works.org	92781, 92787, 107996, 108001, 120834
Bos	http://fossil	11803, 21197, 22584, 32093, 32113, 32124, 34105,
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		184245
Canis	http://fossil	11803, 20123, 20126, 20127, 20129, 20130, 20134,
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Software and		
Algorithm Soapnuke (v1.5.6)	[52]	https://libraries.io/github/BGI-flexlab/SOAPnuke
BWA (v0.7.17)	[52]	http://bio-bwa.sourceforge.net
PSMC	[4]	https://github.com/lh3/psmc
QGIS(v3.2.3)		https://qgis.org/en/site/
N _e curve analysis	This paper	https://github.com/Solenyalyl/Ne-curve-analysis

432

433 CONTACT FOR REAGENT AND RESOURCE SHARING

434 Further information and requests for resources and reagents should be directed to and

435 will be fulfilled by the Lead Contact, Shuaicheng Li (shuaicli@gmail.com).

436

437 METHOD DETAILS

438

439 Mammal and fossil distribution

440	Mammalian SRA data came from different studies, and not all of their locations were
441	recorded. Mammalian place of origin was regarded as the location if a sample was
442	collected from the zoo; domestication location was taken to be the location if it was
443	assured; and when it was ambiguous, sampling location was regarded as the location.
444	For each of the remaining samples without location recorded, we researched the species
445	distribution in the world and then chose a representative place as the location. Our
446	mammals were distributed among 15 genera, and their fossil records were found in
447	http://fossilworks.org/ except for Pan and Gorilla. Those fossils originated in the Calabrian
448	(1.8-0.781 mya), middle Pleistocene (0.781-0.126 mya) or late Pleistocene (0.126-0.0117
449	mya). Mammal location and fossil data distribution were plotted using QGIS (Figure S1
450	and Figure S3).

452 Data collection and preprocessing

We downloaded mammalian Illumina sequence data from public SRA datasets
(https://www.ncbi.nlm.nih.gov/sra/), Chinese BAM files were downloaded from
https://www.ebi.ac.uk/, and non-Chinese BAM files were collected from
http://www.internationalgenome.org/. Including 60 specimens from 35 species, these
SRA data were sequenced by Illumina paired-end sequencing technology using different
sequencing platforms, therefore our first step was transforming all the Illumina sequence
data into uniform Q+33 Illumina format. All the raw sequencing reads were filtered by

460	Soapnuke	[52]	(htt	p://soar	.genomics	.org.cn/)) and each	n was alig	gned to	its res	pective

- 461 reference genome or the most related genome from the same genus by BWA 0.7.17 [53].
- 462 Samples with genome coverage depth less than 6X were excluded.

464 *N*_e curve generation

- 465 The historical *N_e* curve of mammals was inferred by PSMC based on genome data [4]. A
- 466 PSMC-based consensus genome sequence was obtained by using 'mpileup' in
- 467 SAMTOOLS. When the RMS (root-mean-squared) quality of reads was below 20, their
- 468 covered sites were marked N for missing data. The sites whose read depths were less
- than a third or more than twice the average depth across the genome were masked to
- 470 avoid collapsed regions in the assembly. When we ran PSMC, the upper limit of TRMCA
- 471 was set to 15, the initial θ/ρ value was set to 5 and time interval was set to
- 472 '1*4+25*2+1*4+1*6' as Li and Durbin [4] did. The mutation rate and generation time we
- 473 applied are in Table S3. The genome coverage depth, coverage ratio, and false-negative
- 474 ratio are in Table S2. For statistical analysis of these *N_e* curves see sections below.
- 475

476 QUANTIFICATION AND STATISTICAL ANALYSIS

477

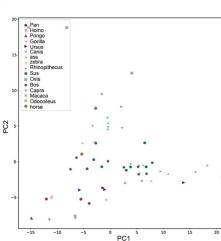
478 **PC** analysis and hierarchical clustering

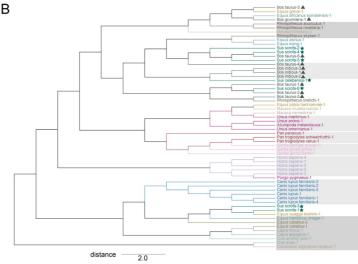
479	PCA was used to identify the principal components of N_e curves. Every N_e curve contains
480	58 points. We standardized these data to unit standard deviation and centralized these
481	points approximately 0, then extracted their principal components to conduct hierarchical
482	clustering. The distance matrix between samples was calculated from Euclidean distance
483	and distance between two clusters was calculated by the UPGMA (unweighted pair group
484	method with arithmetic mean) method.
485	<i>N_e curve extrema analysis</i>
486	To obtain the extrema of every curve, we marked increasing trends as '1' and decreasing
487	trends as '-1' in curves, then the latter trend number was used to subtract the former trend
488	number to obtain the extrema. '-2' and '2' represent crest and trough, respectively. To
489	verify whether these extrema are distributed randomly, we compared these extrema with
490	a uniform distribution to obtain Hopkins statistics. After obtaining a clustering trend, these
491	extrema were clustered by the K-means method to get a pattern. Then we used this
492	pattern to predict every curve extrema into seven clusters by removing those neglectable
493	extrema (difference within 0.1).
494	
495	DATA AND SOFTWARE AVAILABILITY

496 N_e curve analysis code is available from <u>https://github.com/Solenyalyl/Ne-curve-analysis</u>.
497

498 Excel table Legends

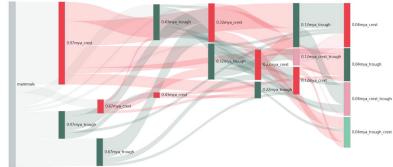
- 499 Table S1. *N*_e curve extrema seven clusters (related to Figure 2,3)
- 500 Table S2. coverage depth, coverage ratio, and false-negative ratio of 60 mammal
- 501 genomes(related to Figure 1-4)



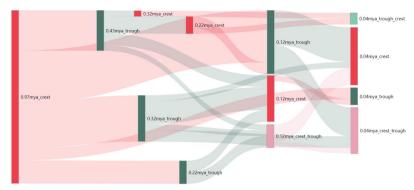


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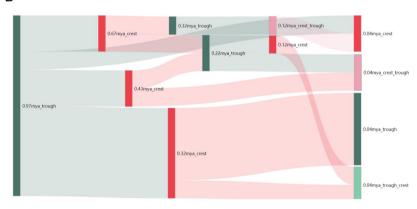
Mammals







В



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